

1 **A new observational evidence of generation and propagation of barotropic Rossby**
2 **waves induced by tropical instability waves in the Northeastern Pacific**

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4 **Kang-Nyeong Lee¹, Chanhyung Jeon², YoungHo Seung¹, Hong-Ryeol Shin³, Seung-Kyu**
5 **Son⁴, and Jae-Hun Park^{1*}**

6 ¹Department of Ocean Sciences, Inha University, Incheon 22212, South Korea.

7 ² Department of Oceanography, Pusan National University, Busan 46241, South Korea.

8 ³Department of Atmospheric Sciences, Kongju National University, Kongju 32588, South
9 Korea.

10 ⁴Deep-sea and Seabed Mineral Resources Research Center, Korea Institute of Ocean Science
11 & Technology, Busan 49111, South Korea.

12
13 Corresponding author: Jae-Hun Park (jaehunpark@inha.ac.kr)

14
15 **Key Points:**

- 16 • In-situ near-bottom current velocity records are coherent with satellite-measured sea
17 surface height related to tropical instability waves.
- 18 • The near-bottom current variations are likely caused by northward propagating
19 tropical instability wave-induced barotropic Rossby waves.
- 20 • Tropical instability wave-induced barotropic Rossby waves vary inter-annually with
21 maxima during the La Niña periods.
22

23 Abstract

24 Tropical instability waves (TIWs) in the equatorial eastern Pacific (EEP) exhibit 25–40-day
25 westward-propagating fluctuations with seasonal and inter-annual variations, which are
26 stronger during July–December and La Niña periods. They likely transfer their energy
27 northward by forming barotropic Rossby waves (BTRWs). Long-term near-bottom current
28 measurements at 10.5°N and 131.3°W during 2004–2013 revealed a spectral peak at 25–40
29 days, where significant coherences were found with satellite-measured sea surface height in a
30 wide region of EEP with maxima approximately 5°N. Simulated deep currents from a data-
31 assimilated ocean model concur with the observed near-bottom currents, and both currents vary
32 seasonally and interannually, consistent with the typical characteristics of TIW. Further
33 analyses using 25–40-day bandpass-filtered barotropic velocity data from the model revealed
34 that they reasonably satisfied the theoretical dispersion relation of TIW-induced BTRW
35 ($BTRW_{TIW}$). We reconfirmed $BTRW_{TIW}$ propagating northward above 10°N in the
36 northeastern Pacific by in-situ observations.

38 Plain Language Summary

39 Tropical instability waves (TIWs), which are located at the boundary between the warm
40 pool and the cold tongue in the eastern Pacific, propagate westward with 25–40-day periods
41 and vary seasonally and interannually, which are stronger during July–December and La Niña
42 periods. Near-bottom velocity measured over a 10-year period at 10.5°N, 131.3°W just above
43 the northern boundary of the waves fluctuates with 25–40-day periods, coinciding with that of
44 sea surface height (SSH) in the equatorial eastern Pacific, especially around 5°N. We find that
45 the wavelike pattern has wave crests oriented southeast-northwest from the model, and that this
46 pattern appears across the study area and has characteristics consistent with TIWs including
47 seasonal and interannual variations with the typical wavenumber and frequency. This pattern
48 was verified to be a barotropic Rossby wave (BTRW) through a model result analysis. Thus,
49 TIWs induce BTRWs that transfer their energy to the abyssal ocean above 10°N in the
50 northeastern Pacific. This study provides a new observational evidence that near-bottom
51 currents vary with BTRWs induced by TIW.

53 1 Introduction

54 Tropical instability waves (TIWs), which propagate westward and have a cusp-like
55 shape with repetitive high amplitudes near 5°N around the boundary of the cold tongue in the
56 equatorial eastern Pacific Ocean, can be observed using satellite-measured sea surface
57 temperature (Legeckis, 1977; Legeckis et al., 1983) and sea surface height (SSH) (Lyman et
58 al., 2005; Farrar, 2011; Holmes and Thomas, 2016; Tchilibou et al., 2018). It is known that
59 TIWs result from instability by interactions between equatorial current system such as the
60 Equatorial Undercurrent, the South Equatorial Current, the North Equatorial Current, and the
61 North Equatorial Countercurrent (Philander, 1976; Lyman et al., 2005). Previous studies
62 described broad ranges of wavenumber and frequency of TIWs depending on measurements
63 utilized for them. Lee et al. (2017) summarized the previous estimates of the wavenumbers
64 and frequencies of TIWs over the spectrum and reported that the TIWs observed by SSH
65 measurements show peak near periods of 33 days and wavelengths of 12°–16° in the
66 wavenumber-frequency spectrum.

67 The waves are representative phenomena with intraseasonal periods in the tropical
68 eastern Pacific Ocean, although these properties are not always remarkable (Chelton et al.,

69 2000; An, 2008; Shinoda et al., 2009). TIWs exhibit seasonal variations in the occurrence of
 70 intense growth from July to December, with more energetic activities during La Niña periods,
 71 linked to the strengthening of upwelling in response to strong trade winds in the equatorial
 72 eastern Pacific (Contreras, 2002; Warner & Moum, 2019).

73 Previous studies have focused mainly on the effects of TIWs near the equatorial ocean
 74 because it is known that the waves play an important role in regional ecosystems and the
 75 balance of heat associated with advection in the equatorial surface ocean (Willett et al., 2006;
 76 Moum et al., 2009). However, Farrar (2011) identified that TIWs can affect their energy up to
 77 approximately 20°N. The longitude-time band-pass filtered SSH shows a structure of TIW at
 78 0°–10°N and a propagation of barotropic Rossby waves (BTRWs) induced by TIW north of
 79 10°N. Furthermore, using both results from barotropic ocean model and newly gridded
 80 satellite-measured SSH with a mapping algorithm without latitudinal variation in its filtering
 81 properties, Farrar et al. (2021) showed that the propagation of the BTRWs continues until
 82 35°N. However, these studies lacked in-situ observations.

83 Here, we used 10-year-long in-situ near-bottom current measurements that were
 84 recorded at a site located north away from the active region of TIW. The in-situ near-bottom
 85 current measurements clearly show that the energy of the TIW-induced BTRWs propagate
 86 northward. The processes of energy propagation in the form of BTRWs were also analyzed
 87 through the satellite-measured SSH as well as the results of data-assimilated numerical
 88 simulation (GLORYS12V1). In addition, the long-term in-situ measurements, satellite
 89 measurements, and results of GLORYS12V1 between 2004 and 2013 enable the verification
 90 of interannual variations according to the El Niño-Southern Oscillation (ENSO).

91 **2 Data and Methods**

92 **2.1 In-situ and satellite measurements and GLORYS12V1 model results**

93 Long-term, half-hour interval near-bottom current data (U_{obs} , V_{obs}) were recorded at a
 94 depth of ~5000 m in the northeastern Pacific (10.5°N, 131.3°W; black star in Figure 1a) from
 95 August 21, 2004 to July 27, 2013. The observations were conducted as part of the Korea
 96 Deep Ocean Study (KODOS). To compare in-situ data with other data explained below, the
 97 former were averaged over a day.

98 Farrar et al. (2021) noted that the SSH data product by Copernicus Climate Change
 99 Service causes barotropic signals with 30-day periods to disappear at higher than 20°N due to
 100 a mapping algorithm. They produced a special-purpose gridded SSH product which has
 101 latitudinally uniform filtering properties. In this paper, we used the newly gridded SSH data
 102 product (hereafter referred to as Farrar SSH) with a space-time grid of $0.5^\circ \times 0.5^\circ \times 3$ days to
 103 conduct the spectral analysis and squared coherency analysis with our in-situ data
 104 subsampled at a 3-day interval. The domain used was 0°–20°N and 140°–80°W during the
 105 same period of near-bottom current measurements.

106 We also used the results of a data-assimilated global ocean reanalysis numerical
 107 simulation (GLORYS12V1) to investigate the characteristics of TIW-induced BTRWs. The
 108 GLORYS12V1 product is provided by the Copernicus Marine Environment Monitoring
 109 Service (CMEMS), and its component is the Nucleus for a European Model of the Ocean
 110 (NEMO) platform. The daily mean GLORYS12V1 outputs have a spatial resolution of $1/12^\circ$
 111 $\times 1/12^\circ$. The selected domain for the analyses is the same as that of Farrar SSH, but the data
 112 cover the period from January 1, 2004 to December 31, 2013. The velocity results at 4833-m
 113 depth filtered by using a band-pass filter with cutoff periods of 25–40 days are consistent
 114 with filtered in-situ near-bottom current measurements, showing high correlation of ~0.8.

115 2.2 Pre-processing of squared coherency

116 The squared coherency (hereafter referred to as coherence) between Farrar SSH and the
 117 time series of in-situ near-bottom current measurements was performed as follows. Spectral
 118 analysis was applied to 1088 -long time series with 3-day interval from August 21, 2004, to
 119 July 27, 2013. A hamming window of length 192 days was used on the segment, and a 50%
 120 overlap was used to increase the number of segments. The 95% significance level,
 121 determined by the number of segments and the window, is 0.137 (Thomson & Emery, 2014).

122 2.3 Complex empirical orthogonal function analysis

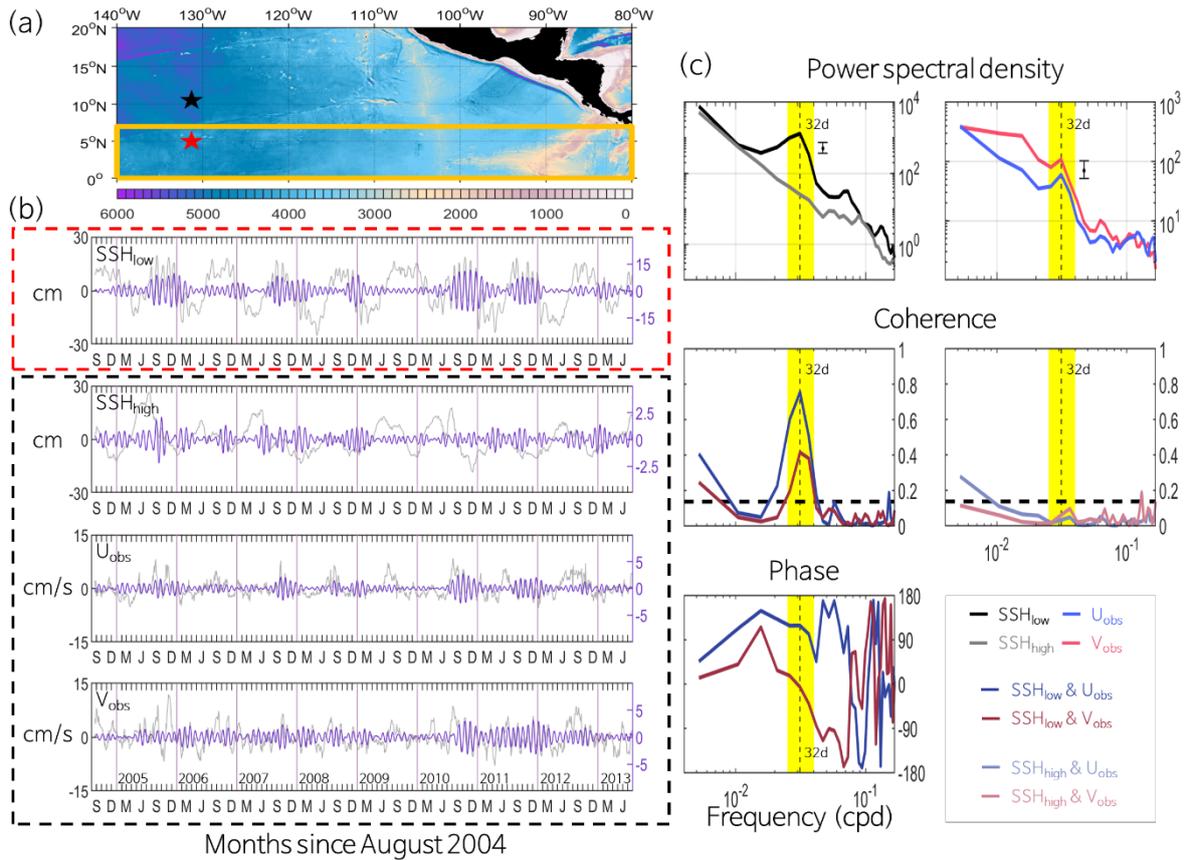
123 The Complex empirical orthogonal function (CEOF) analysis (Hernández-Guerra &
 124 Nykjaer, 1997) using the barotropic velocity results, calculated from the depth average of the
 125 numerical simulation, requires a preprocessing procedure. The results of the numerical
 126 simulation were filtered using a longitude-latitude-time band-pass filter (zonal wavelengths
 127 of 9° – 20° in longitude, meridional wavelengths of 9° – 20° in latitude, and periods of 25–40
 128 days). The longitudinal band-pass filter has a variable cut-off length depending on the
 129 latitudes considered; however, the latitudinal band-pass filter has a constant cut-off length for
 130 all longitudes. These filtering steps were performed sequentially, first for longitude, next for
 131 latitude, and lastly for time. The three dimensions (longitude-latitude-time) filtered data were
 132 converted to two dimensions (spatio-temporal section) and the two components were
 133 concatenated along the row to consider a spatial relationship between them. The results of
 134 CEOF analyses are shown separately for zonal (U_{bt}) and meridional (V_{bt}) components.

135 3 Results

136 To compare the Farrar SSH and in-situ near-bottom current velocity (U_{obs}, V_{obs}) with
 137 each other, two time series of SSH located at different latitudes, indicated by black and red
 138 stars in Figure 1a, were used. One is located at the mooring observation site (SSH_{high}), and
 139 the other is located at $5^{\circ}N, 131.3^{\circ}W$ (SSH_{low}). Figure 1b shows the time series of Farrar
 140 SSH_{low} , SSH_{high} , and in-situ near-bottom current velocity (U_{obs}, V_{obs}) that were filtered by
 141 using a band-pass filter with cutoff periods of 25–40 days. Gray lines superimposed on the
 142 filtered data show the original time series. The time series corresponding to a red star are
 143 surrounded by a box with red dashed lines, and those to a black star are surrounded by a box
 144 with black dashed lines. The maximum speed of the original (filtered) U_{obs} and V_{obs} are 13.5
 145 (2.8) cm/s and 16.7 (3.2) cm/s. The filtered time series of U_{obs} and V_{obs} exhibit similar
 146 variations to SSH_{low} in approximately a month period, which is consistent with the temporal
 147 variation of TIWs reported by Lyman et al., (2007). They are strengthened during the late
 148 summer and early winter months, with inter-annual variations. In contrast, the SSH_{high} shows
 149 no resemblance to others and has substantially smaller values than the original time series.
 150 The results of the spectral analysis also show the same tendency. The spectral peak around
 151 the periods of 32 days clearly shows that the filtered time series has the similar periodicity to
 152 the TIW (top panels in Figure 1c). In contrast, the power spectral density (PSD) of the
 153 SSH_{high} does not show any significant peaks around that period (gray line in Figure 1c).

154 Coherences between the SSH_{low} and the U_{obs} exhibit a maximum value (> 0.75) at the
 155 periods of 32 days and the V_{obs} show higher values (~ 0.4) than the significance level around
 156 the periods of 32 days (middle panel in Figure 1c). In the 32-day periods, the SSH_{low} leads
 157 the U_{obs} by 119° , and the V_{obs} leads the SSH_{low} by -7° . Conversely, coherences between the
 158 SSH_{high} and either the U_{obs} or the V_{obs} appear to be much smaller than the significance level
 159 (0.137) in the 32-day periods. This disparate results seen at two latitudes will be discussed in
 160 Figure 3, by using Farrar SSH data and numerical simulation results.

161

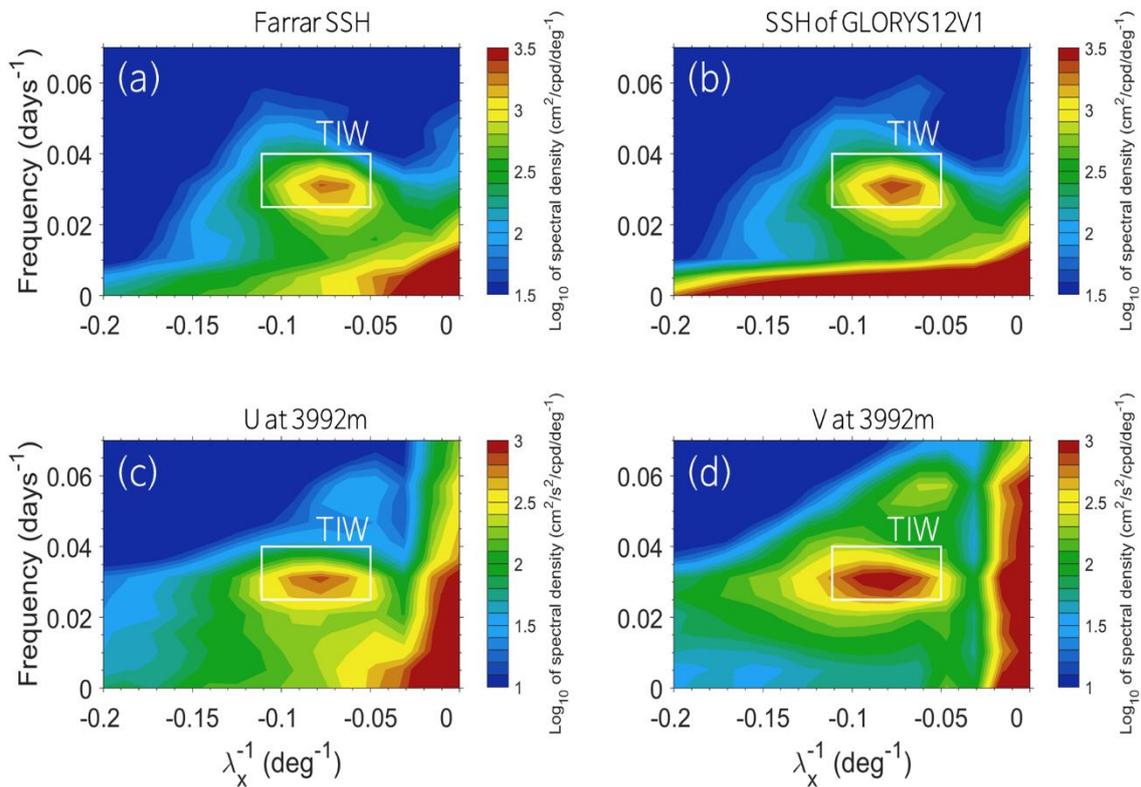


162 **Figure 1.** (a) Bathymetry of the numerical model (GLORYS12V1) domain, with black and
 163 red stars indicating, respectively, the location of the mooring observation site (10.5°N,
 164 131.3°W) and SSH_{low} (5°N, 131.3°W). (b) Bandpass filtered time series (purple lines) of
 165 SSH_{low} , SSH_{high} , U_{obs} , and V_{obs} , having periods of 25–40 days superimposed on their original
 166 time series (light gray lines). Note that SSH data are from Farrar et al. (2021) and the color of
 167 time series matches with that of the axis of ordinates. (c) Power spectral density of SSH_{low}
 168 (black line), SSH_{high} (gray line), U_{obs} (blue line), and V_{obs} (pink line). Vertical bars indicate
 169 the 95% confidence interval. Coherences and phases between SSH_{low} and U_{obs} (SSH_{low} and
 170 V_{obs}), and SSH_{high} and U_{obs} (SSH_{high} and V_{obs}) are represented by a dark blue line (dark pink
 171 line), and a light blue line (light pink line), respectively. The phases between the SSH_{high} and
 172 the in-situ near-bottom current data are not shown here because of the low coherences
 173 between them. Horizontal dashed lines denote the 95% significance level (0.137). Vertical
 174 dashed lines indicate 32-day periods and the yellow shaded areas indicate a TIW frequency
 175 band (periods of 25–40 days).

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178 The zonal wavenumber-frequency power spectral density (two-dimensional PSD)
 179 averaged over 0°–7°N for Farrar SSH resembles the spectrum shown in Farrar (2011), which
 180 is thought to be a common feature associated with the TIW (Figure 2a). The domain used was
 181 0°–7°N, 140°–80°W (yellow box in Figure 1a) and the observation period is from January 1,
 182 2004 to December 31, 2013. Two-dimensional PSDs from longitude-time sections of the data
 183 at different latitudes are averaged, resulting in a function of negative zonal wavenumber and
 184 frequency (Figure 2). The resulting two-dimensional PSD may allow us to identify the spatio-
 185 temporal character of the observed features in the area of active TIWs (the regions

186 surrounded by a yellow box in Figure 1a). Two-dimensional PSDs of the numerical
 187 simulation results show high values in the frequency band of periods 25–40 days and in the
 188 wavenumber band of wavelength 9°–20° with concentrated values near the 32-day periods
 189 (the domains surrounded by the boxes with white lines in Figures 2b–d). It is suggested that
 190 TIWs can affect the currents in the deep layer because the distributions of energy of velocity
 191 components at 3992-m depth from GLORYS12V1 shows similar characters to those of
 192 TIWs.



193 **Figure 2.** Zonal wavenumber-frequency power spectral density (PSD) averaged over 0°–7°N
 194 for (a) Farrar SSH, and numerically simulated (b) SSH and (c, d) velocity components at
 195 3992-m depth from GLORYS12V1. White box is the range of frequency and wavenumber
 196 for the typical TIWs which has periods of 25–40 days and zonal wavelengths of 9°–20°.

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200 To extend the coherence analysis shown in Figure 1c, we calculated the coherences and
 201 phases between gridded Farrar SSH and in-situ near-bottom velocity components (U_{obs} , V_{obs}),
 202 and mapped them by averaging over the frequency range of periods 25–40 days (Figures 3a–
 203 d). The coherence map between the SSH and U_{obs} (V_{obs}) exhibits high values larger than 0.6
 204 (0.4) especially south (southwest) of the mooring observation site. These high coherences
 205 strongly suggest that both the SSH and the observed near-bottom current are related with
 206 TIW (see also Farrar, 2011; Holmes & Thomas, 2016; Farrar et al., 2021). The significant
 207 coherence also exists to the north of the mooring observation site except along the latitude
 208 around 10°N.

209 The positive phase relationship in the region of high coherence (> 0.6) between SSH
 210 and U_{obs} suggests that the SSH which reflects the TIWs in this region leads U_{obs} . In contrast,
 211 the negative phase relationship in the same region between SSH and V_{obs} suggests that V_{obs}

212 leads the SSH. These results about the phase relationship strongly indicate a southwestward
213 phase propagation at periods 25–40 days. There is an abrupt change in phase across the
214 latitude 10°N consistently with the cross-spectral phase estimated in the previous study where
215 satellite-measured gridded SSH data relative to 5°N are used (Farrar, 2011). The speculation
216 by Farrar (2011) is that this abrupt change is probably caused by a coherent superposition of
217 TIWs with barotropic Rossby waves (BTRWs).

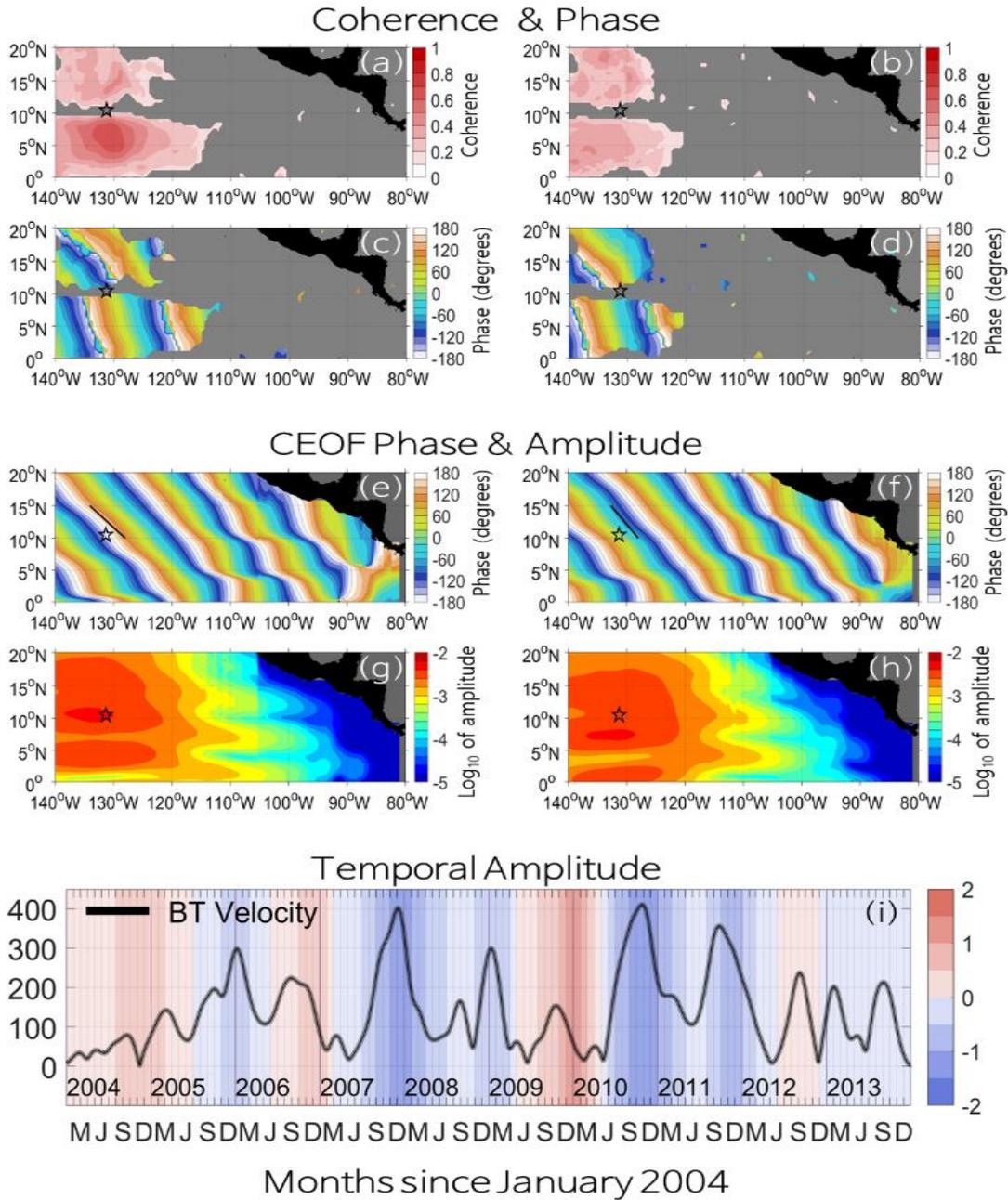
218 To verify the distribution of this BTRWs from numerically simulated barotropic
219 velocity fields, we assume that TIW and BTRW have the same frequency and zonal
220 wavenumber, because the TIWs will induce the BTRWs (Farrar, 2011). We would expect
221 that the filtered barotropic velocity exhibits the BTRW induced by TIW. Further argument
222 regarding whether the filtered barotropic velocity is BTRW induced by TIW is shown in the
223 discussion section using the dispersion relation of BTRW.

224 The CEOF analysis of the filtered barotropic velocity from GLORYS12V1 clearly
225 supports the speculation by Farrar (2011) as follows. The four maps in Figures 3e–h exhibit
226 the first-mode CEOF phase, and the amplitude of the filtered barotropic velocity fields (U_{bt} ,
227 V_{bt}) obtained the GLORYS12V1. The phase progresses southwestward almost uniformly.
228 The amplitude maps also show high values in the west of 115°W with slightly low values
229 along the equator and 7°N in Figure 3g (along 4°N in Figure 3h). These tolerably uniform
230 phase and amplitude indicate that barotropic signal with the wavelengths and periods of TIW
231 exists over the entire study domain. The SSH data include both signal of TIW located on 0° –
232 10°N and this barotropic signal verified from the CEOF analysis. This coexistence induces a
233 node, which explains the non-significant coherence along approximately 10°N in Figure 3a-b.

234 In addition, TIWs are known to strengthen during the La Niña periods when the sea
235 surface temperature is lower and sea surface pressure is higher than those in normal years in
236 the equatorial eastern Pacific (Wang & Fiedler, 2006). The time series of the amplitude of the
237 principal component of the first CEOF mode for the filtered barotropic velocity exhibits an
238 inter-annual variation similar to that of the TIWs (Figure 3i). Large amplitudes are seen
239 around La Niña periods (blue shades in Figure 3i). The largest amplitude occurs during the
240 most intense La Niña periods (2007 and 2010).

241

242



243 **Figure 3.** (a, b) Coherence and (c, d) phase maps between gridded Farrar SSH and observed
 244 near-bottom velocity components (U_{obs} and V_{obs}). (a, b, c, d) The left is for SSH and U_{obs} ,
 245 and the right is for SSH and V_{obs} . The region with coherences under 95% significance level is
 246 gray-colored. (e, f, g, h) The first CEOF mode of the spatio-temporal band-pass filtered
 247 barotropic velocity components (U_{bt} , V_{bt}) obtained from the numerical simulation
 248 (GLORYS12V1) during 10 years from 2004 to 2013. The percentage of the first CEOF mode is
 249 56.5%. The left is for U_{bt} and the right is for V_{bt} . The upper part is the phase and lower part
 250 is the amplitude. (e, f) Superimposed black lines are co-phase line calculated from the CEOF
 251 phase. Black stars denote the location of the mooring observation site. (i) The time series of
 252 the amplitude of the principal component of the first CEOF mode for the barotropic current
 253 velocity obtained from GLORYS12V1. Niño 3.4 indices
 254 (<https://climatedataguide.ucar.edu/climate-data/nino-sst-indices-nino-12-3-34-4-oni-and-tni>)
 255 based on a 30-year data are shown using background colors. Red denotes El Niño and blue

256 denotes La Niña periods, respectively.

257

258 **4 Discussions**

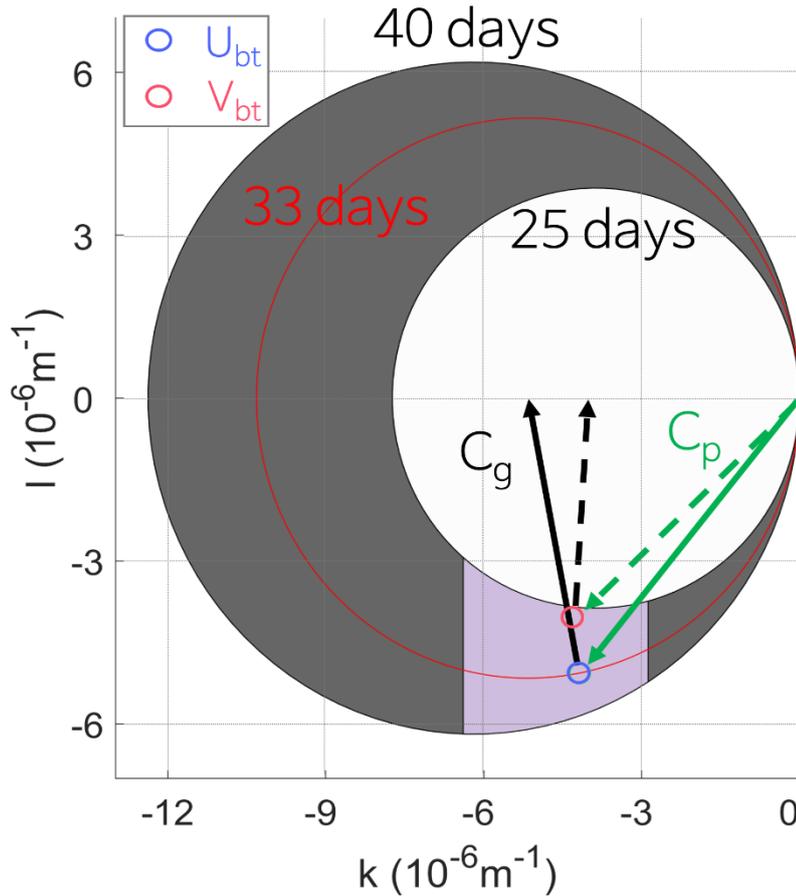
259 To verify our expectation that the CEOF phase of the filtered barotropic velocity is
 260 associated with TIW-induced BTRW, we compare the wavenumbers calculated from the
 261 CEOF phases to the wavenumbers obtained theoretically using the dispersion relation of the
 262 BTRW:

$$\omega = \frac{-\beta k}{k^2 + l^2} \quad (1),$$

263 where ω is frequency, and k and l are the zonal and meridional wavenumbers, respectively. β
 264 is the gradient of the Coriolis parameter at specific latitudes, and we take β at 10.5°N to be
 265 $2.25 \times 10^{-11} \text{ m}^{-1} \text{ s}^{-1}$. Note that, in Eq. (1), the Rossby radius is assumed to be much larger
 266 than the wavelength. The comparison was performed under the assumption that the TIW and
 267 TIW-induced BTRW have the same frequency and zonal wavenumber (Farrar, 2011).

268 The frequencies and zonal wavenumbers of TIW were estimated to be $1.8 \times 10^{-6} \text{ s}^{-1} <$
 269 $\omega < 2.9 \times 10^{-6} \text{ s}^{-1}$ (periods of 25–40 days) and $-6.4 \times 10^{-6} \text{ m}^{-1} < k < -2.9 \times 10^{-6} \text{ m}^{-1}$ (zonal
 270 wavelengths of 9° – 20° of longitude) based on the two-dimensional PSDs of Farrar SSH (see
 271 Figure 2a). The meridional wavenumbers of BTRW were calculated by substituting the
 272 frequencies and zonal wavenumbers of TIW into Eq. (1). The theoretically possible range of
 273 frequencies and wavenumbers of TIW-induced BTRW appears in the wavenumber space to
 274 be the purple region in Figure 4. On the other hand, zonal and meridional wavenumbers are
 275 also calculated using the first-mode CEOF phases as follows. The zonal wavenumbers of U_{bt}
 276 and V_{bt} by using phases at two points (10.5°N , 126.3°W and 10.5°N , 136.3°W) are estimated
 277 to be $-4.18 \times 10^{-6} \text{ m}^{-1}$ and $-4.31 \times 10^{-6} \text{ m}^{-1}$ and the meridional wavenumbers of U_{bt} and V_{bt}
 278 by using phases at the two points (15.5°N , 131.3°W and 5.5°N , 131.3°W) are estimated to be
 279 $-5.06 \times 10^{-6} \text{ m}^{-1}$ and $-4.03 \times 10^{-6} \text{ m}^{-1}$. Note that the co-phase line is shown as the black lines
 280 in the CEOF phase maps (Figures 3e and 3f). The estimated wavenumbers based on U_{bt} and
 281 V_{bt} , are marked by blue and pink small circles in the wavenumber space (Figure 4). It is quite
 282 encouraging that they fall within the possible range of frequency and wavenumber estimated
 283 earlier, supporting the fact that the first CEOF mode is quite compatible with the TIW-
 284 induced BTRW. The direction of group velocity corresponding to the estimated possible
 285 wave frequencies and wavenumbers is northward (Figure 4). This result suggests that our in-
 286 situ near-bottom current measurements enable fluctuation due to the northward propagation
 287 of energy of TIW-induced BTRW estimated by CEOF analysis of GLORYS12V1 results.

288 The non-significant coherence between our in-situ near-bottom current measurements
 289 and Farrar SSH along 10°N indicates that the barotropic signals in SSH near 10°N is not
 290 correlated with the in-situ near-bottom current. This zonally non-correlated band could be
 291 observational evidence to support the speculation of Farrar (2011) that the barotropic signal
 292 in SSH is distorted by the superposition of TIWs and TIW-induced BTRWs.



293 **Figure 4.** Dispersion relation curves of BTRW with periods of 25-40 days. The purple
 294 shading indicates theoretically possible ranges of frequency and wavenumber for TIW-
 295 induced BTRW. A large red circle corresponds to the 33-day period BTRWs. Blue and pink
 296 small circles are wavenumbers estimated from numerical results U_{bt} and V_{bt} . Green and black
 297 solid (dashed) arrows are, respectively, phase and group velocities of TIW-induced BTRW
 298 obtained from U_{bt} (V_{bt}).
 299

300 5 Conclusion

301 Using long-term in-situ near-bottom current measurements, this study provides new
 302 evidence to confirm that the TIW-induced BTRWs propagate their energy northward above
 303 10°N in the northeastern Pacific Ocean. The filtered time series of in-situ near-bottom current
 304 velocity shows that the TIW-induced BTRWs induce a maximum velocity of approximately 3
 305 cm/s at the near bottom and have variations similar to those of TIWs. Our results can be the
 306 answer to the question about whether barotropic waves can actually exist in the deep ocean
 307 where the near-bottom velocities are close to zero due to bottom boundary condition
 308 (LaCasce, 2017).

309 It has been also evidenced from numerical simulation that this energy propagation was
 310 caused by the BTRWs, which showed inter-annual variations because the waves were derived
 311 from TIWs. The inter-annual variations of filtered velocity from numerical simulation and
 312 filtered in-situ observation suggest that the abyssal ocean responds to climate change, ENSO
 313 over the northeastern Pacific Ocean.

314 Our observation suggests that TIW-induced BTRWs transported the energy of the
 315 equatorial eastern Pacific Ocean to the abyssal ocean in high latitudes. The effects of TIWs

316 transported to the abyssal ocean in a low energy environment, due to the lesser vertical
 317 gradient of density and variation of current, can lead to turbulence (Aleynik et al., 2017). The
 318 response of the bottom current is meaningful in that it is possible to affect the advection of
 319 abyssal resources, because the mooring observation site is located in the Clarion-Clippertone
 320 zone. Thus, the long-term in-situ near-bottom current velocity is also expected to improve the
 321 understanding of the distributions of mineral deposits and be used as an evaluation element in
 322 terms of abyssal mining.

323

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 328 Korea.

329

330 **Open Research**

331 The filtered near-bottom current velocity data used in figures can be downloaded
 332 https://github.com/KNLeeinha/KOMO_CM.git and will be deposited Zenodo permanently if
 333 the manuscript is accepted. A newly gridded SSH data product was provided by Farrar et al.
 334 (2021), at <https://doi.org/10.5281/zenodo.4541592>. GLORYS12V1 reanalysis data were
 335 provided by the CMEMS, from their web site at <https://doi.org/10.48670/moi-00021>.

336

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