

1 **How sphericity combines with the age and width of**  
2 **slabs to dictate the dynamics and evolution of**  
3 **subduction systems on Earth**

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8 **Key Points:**

- 9 • The dynamics of subduction systems depend on the thickness, density (combined  
10 approximating age) and width of the downgoing plate.  
11 • Subducting slabs in a spherical geometry exhibit a greater effective strength than  
12 equivalent slabs in a Cartesian geometry.  
13 • The effect of Earth's curvature becomes significant when simulating subduction  
14 systems wider than  $\sim 2000$  km.

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**Abstract**

The role of Earth’s spherical geometry in modulating the evolution of subduction zones is poorly understood. Here, we simulate multi-material free-subduction in a 3-D spherical shell domain, to investigate the effect of plate thickness, density (combined approximating age) and width on the evolution of subduction systems. To isolate the role of sphericity, we compare results with equivalent Cartesian models. The first-order predictions of our spherical cases are generally consistent with existing Cartesian studies: (i) slabs retreat more, at a shallower dip, as plate age increases, due to increased bending resistance and sinking rates; and (ii) wider slabs can develop along-strike variations in trench curvature, trending towards a ‘W’-shape, due to toroidal flow at slab edges. We find, however, that these along-strike variations are restricted to older, stronger, retreating slabs. When compared to slabs in Cartesian models, in a spherical domain: (i) slabs descend faster, due to the convergence of downwelling material with depth; (ii) these faster sinking rates reduce the time available for bending at the trench, resulting in effectively stronger slabs; (iii) the curvature of slabs increases their effective strength; and (iv) the curvature of the transition zone tends to enhance slab stagnation. These differences between spherical and Cartesian cases become more prominent as slab width increases. Taken together, our results suggest that Cartesian models are suitable for simulating narrow subduction zones, but spherical models should be utilised when investigating subduction zones wider than  $\sim 2000$  km: at such length-scales, the consequences of Earth’s curvature cannot be ignored.

**Plain Language Summary**

Subduction zones are locations where Earth’s tectonic plates collide, and the denser plate subsequently descends into the mantle. They exert important controls on surface plate motions and plate boundary deformation, and help to organise underlying mantle flow. As a result, this important process has been extensively studied through both analogue and computational models. However, most of these models have been undertaken in a rectangular box instead of in an Earth-like sphere. Here, we model the dynamics and evolution of subduction systems on a sphere, and find that although our results are similar to the general findings that have been made with box models, there are important differences between these two setups. The curved shape of tectonic plates on a sphere makes them effectively stronger than in the box models and this affects their velocities and shape. They also sink faster in the spherical domain. We find that while box models are suitable for studying narrow plates, we need to model in a spherical setting to study wider subduction zones. Using this spherical setup, we show how the age and width of a subducting plate combine to dictate the evolution of subduction zones on Earth.

**1 Introduction**

During subduction, the oceanic lithosphere of one tectonic plate dives beneath another at a convergent margin and is recycled into Earth’s mantle (e.g., Stern, 2002; Kearey et al., 2009). As subducted slabs descend, their negative buoyancy provides a key driving force for plate tectonics, and they continue to influence surface processes in a number of ways (e.g., Forsyth & Uyeda, 1975; Lithgow-Bertelloni & Richards, 1998; Wheeler & White, 2002). Seismic images of Earth’s interior reveal that when slabs descend towards the mantle transition zone, at depths of 410–660 km that coincide with several mineralogical phase transformations and a likely viscosity increase, some stall and are horizontally deflected (e.g., the Ryuku, Izu-Bonin and Honshu slabs), some thicken and buckle (e.g., the Marianas slab), whilst others appear to pass through unhindered (e.g., the Cocos and Antilles slabs): their imaged morphologies are far from uniform (e.g., Karato et al., 2001; Li et al., 2008; Goes et al., 2017; van der Meer et al., 2018). The dominant con-

65 trols on such variations remain unclear, and likely vary between different subduction zones,  
66 due to complexities arising from non-linear and multi-scale interactions between several  
67 aspects of the mantle system, including downgoing and overriding plate properties, global  
68 mantle flow, mineral phase changes and material rheology (e.g., Karato et al., 2001; Čížková  
69 et al., 2002; Capitanio et al., 2007; Schellart et al., 2007; Goes et al., 2008; Stegman, Far-  
70 rington, et al., 2010; Garel et al., 2014; Goes et al., 2017; Agrusta et al., 2017).

71 The observational record does not support a clear correlation between slab mor-  
72 phology and subducting plate age (e.g., Lallemand et al., 2005; Sdrolias & Müller, 2006;  
73 Goes et al., 2011), implying that variations in a slab's age along the trench, trench width,  
74 local sources of buoyancy such as oceanic plateaus and aseismic ridges, complexities as-  
75 sociated with overriding plates and regional tectonics, can all affect the evolution of sub-  
76 ducting slab morphology. Several studies have investigated the factors controlling the  
77 evolution of subducted slabs, mostly in either 2-D or 3-D Cartesian domains, in an at-  
78 tempt to reconcile predictions from geodynamical modelling with the observed morpholo-  
79 gies (e.g., Čížková et al., 2007; Schellart et al., 2007; Capitanio et al., 2010; Mason et al.,  
80 2010; Stegman, Farrington, et al., 2010; Sharples et al., 2014; Garel et al., 2014; Holt et  
81 al., 2015).

82 The age of a subducting slab determines its thermal structure, which controls slab  
83 thickness, density and rheology. In turn, these control slab strength and buoyancy, which  
84 combine to determine the rate of trench retreat (e.g., Bellahsen et al., 2005; Capitanio  
85 et al., 2007; Di Giuseppe et al., 2008; Schellart, 2008; Ribe, 2010; Stegman, Farrington,  
86 et al., 2010; Garel et al., 2014; Goes et al., 2017). It is well-established that trench-motion  
87 history correlates with slab morphology (e.g., van der Hilst & Seno, 1993; Faccenna et  
88 al., 2001; Goes et al., 2017). Goes et al. (2008) suggest that older, colder, oceanic litho-  
89 sphere is stronger due to the temperature dependence of viscosity, and that this drives  
90 significant trench retreat, with slabs more likely to lie flat at the mantle transition zone;  
91 conversely, younger lithosphere is weaker and subducts with less trench retreat, tend-  
92 ing to buckle at the mantle transition zone. This direct link between slab age and the  
93 style of slab-transition zone interaction is supported by laboratory and numerical sim-  
94 ulations (e.g., Schellart, 2004; Bellahsen et al., 2005; Capitanio et al., 2007; Fucicello  
95 et al., 2008; Garel et al., 2014; Goes et al., 2017).

96 Slab width also plays an important role in determining the evolution of subduc-  
97 tion zones, affecting the shape and curvature of the trench, through an influence on the  
98 rate of trench retreat (e.g., Stegman et al., 2006; Schellart et al., 2007; Stegman, Schel-  
99 lart, & Freeman, 2010; Strak & Schellart, 2016). Schellart et al. (2007) advocate an in-  
100 versely proportional relationship between trench migration rates and the width of sub-  
101 duction zones. In their models, wider slabs develop upper mantle stagnation zones, where  
102 the centre of the trench exhibits negligible trench retreat compared to the edges of the  
103 trench. Such a relationship is observed in large subduction systems such as the South  
104 American subduction zone, and could be responsible for the varying styles of slab mor-  
105 phology along wide trenches.

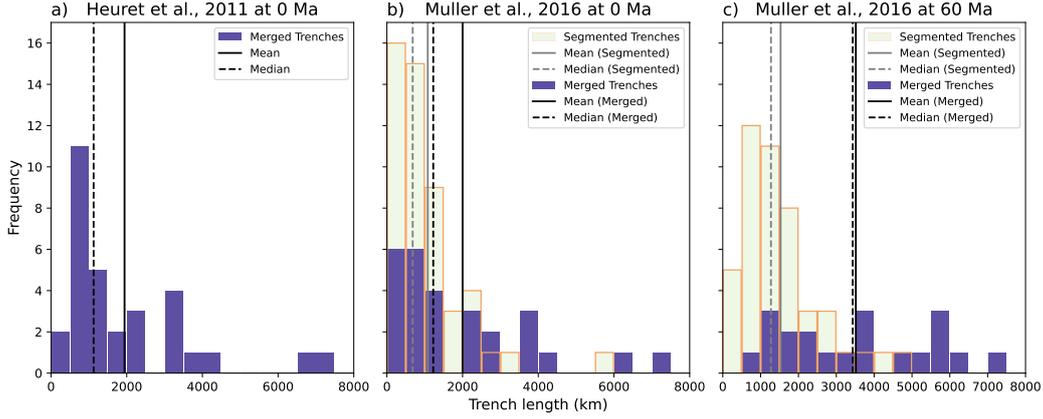
106 Plate interactions and lateral variations in plate properties also influence the ge-  
107 ometry and evolution of subducting plates. The structure and motion of overriding plates  
108 have an effect on slab dip, trench migration, and slab interaction with the mantle tran-  
109 sition zone (e.g., Jarrard, 1986; Lallemand et al., 2005; Heuret et al., 2007; Capitanio  
110 et al., 2010; van Dinther et al., 2010; Garel et al., 2014). The subduction of locally thick-  
111 ened oceanic lithosphere, such as oceanic plateaus, aseismic ridges or seamount chains  
112 has been proposed to influence the shape of the trench and change the geometry of sub-  
113 ducting slabs (e.g., Cross & Pilger, 1982; Gutscher, Malavieille, et al., 1999; Martinod  
114 et al., 2005; Capitanio et al., 2011). The higher compositional buoyancy of oceanic plateaus  
115 and ridges could resist slab sinking into the mantle and, thus, potentially lead to flat slab  
116 subduction (e.g., van Hunen et al., 2002; Mason et al., 2010). These factors add to the

117 complexity of observed slab morphology, and may explain why no simple correlations to  
118 plate properties can be made to explain slab morphology.

119 Existing numerical and laboratory studies in an enclosed Cartesian domain pro-  
120 vide valuable insight into the sensitivity of slab morphology to a number of controlling  
121 parameters. However, Earth’s mantle is a spherical shell, where gravity points radially  
122 towards the center and is bounded only by Earth’s surface and the core-mantle bound-  
123 ary at  $\sim 2890$  km depth: there are no side boundaries. In addition, the mantle closes  
124 in upon itself under geometrical and gravitational constraints, requiring 10% shorten-  
125 ing in any lateral direction when a slab descends from Earth’s surface to the transition  
126 zone. Earth’s surface is also curved in two orthogonal directions and this ‘double cur-  
127 vature’ likely increases the geometric stiffness of slabs (e.g. Mahadevan et al., 2010). As  
128 a consequence, the applicability of Cartesian simulations for investigating the evolution  
129 of subduction systems on Earth remains unclear, particularly for wider subduction zones.

130 Very few studies have investigated the role of Earth’s spherical geometry in con-  
131 trolling the dynamical evolution of subduction systems. To our knowledge, Morra et al.  
132 (2009) were the first to use spherical models at the planetary scale to demonstrate that  
133 during subduction, Earth’s curvature can drive the development of concave curvatures  
134 at plate edges and, for wider plates, complex folding at the centre that becomes more  
135 pronounced at depth. Morra et al. (2012) incorporated a viscosity jump at 660 km depth,  
136 and demonstrate that slab-transition-zone interaction can drive lateral heterogeneity in  
137 trench behaviour. The Boundary Element Method (BEM) used in these studies (e.g.,  
138 Pozrikidis, 1992; Morra et al., 2007) has also been applied to examine the influence of  
139 overriding plates on subduction (Butterworth et al., 2012), and intra-plate deformation  
140 of the Pacific plate in the early Cenozoic (Butterworth et al., 2014). The BEM approach  
141 has many advantages over traditional finite element approaches, including increased nu-  
142 merical efficiency. However, there are also important limitations, including difficulties  
143 in simulating anything but isoviscous plates. This is a major shortcoming as a growing  
144 body of (Cartesian) studies demonstrate that complex plate rheology is fundamental to  
145 reproducing the dynamics of subduction on Earth (e.g., OzBench et al., 2008; Capitanio  
146 et al., 2010; Stegman, Farrington, et al., 2010; Garel et al., 2014; Király et al., 2017).

147 In this paper, we build on the insights gained from these previous Cartesian stud-  
148 ies, to examine the role of Earth’s sphericity in controlling subduction dynamics in sim-  
149 ulations that incorporate a composite visco-plastic plate rheology. Our aim is to inves-  
150 tigate the effect of subducting plate age and width on slab morphology using 3-D spher-  
151 ical numerical models of free subduction, and to isolate the role of sphericity by com-  
152 paring results to Cartesian simulations. We use Fluidity (e.g., Davies et al., 2011; Kramer  
153 et al., 2012; Davies et al., 2016; Kramer et al., 2021), an anisotropic, adaptive, unstruc-  
154 tured mesh computational modelling framework, to examine comparable cases, for a range  
155 of slab densities, thicknesses and widths, in both Cartesian and spherical geometries. We  
156 aim to identify the critical threshold beyond which Cartesian models are no longer ap-  
157 propriate for examining subduction systems on Earth, and the sensitivity of this thresh-  
158 old to plate properties. We examine three combinations of plate thickness and density  
159 to estimate the effect of different plate ages ranging from young ( $\sim 10$  Myr, estimated  
160 using a half space cooling model) to old ( $\sim 140$  Myr), noting that the range of subduct-  
161 ing plate ages on Earth is 0 - 160 Myr (e.g. Müller et al., 2016). Motivated by a com-  
162 pilation of global trench lengths at the present-day (Heuret et al., 2011) and Cenozoic  
163 Era reconstructions (Müller et al., 2016), we examine trench widths of 1200, 2400, 3600  
164 and 4800 km. As illustrated in Figure 1, at the present-day, most trenches are less than  
165 5000 km wide, with mean and median widths of 1940 km and 1130 km for the dataset of  
166 Heuret et al. (2011) and 2000 km and 1230 km for the dataset of Müller et al. (2016). At  
167 60 Ma (where there are less data on narrow trenches), mean and median widths are 3520 km  
168 and 3430 km (Müller et al., 2016). It is noteworthy that very few trenches exceed 6000 km  
169 in width: at present, the South America trench is 7060 km wide, and the Sumatra-Andaman-



**Figure 1.** (a) Present-day trench length histogram compiled from rupture segment lengths by Heuret et al. (2011) and combined here for commonly recognised continuous trenches (see Supplementary Information Figure S1); (b) and (c) trench length histograms based on global tectonic plate reconstructions, at the present-day and at 60 Ma respectively Müller et al. (2016). The reconstructed trenches are segmented based on changes in lower or upper plate characteristics (green bars with orange outline). Based on lower plate properties, we merged segments that likely subduct as coherent slabs (purple bars). Note that corresponding maps of trench segments are provided in Figure S1.

170 Java-Timor trench exceeds 6040 km width; at 60 Myr, the South America and Aleutian  
 171 trenches were  $\sim 7100$  and  $\sim 6070$  km wide, respectively (Müller et al., 2016). Global  
 172 maps of the trenches included in the compilations of Figure 1 are provided in Supple-  
 173 mentary Figure S1.

174 The paper is structured as follows. We first describe the governing equations, ma-  
 175 terial properties, initial and boundary conditions and other aspects of our numerical model  
 176 setup, in addition to listing the different cases examined. We subsequently present a sys-  
 177 tematic quantitative comparison between simulations in Cartesian and spherical domains,  
 178 to demonstrate: (i) how slab thickness and density (approximating slab age) and slab  
 179 width affects the evolution of subducting slabs; and (ii) the significance of Earth’s spher-  
 180 icality in modulating subduction dynamics. We end by discussing our results and their im-  
 181 plications for an improved understanding of subduction on Earth.

## 182 2 Model Description

### 183 2.1 Governing Equations and Numerical Strategy

184 We simulate multi-material free-subduction of a composite visco-plastic plate into  
 185 an ambient mantle, in both 3-D Cartesian and 3-D hemispherical shell domains, which  
 186 extend from the surface to a depth of 2890 km. Assuming incompressibility, the govern-  
 187 ing equations for this problem are the continuity equation,

$$188 \quad \nabla \cdot u = 0 \quad (1)$$

189 the Stokes equation,

$$190 \quad -\vec{\nabla} p + \nabla \cdot \left[ \mu \left( \vec{\nabla} u + \left( \vec{\nabla} u \right)^T \right) \right] = g \Delta \rho \Gamma \hat{k} \quad (2)$$

191 and an advection equation for composition,

$$192 \quad \frac{\partial \Gamma}{\partial t} + \mathbf{u} \cdot \vec{\nabla} \Gamma = 0 \quad (3)$$

193 where  $u$  is velocity,  $p$  the pressure,  $\mu$  the viscosity,  $\rho$  the density,  $g$  gravity acceleration,  
 194  $\hat{k}$  unit vector in the direction opposite gravity, and  $\Gamma$  the material volume fraction ( $\Gamma =$   
 195  $1$  in a region occupied by a given material and  $\Gamma = 0$  elsewhere).

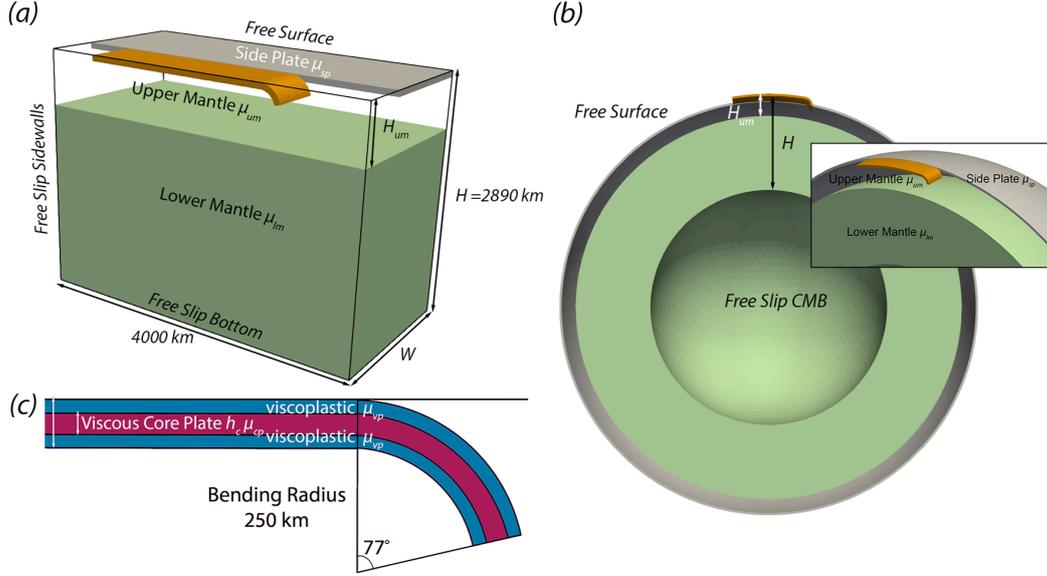
196 Simulations are carried out using Fluidity (e.g., Davies et al., 2011; Kramer et al.,  
 197 2012), a computational modelling framework supporting finite element and control vol-  
 198 ume discretisations, which has recently been validated in a spherical shell domain against  
 199 an extensive set of analytical solutions introduced by Kramer et al. (2021). In the con-  
 200 text of this study, the framework has several ideal features. Fluidity: (i) can run in 2-  
 201 D and 3-D, Cartesian and spherical domains; (ii) uses an unstructured mesh, which en-  
 202 ables the straightforward representation of complex geometries and materials; (iii) dy-  
 203 namically optimizes this mesh, across parallel processors, providing increased resolution  
 204 in areas of dynamic importance, thus allowing for accurate simulations across a range  
 205 of length-scales, within a single model; (iv) enhances mesh optimization using anisotropic  
 206 elements; (v) can employ a free-surface boundary condition, which is important for cor-  
 207 rectly capturing slab decoupling from the surface (Kramer et al., 2012); (vi) utilises the  
 208 highly-scalable parallel linear system solvers available in PETSc (Balay et al., 1997), which  
 209 can efficiently handle sharp, orders of magnitude variations in viscosity; and (vii) has a  
 210 novel interface-preservation scheme, which conserves material volume fractions and al-  
 211 lows for the incorporation of distinct materials (Wilson, 2009). In this study, Fluidity’s  
 212 adaptive mesh capabilities are utilised to provide a local resolution of 3 km in regions of  
 213 dynamic significance (i.e. at the interface between materials and in regions of strong ve-  
 214 locity and viscosity contrasts), with a coarser resolution of up to 300 km elsewhere. It  
 215 is this adaptive mesh functionality that makes our global spherical simulations compu-  
 216 tationally tractable.

## 217 **2.2 Geometry, Boundary Conditions and Material Properties**

218 The configuration of our models are inspired by Stegman, Farrington, et al. (2010)  
 219 and Garel et al. (2014). Both Cartesian and spherical simulations utilise the symmetry  
 220 of the model to halve the computational domain’s extent.

221 For Cartesian simulations (Figure 2a), the domain is 4000 km long, 2890 km deep,  
 222 whilst the width ( $W$ ) depends on the width of the plate ( $w$ ) where  $w/W = 0.3$ . When  
 223 non-dimensionalised with characteristic depth  $H = 2890$  km, the domain depth becomes  
 224 1. The Cartesian model has a free-surface, with free-slip conditions elsewhere, includ-  
 225 ing the symmetric mid-plane. The gravity direction is vertical. For spherical simulations  
 226 (Figure 2b), the domain is a hemispherical shell with outer and inner radii that corre-  
 227 spond to Earth’s surface and core-mantle-boundary (CMB), respectively (Figure 2). When  
 228 non-dimensionalised, the hemispherical shell has an outer radius of 2.22 and an inner ra-  
 229 dius 1.22, thus the computational domain has thickness of 1, and is equivalent to its Carte-  
 230 sian counterpart. The spherical model has a free-surface boundary condition on the outer  
 231 surface, with a free-slip condition on the symmetry plane and CMB. The gravity direc-  
 232 tion points radially towards the centre of the sphere.

The subducting plate length ( $L$ ) is 2200 km. In Cartesian models, the tail of the  
 plate is 600 km from the edge of the domain. The initial slab tip geometry is prescribed  
 with a bending radius of 250 km and an angle of  $77^\circ$  (Figure 2c). The subducting litho-  
 sphere is a composite plate comprising a core isoviscous layer embedded in upper and  
 lower visco-plastic layers with viscosities following a von Mises law, building on OzBench  
 et al. (2008). Upper and lower plates are assigned the minimum viscosity between the  
 Newtonian viscosity  $\mu_{\text{Newt}}$  and an effective von Mises viscosity  $\mu_{\text{vM}}$ , such that purely vis-  
 cous deformation occurs as long as the second invariant of the stress tensor  $\tau_{II} = 2\mu\dot{\epsilon}_{II}$



**Figure 2.** Setup of our simulations in: (a) a Cartesian geometry; and (b) a spherical geometry. In both geometries, we exploit the symmetry of the system, allowing us to halve the computational domain’s extent, whilst bottom and top (inner and outer) boundaries approximate Earth’s core-mantle-boundary and surface, respectively. (c) Initial slab tip geometry of our layered visco-plastic plates.

(where  $\dot{\epsilon}_{II}$  is the second invariant of strain rate tensor) does not reach the critical yield stress,  $\tau_{yield}$ . The effective viscosity of visco-plastic layers is given by:

$$\mu_{vM} = \begin{cases} \frac{\tau_{II}}{2\dot{\epsilon}_{II}}, & \text{if } \tau < \tau_{yield} \\ \frac{\tau_{yield}}{2\dot{\epsilon}_{II}}, & \text{if } \tau \geq \tau_{yield} \end{cases} \quad (4)$$

At material interfaces, the average viscosity is calculated through a geometric mean,

$$\mu_{ave} = \mu_1^{\Gamma_1} * \mu_2^{\Gamma_2}, \quad (5)$$

233 where  $\mu_i$  is the viscosity of material  $i$ , and  $\Gamma_i$  is the relative volume fraction of material  
 234  $i$  in the vicinity of the finite-element node at which the effective viscosity  $\mu_{ave}$  is needed.

235 A side plate covers the entire domain adjacent to the subducting plate. It has the  
 236 same thickness as the plate, and is placed 22 km away from the plate’s edge. It is 1000  
 237 times more viscous than adjacent upper mantle material, and is required to prevent lat-  
 238 eral flow narrowing the width of downgoing plate (as in Holt et al., 2017). The lower man-  
 239 tle is 50 times more viscous than the upper mantle, with the viscosity jump occurring  
 240 at 660 km depth. Model parameters common to all simulations are listed in Table 1.

### 241 2.3 Cases Examined and Quantitative Model Diagnostics

242 We investigate 9 cases, across a wide parameter-space, in both Cartesian and spheri-  
 243 cal coordinate systems, with an additional three spherical cases at a plate width of 3600 km,  
 244 totalling 21 cases. We systematically varied plate thickness ( $h$ ), core plate thickness ( $h_c$ ),  
 245 density contrast between plate and adjacent mantle ( $\Delta\rho$ ) and plate width ( $w$ ) to exam-  
 246 ine how these influence the evolution of subduction and slab morphologies. Our choices  
 247 are motivated by subduction regime diagrams, as a function of plate age and width, from

**Table 1.** Parameters common to all simulations.

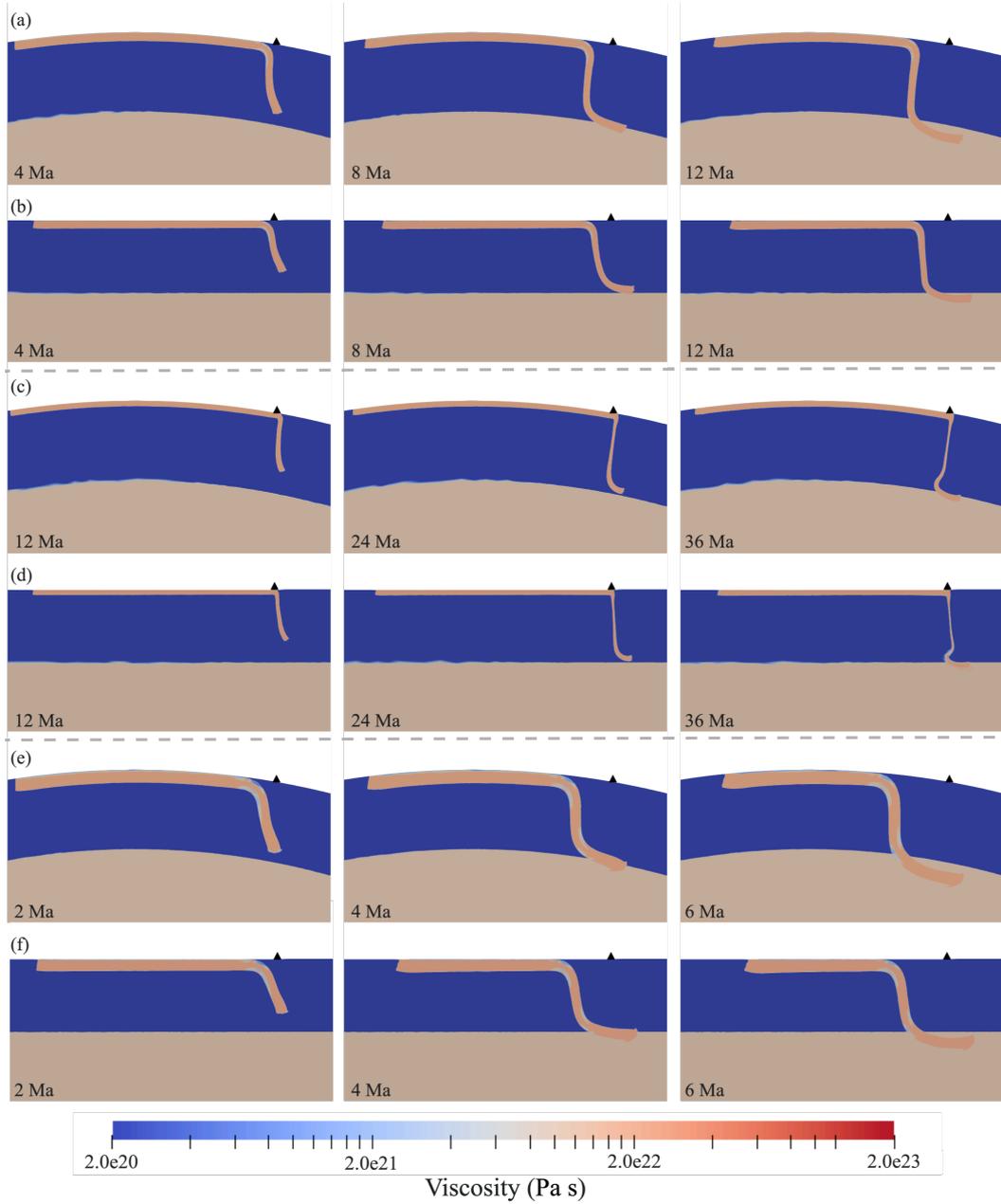
Parameter	Symbol	Value
Gravitational acceleration	$g$	10 m/s <sup>2</sup>
Characteristic depth (whole mantle)	$H$	2890 km
Depth of upper mantle	$H_{um}$	660 km
Upper mantle reference viscosity	$\mu_{um}$	$2.0 \times 10^{20}$ Pa s
Lower mantle reference viscosity	$\mu_{lm}$	$50 \times \mu_{um}$
Core plate viscosity	$\mu_{cp}$	$100 \times \mu_{um}$
Initial viscosity of visco-plastic layer	$\mu_{Newt}$	$100 \times \mu_{um}$
Side plate viscosity	$\mu_{sp}$	$1000 \times \mu_{um}$
Mantle density	$\rho$	3300 kg/m <sup>3</sup>
Yield stress	$\tau_{yield}$	100 MPa

**Table 2.** Simulations examined and associated model parameters.

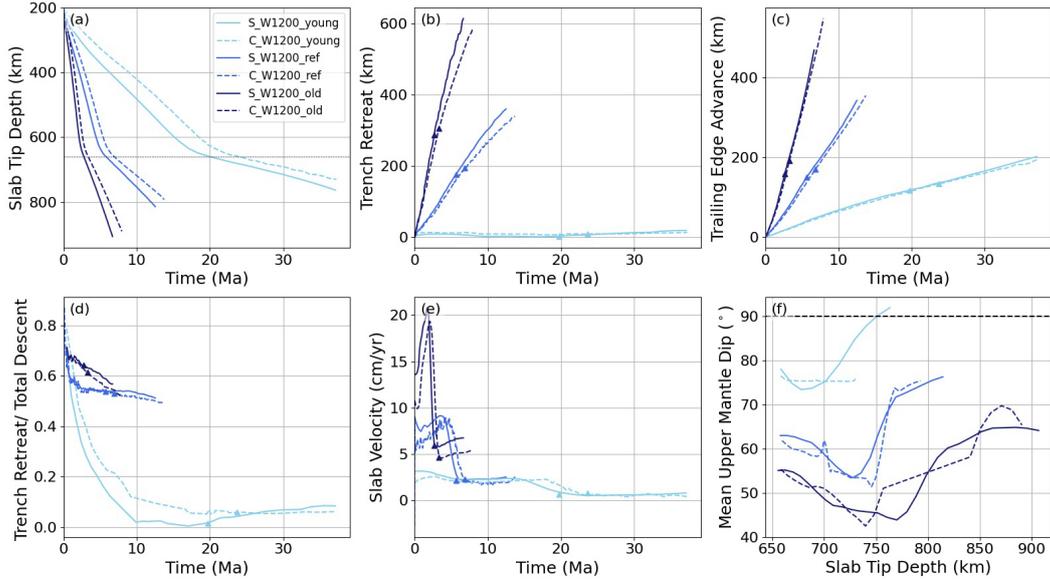
Case	$h$ (km)	$h_c$ (km)	$\Delta\rho$ (kg m <sup>-3</sup> )	$w$ (km)	Domain Type
W1200_young	45	15	40	1200	Cartesian & Spherical
W1200_ref	70	30	80	1200	Cartesian & Spherical
W1200_old	100	40	120	1200	Cartesian & Spherical
W2400_young	45	15	40	2400	Cartesian & Spherical
W2400_ref	70	30	80	2400	Cartesian & Spherical
W2400_old	100	40	120	2400	Cartesian & Spherical
W4800_young	45	15	40	4800	Cartesian & Spherical
W4800_ref	70	30	80	4800	Cartesian & Spherical
W4800_old	100	40	120	4800	Cartesian & Spherical
W3600_young	45	15	40	3600	Spherical
W3600_ref	70	30	80	3600	Spherical
W3600_old	100	40	120	3600	Spherical

248 other studies (e.g., Stegman, Schellart, & Freeman, 2010; Garel et al., 2014; Goes et al.,  
249 2017). The combinations of plate thickness and density contrast produce a range of sub-  
250 duction behaviour from a vertical-folding type young plate to a retreating and flatten-  
251 ing old plate. In the following sections, the plate widths refer to the full widths of the  
252 plate. In practice, we only simulate half of the width exploiting the symmetry of the do-  
253 main. When combined, these cases allow us to compare the effect of plate age, which in-  
254 fluences the thickness and density contrast of a slab, and plate width, on the evolution  
255 of subduction. Case names, alongside their key parameter values, are listed in Table 2.

256 To quantify how these parameters influence results, we have calculated several di-  
257 agnostic outputs from these cases. When doing so, the boundary of the slab is defined  
258 as the 0.5 contour of the mantle material volume fraction (material volume fraction =  
259 1 when the material is mantle, 0 otherwise). Based on this contour, we extract the slab  
260 tip depth, the trench location and the trailing edge position, as well as rates of slab de-  
261 scent, trench retreat and plate advance. We calculate the average slab dip in the upper  
262 mantle from the surface to 650 km depth, with respect to the direction of gravity at the  
263 slab centre at 325 km depth. In Cartesian domains, the direction of gravity is always ver-  
264 tical, whereas for spherical models, the direction of gravity is radially towards the cen-  
265 tre of the sphere from the point of measurement. The measurements are taken at the  
266 symmetry plane unless otherwise specified. We also trace the evolution of trench geom-  
267 etry relative to the initial trench shape.



**Figure 3.** Snapshots illustrating the spatio-temporal evolution of slab morphology through the viscosity field, for spherical and Cartesian models with a plate width of 1200 km: (a) spherical; and (b) Cartesian cases, with  $H = 70$  km and  $\Delta\rho = 80$  kg m $^{-3}$ ; (c) spherical and (d) Cartesian models with  $H = 45$  km and  $\Delta\rho = 40$  kg m $^{-3}$ ; (e) spherical and (f) Cartesian models with  $H = 100$  km and  $\Delta\rho = 120$  kg m $^{-3}$ .



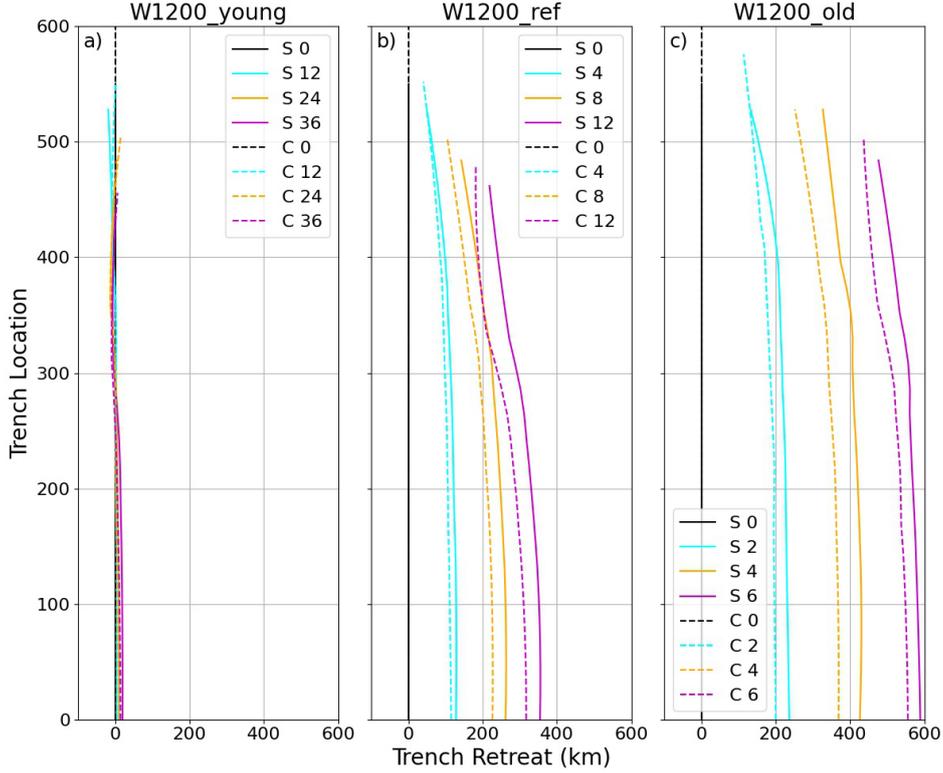
**Figure 4.** Comparison between spherical and Cartesian simulations with a plate width of 1200 km: (a) slab tip depth, as a function of time, where the upper–lower mantle boundary is indicated by the black dotted line at 660 km depth; (b) amount of trench retreat; (c) amount of plate advance, measured at the plate’s trailing edge; (d) ratio of trench retreat to total descent, which is the sum of trench retreat and trailing edge advance; (e) slab sinking velocity; and (f) average slab dip in the upper mantle, with the black dashed line indicating a vertical slab with dip angle of  $90^\circ$ . Triangles indicate the time of slab tip transition-zone interaction. All measurements are taken at the symmetry plane.

## 3 Results

### 3.1 Reference Case

Case W1200\_ref is selected as our reference, given its mid-range plate density and thickness, and width that sits towards the lower end of trench lengths on Earth. The temporal evolution of this case, in both spherical and Cartesian domains, is illustrated in Figure 3(a,b), and both yield similar slab morphologies. As subduction initiates, the slab tip steepens. During the upper mantle sinking phase (Figure 4a), the trench steadily retreats from its initial position (Figure 4b) with  $\sim 50\%$  of subduction accommodated via this trench retreat (over  $60\%$  in the early stages), despite the trailing edge of the plate advancing steadily (Figure 4c,d). As the trench retreats, it develops a concave ‘C’ shape, as illustrated in Figure 5(b). Following interaction with the viscosity jump at 660 km depth, the slab tip is deflected, the slab sinking rate reduces substantially (Figure 4a,e), and the upper mantle section of the slab steepens (Figure 4f). The slab then slowly sinks into the lower mantle.

Coupling of the sinking plate with adjacent mantle drives toroidal and poloidal mantle flow (e.g., Schellart, 2004; Funiello et al., 2006; Stegman et al., 2006). Figure 6(a-c) illustrates tangential flow at 300 km depth for the spherical case at different stages of subduction: the toroidal cell around the edge of the plate drives the increasing concavity of the trench (Figure 5b). Figure 6(d-f) shows vertical cross-sections through the symmetry plane: two poloidal cells can be identified as the slab sinks in the upper mantle, one above the downgoing plate in the mantle wedge, and the other in the sub-slab re-



**Figure 5.** Spatio-temporal evolution of trench location in spherical (S, solid) and Cartesian (C, dashed) simulations, at a plate width of 1200 km. Times given in Myr since simulation initiation. (a)  $H = 45$  km,  $\Delta\rho = 40$  kg m<sup>-3</sup>; (b)  $H = 70$  km,  $\Delta\rho = 80$  kg m<sup>-3</sup>; (c)  $H = 100$  km,  $\Delta\rho = 120$  kg m<sup>-3</sup>.

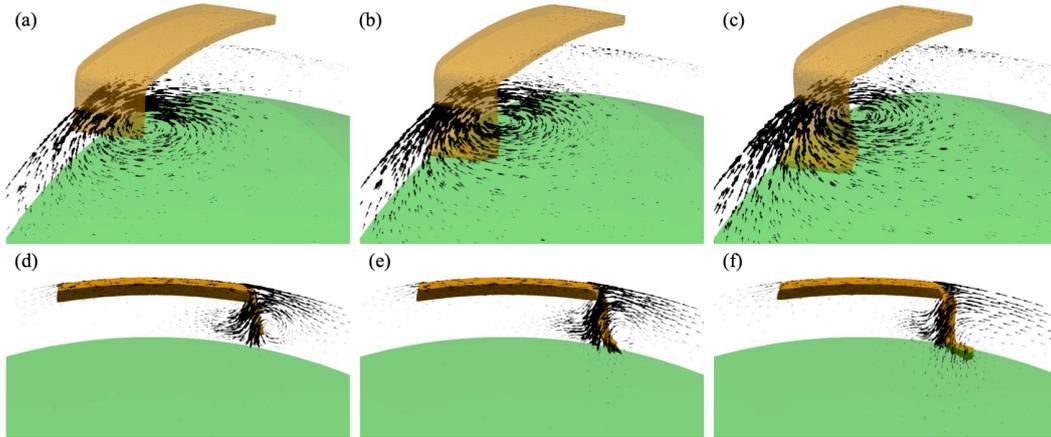
289 gion. During the upper-mantle phase of subduction, the mantle wedge cell is more promi-  
 290 nent, while flow velocities in this cell diminish as the slab tip deflects and sinks into the  
 291 more viscous lower mantle.

292 Cartesian and spherical models generally evolve in a similar manner in space and  
 293 time. However, there are subtle differences that persist across all cases examined: (i) spher-  
 294 ical models exhibit faster sinking rates than their Cartesian counterparts – for the refer-  
 295 ence case, the spherical model displays a maximum sinking velocity of 9 cm/yr, which  
 296 is  $\sim 1.3$  cm/yr faster than the equivalent Cartesian case (Figure 4a,e); and (ii) the rate  
 297 of trench retreat is higher for spherical models – the reference spherical case retreats  $\sim$   
 298 10% faster than its Cartesian counterpart (Figure 4b) and, as a result, the shape of the  
 299 trench evolves differently, with curvature enhanced for spherical cases at a given stage  
 300 of model evolution (Figure 5b).

### 301 **3.2 Influence of subducting plate age**

302 The two cases, W1200\_young and W1200\_old, were designed to demonstrate how  
 303 plate thickness and density modify subduction dynamics. Our parameter values approx-  
 304 imate younger (decreased  $\Delta\rho$  and  $H$ ) and older (increased  $\Delta\rho$  and  $H$ ) slabs, respectively.

305 The younger slab (W1200\_young) stretches and sinks almost vertically through the  
 306 upper mantle as it subducts, folding upon interaction with the transition zone (Figures

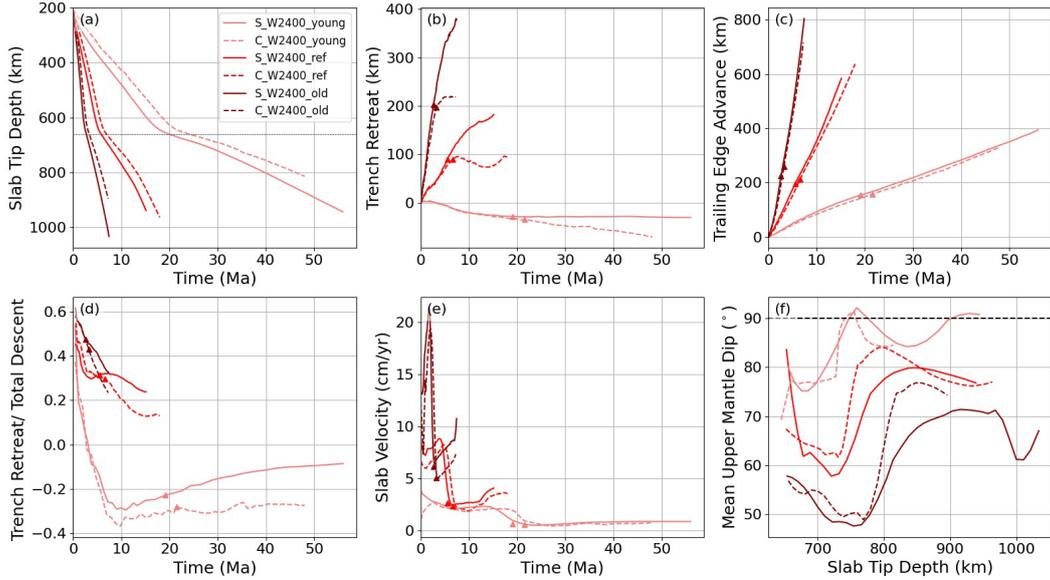


**Figure 6.** Snapshots of upper mantle flow regime from a spherical model at a plate width of 1200 km, plate thickness of 70 km and  $\Delta\rho$  of  $80\text{ kg m}^{-3}$ . (a)-(c) Tangential flow at 300 km depth, highlighting the toroidal flow cell at the edge of the plate. The largest arrows represent a tangential velocity magnitude of 3.7 cm/yr (radial component of velocity removed); (b)-(d) Poloidal flow cells, in the mantle wedge and sub-slab regions, at corresponding times. The largest arrow in the bottom panels represent velocity magnitude of 9.4 cm/yr. As the slab tip interacts with the mantle transition zone, the poloidal cell diminishes as the viscosity increase in the lower mantle prevents return flow beneath the slab tip.

307 3c,d and 4e). Trench location generally remains fixed (Figure 4b) and its shape does not  
 308 evolve much over time (Figure 5a). Excluding the initial phase of subduction, trench re-  
 309 treat is minimal: within 5 Myr of subduction initiation,  $\sim 80\%$  of subduction is accom-  
 310 modated by plate advance (Figure 4d).

311 The older case (W1200\_old) exhibits the fastest sinking, trench retreat and plate  
 312 advance velocities among all cases examined at this width (Figure 4). The slab tip sinks  
 313 in the upper mantle at a shallower angle than the younger cases (Figure 4f). It is de-  
 314 flected at the mantle transition zone, and the sinking rate decreases as the slab moves  
 315 into the lower mantle (Figure 4a,e). Similar to the reference case, after reaching the tran-  
 316 sition zone, the upper mantle portion of the slab gradually steepens (Figure 3e,f). Trench  
 317 retreat is substantial and accommodates the majority of subduction (the trench retreat:total  
 318 descent ratio remains above  $\sim 55\%$  throughout the simulation – Figure 4b,d), with the  
 319 trench developing a concave curvature over time (Figure 5c).

320 The cases examined at 1200 km width clearly display a range of behaviours, with  
 321 a strong sensitivity to the thickness and density and, hence, age, of the subducting slab.  
 322 The younger plate exhibits the weakest behaviour, manifest by a steeper upper mantle  
 323 subduction angle and minimal trench retreat, with subduction principally accommodated  
 324 via plate advance. This case falls into the vertical folding regime (e.g., Schellart, 2008;  
 325 Stegman, Farrington, et al., 2010; Garel et al., 2014; Goes et al., 2017). The older plate  
 326 is the strongest: it sinks faster, has a shallower upper mantle subduction angle, and drives  
 327 significant trench retreat, with the majority of subduction accommodated via this re-  
 328 treat (Figure 4). This case falls into the weak retreat regime (e.g., Schellart, 2008; Stegman,  
 329 Farrington, et al., 2010; Garel et al., 2014; Goes et al., 2017). As expected, the reference  
 330 case has an intermediate strength, with sinking and trench-retreat rates, in addition to  
 331 the slab dip angle, all between those of the older and younger cases (Figure 4). As trench



**Figure 7.** Comparison between spherical and Cartesian simulations with a plate width of 2400 km: (a) slab tip depth, as a function of time, where the upper–lower mantle boundary is indicated by the black dotted line at 660 km depth; (b) amount of trench retreat; (c) amount of plate advance, measured at the plate’s trailing edge; (d) ratio of trench retreat to total descent, which is the sum of trench retreat and trailing edge advance; (e) slab sinking velocity; and (f) average slab dip in the upper mantle, with the black dashed line indicating a vertical slab with dip angle of  $90^\circ$ . Triangles indicate the time of slab tip transition-zone interaction. All measurements are taken at the symmetry plane.

332 retreat accounts for slightly more than  $\sim 50\%$  of the total subduction in the reference  
 333 case, it also falls into the weak retreat regime.

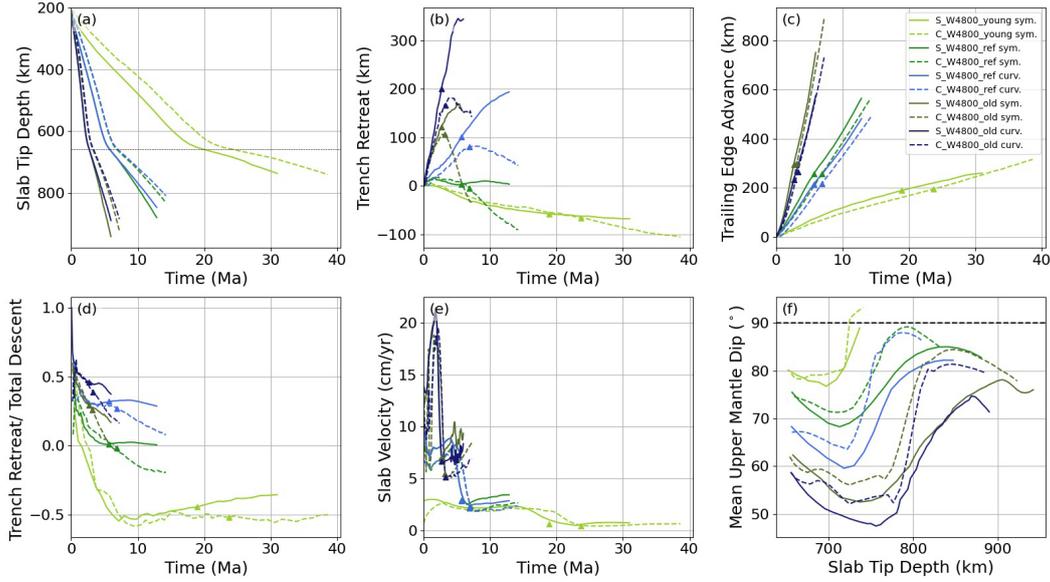
334 Although our spherical and Cartesian models are similar morphologically for the  
 335 1200 km wide plates, spherical models display consistently faster sinking rates than their  
 336 Cartesian counterparts: the older case exhibits the greatest difference in maximum sink-  
 337 ing velocity ( $\sim 1.6$  cm/yr) between comparable spherical and Cartesian cases, followed  
 338 by the reference case ( $\sim 1.3$  cm/yr) and the younger case ( $\sim 0.9$  cm/yr). For the older  
 339 and reference cases, which are both in the weak retreat regime, spherical models exhibit  
 340 faster trench retreat rates than their Cartesian counterparts. The difference in trench  
 341 velocity is negligible between the younger spherical and Cartesian cases, both of which  
 342 fall into the vertical folding regime and display minimal trench motion.

### 333 3.3 Effect of subducting plate width

344 We next examine cases with the same density and thickness values as in the previ-  
 345 ous section, but at larger widths of 2400 km and 4800 km.

#### 346 3.3.1 Less retreat

347 For cases that share a common plate age (thickness and density), a larger plate width  
 348 reduces trench retreat. As the slab tries to maintain its sinking rate, this results in stronger  
 349 bending at the trench. The dynamical behaviour can shift regimes, especially at the cen-

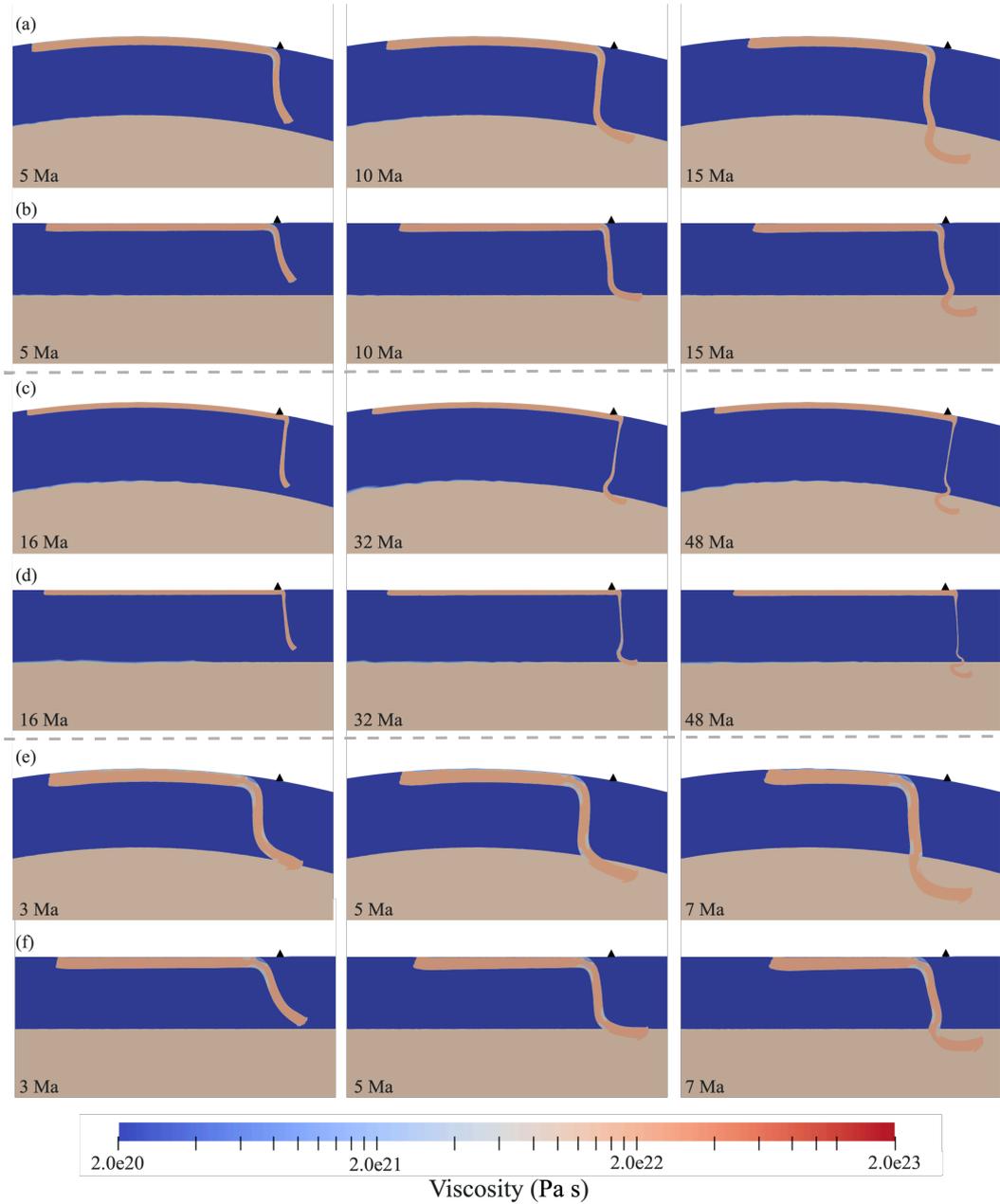


**Figure 8.** Comparison between spherical and Cartesian simulations with a plate width of 4800 km. Measurements are taken from the centre of the slab (i.e., the symmetry plane, abbreviated to sym.) and the location of most trench retreat, which is at the centre of the concave curvature (curv.). (a) slab tip depth, as a function of time, where the upper–lower mantle boundary is indicated by the black dotted line at 660 km depth; (b) amount of trench retreat; (c) amount of plate advance, measured at the plate’s trailing edge; (d) ratio of trench retreat to total descent, which is the sum of trench retreat and trailing edge advance; (e) slab sinking velocity; and (f) average slab dip in the upper mantle, with the black dashed line indicating a vertical slab with dip angle of  $90^\circ$ . Triangles indicate the time of slab tip transition-zone interaction.

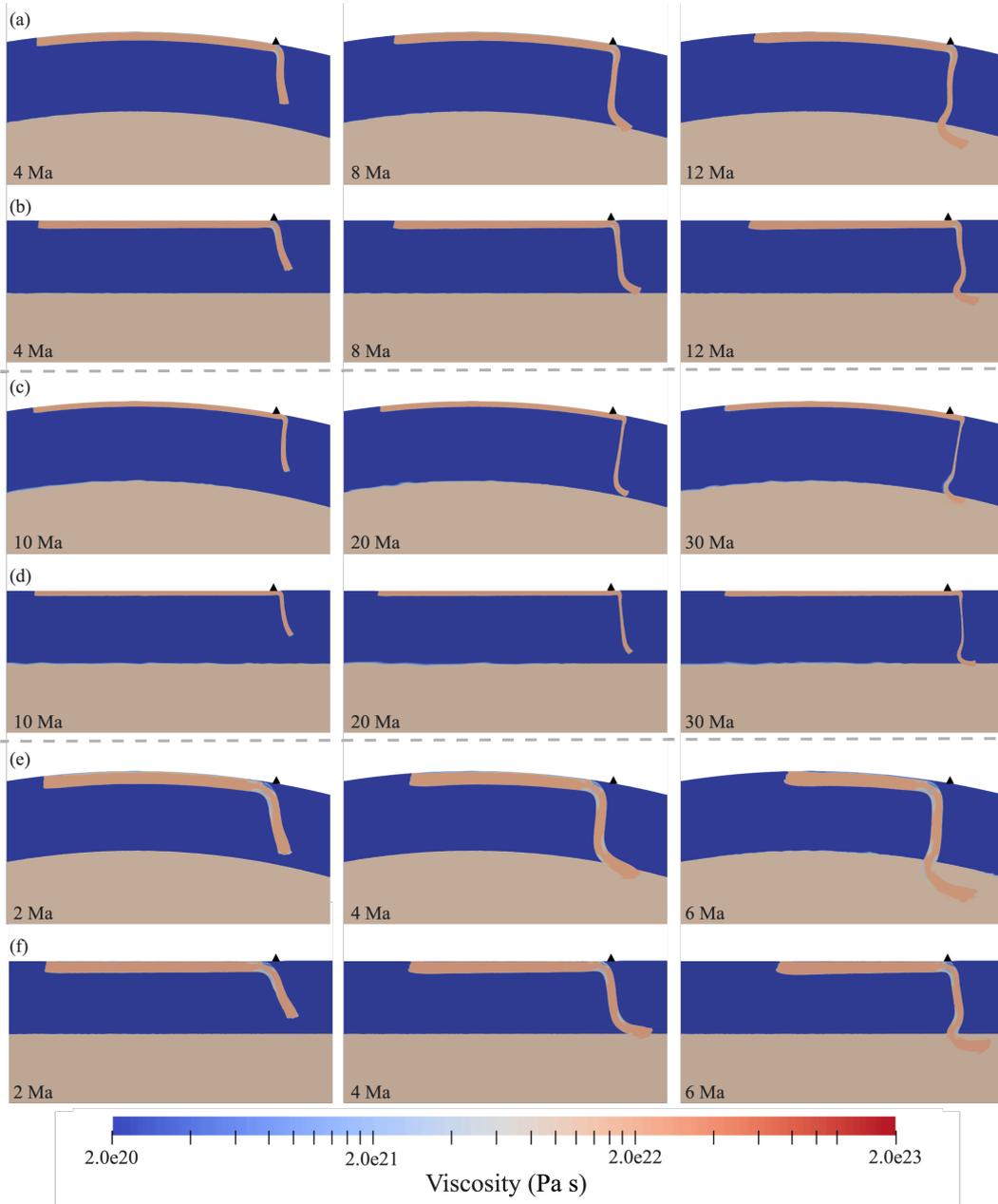
350 tre of the plate, where increased slab width causes slabs to steepen at the trench, with  
 351 the trench sometimes advancing. The behaviour at the centre of the plate thereby shifts  
 352 towards a ‘bending mode’, where slab bending at the trench takes up a significant part  
 353 of the potential energy of the slab, as opposed to a ‘sinking mode’, where bending at the  
 354 trench uses only 10-20% of the potential energy, and slab sinking is, in part, achieved  
 355 through trench retreat (e.g. Capitanio et al., 2007; Ribe, 2010).

356 For younger cases, at both widths (W2400\_young and W4800\_young), slabs stretch  
 357 and sink steeply in the upper mantle, at a dip of  $> 75^\circ$  (Figures 7f and 8f), eventually  
 358 buckling upon interaction with the transition zone at 660 km depth (Figures 9c,d and  
 359 10c,d), like their narrower counterpart. However, as plate width increases, the rate of  
 360 trench advance also increases. Upon interaction with 660 km, the 1200 km cases display  
 361 minimal trench motion (Figure 4b), whereas the trench has advanced  $\sim 30$  and  $\sim 50$  km  
 362 for the 2400 km and 4800 km wide cases, respectively (Figures 7b and 8b). This folding,  
 363 with some advance, is a characteristic of a ‘fold-and-retreat’ bending mode (e.g., Schel-  
 364 lart, 2008; Stegman, Farrington, et al., 2010; Goes et al., 2017), and the centre of wide  
 365 young slabs display behaviour between a vertical folding and fold-and-retreat mode.

366 For wider cases at the reference age (W2400\_ref and W4800\_ref), slabs retreat prior  
 367 to interacting with the transition zone. At the symmetry plane, they steepen and buckle  
 368 following interaction (Figures 9a,b and 10a,b), thus demonstrating stronger bending at  
 369 the trench in comparison to the 1200 km wide case, which displayed a deflect-and-sink



**Figure 9.** Snapshots illustrating the spatio-temporal evolution of slab morphology through the viscosity field, for spherical and Cartesian models with a plate width of 2400 km: (a) spherical; and (b) Cartesian cases, with  $H = 70$  km and  $\Delta\rho = 80$  kg m<sup>-3</sup>; (c) spherical and (d) Cartesian models with  $H = 45$  km and  $\Delta\rho = 40$  kg m<sup>-3</sup>; (e) spherical and (f) Cartesian models with  $H = 100$  km and  $\Delta\rho = 120$  kg m<sup>-3</sup>.



**Figure 10.** Snapshots illustrating the spatio-temporal evolution of slab morphology through the viscosity field, for spherical and Cartesian models with a plate width of 4800 km: (a) spherical; and (b) Cartesian cases, with  $H = 70$  km and  $\Delta\rho = 80$  kg m<sup>-3</sup>; (c) spherical and (d) Cartesian models with  $H = 45$  km and  $\Delta\rho = 40$  kg m<sup>-3</sup>; (e) spherical and (f) Cartesian models with  $H = 100$  km and  $\Delta\rho = 120$  kg m<sup>-3</sup>.

behaviour (Figure 3a,b). As plate width increases, the upper mantle portion of the slab steepens and the dip angle increases (Figure 4f, Figure 7f and Figure 8f). Buckled slabs with a width of 4800 km have maximum dips that exceed those of the 2400 km wide case by  $\sim 4^\circ$ . At the symmetry plane, the trench retreat:total slab descent ratio decreases with plate width ( $\sim 0.5$  and  $\sim 0.2$ , for 1200 km and 2400 km wide cases, respectively, and  $\sim -0.2 - 0$  for the 4800 km wide case), indicating less of a role for trench retreat in accommodating subduction. This is most clearly demonstrated for the Cartesian W4800\_ref simulation, which transitions from retreat at a width of 1200 km, to advance at a width of 4800 km (Figures 4d, 7d and 8d).

For older cases at widths of 2400 km (W2400\_old) and 4800 km (W4800\_old), slabs sink with shallower angles than corresponding reference cases in the upper mantle (Figures 7f and 8f), deflecting at transition zone depths and, subsequently, sinking through into the lower mantle (Figures 9e,f and 10e,f). As plate width increases from 2400 km to 4800 km, the maximum upper mantle dip angle increases by  $\sim 7^\circ$ . The trench retreat:total slab descent ratio also decreases as slab width increases, from  $\sim 0.6$  to  $\sim 0.3$  to  $< 0.2$  for 1200 km, 2400 km and 4800 km wide slabs, respectively. The widest Cartesian case even exhibits trench advance, following slab transition-zone interaction. While the older 1200 km and 2400 km wide cases, in addition to the spherical 4800 km case, all fall into the weak retreat regime, the Cartesian W4800\_old case begins to develop a buckling fold and shifts towards the vertical folding regime.

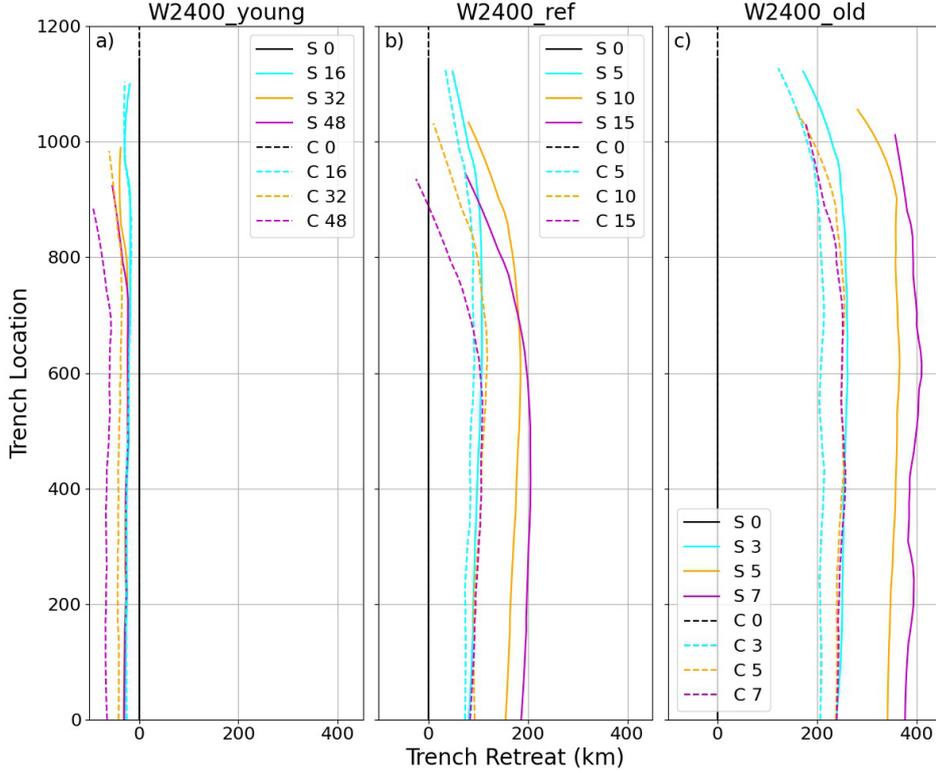
Taken together, our results demonstrate that as plate width increases, slabs display less of a tendency to retreat, as evidenced by a reduction in the trench retreat:total slab descent ratio across all three ages examined and, as a consequence, they bend more strongly at the trench.

### 3.3.2 Trench curvatures and along strike variations in morphology

Different trench shapes are observed across the simulations examined, which can be categorised into 3 types: (i) ‘I’-type, where the trench is reasonably straight (e.g., Figure 5a); (ii) ‘C’-type, where trench retreat is strongest in the centre of the slab relative to its edges (e.g., Figure 5b); and (iii) ‘W’-type, where trench retreat is low in the centre of the slab and at the edges, and higher in between (‘S’ curvature in half-width, as shown in Figure 12b).

We find that ‘I’-type trenches develop for younger cases across all plate widths: trenches remain reasonably straight, aside from a slight curvature adjacent to the slab edge (Figures 5a, 11a and 12a). ‘C’-type trenches develop for narrow plates that are retreating, for example, in cases W1200\_ref and W1200\_old (Figure 5b,c). For stronger plates that have moderate width, such as case W2400\_old (Figure 11b), the trench develops a gentle curvature close to the edge, but the bulk of the trench remains approximately straight throughout the simulation, in an elongated ‘C’ shape. As slab width increases, ‘W’-type trenches develop on slabs that would have ‘C’-type trenches at a narrower width. This is exemplified by comparing cases W2400\_ref and W4800\_ref. Case W2400\_ref develops a concave curvature at the edges, with the centre of the trench retreating slightly less than the edge (Figure 11b), placing it at the transition between ‘C’- and ‘W’-type trenches. Conversely, case W4800\_ref displays a ‘W’-type curvature (Figure 12b). Similarly, for older slabs trenches develop into a ‘W’ shape (‘S’ in half-width in Figure 12 c). In case W4800\_old, the curvature increases following slab transition-zone interaction and the difference in trench retreat between the centre and the region of most retreat also increases (Figure 8b).

Taken together, our results demonstrate that the evolution of trench shape is dependent on both slab age and slab width. Younger and weaker slabs that are in the vertical folding regime develop mostly straight ‘I’-type trenches, regardless of slab width. For older cases that can drive trench retreat, trench curvatures evolve from a ‘C’-shape



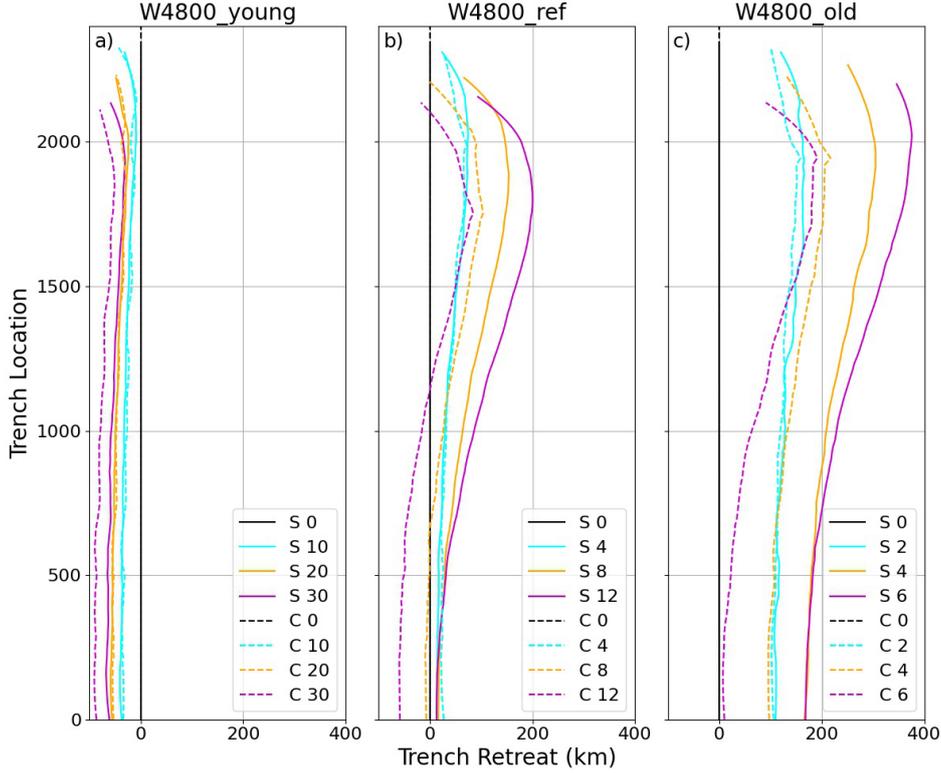
**Figure 11.** Spatio-temporal evolution of trench location in spherical (S, solid) and Cartesian (C, dashed) simulations, at a plate width of 2400 km. Times given in Myr since simulation initiation. (a)  $H = 45$  km,  $\Delta\rho = 40$  kg m $^{-3}$ ; (b)  $H = 70$  km,  $\Delta\rho = 80$  kg m $^{-3}$ ; (c)  $H = 100$  km,  $\Delta\rho = 120$  kg m $^{-3}$ .

421 in narrower plates to a ‘W’-shape in wider plates, with slabs of greater strength transi-  
 422 tioning to a ‘W’ shape at a greater width.

423 Slab morphologies evolve with trench shape. For weaker cases with an ‘I’-type trench,  
 424 subducting slab morphology is relatively uniform along strike (Figure 13a,d). For stronger  
 425 wide retreating cases that develop a ‘W’-type trench, along-strike variations in trench  
 426 retreat translate into morphological variations at depth (Figures 13b,c,e,f and 8a): at  
 427 the symmetry plane, the slab is steep and buckles at the transition zone, with dips up  
 428 to 9° larger than the slab at the location of most retreat, which deflects at transition-  
 429 zone depths (Figure 8d,f).

### 430 3.3.3 Spherical versus Cartesian

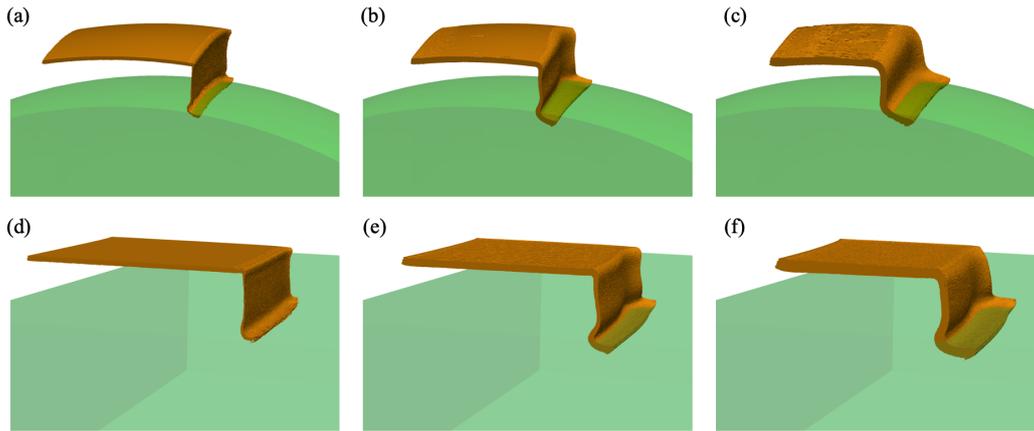
431 We find that, regardless of plate width, all spherical cases evolve faster than their  
 432 Cartesian counterparts, displaying elevated descent rates. Retreating cases also display  
 433 a shallower upper mantle dip angle (a difference of  $\sim 5^\circ$  for reference age slabs, and  $\sim$   
 434  $3^\circ$  for older slabs – Figures 7 and 8). In addition, we find that Cartesian cases are more  
 435 prone to move into a bending mode with increasing plate width. For the younger advanc-  
 436 ing simulations, Cartesian trenches continue to advance at the same rate following slab  
 437 transition-zone interaction, whereas the rate of trench advance is reduced in spherical  
 438 cases. This leads to higher plate advance to total slab descent ratio for Cartesian mod-  
 439 els (Figures 7d and 8d), a characteristic of bending-mode subduction behaviour. For ref-



**Figure 12.** Spatio-temporal evolution of trench location in spherical (S, solid) and Cartesian (C, dashed) simulations, at a plate width of 4800 km. Times given in Myr since simulation initiation. (a)  $H = 45$  km,  $\Delta\rho = 40$  kg m $^{-3}$ ; (b)  $H = 70$  km,  $\Delta\rho = 80$  kg m $^{-3}$ ; (c)  $H = 100$  km,  $\Delta\rho = 120$  kg m $^{-3}$ .

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reference age simulations, the Cartesian case stops retreating after interaction with the tran-  
sition zone at 2400 km width, and even switches from trench retreat to trench advance  
after reaching the lower mantle for the 4800 km width case (Figures 7b and 8b).

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We find that a width of 2400 km is at the tipping point of the Cartesian reference  
model switching from a retreating regime to an advancing regime, after interaction at  
the transition zone. Spherical cases at this width, on the other hand, continue to retreat  
or stagnate after interacting with 660 km. The spherical W3600\_ref model, however, ex-  
hibits trench advance at the symmetry plane after interaction with the lower mantle (Fig-  
ure S2b). Taken together, this suggests that the tipping point from retreating to advanc-  
ing for the spherical reference case is at an increased width of  $\sim 3600$  km. At a width  
of 4800 km, the trench at the symmetry plane of the Cartesian case advances but, in com-  
parison, the spherical cases behave stronger, evolving with ongoing trench retreat (Fig-  
ure 8b). For older cases, only the Cartesian 4800 km case develops buckling at the cen-  
tre of the slab due to its steep angle when hitting the transition zone; the correspond-  
ing spherical case, although steepened, remains sufficiently strong to resist vertical fold-  
ing (Figure 10e,f). The strength of the spherical plate is large enough that the increased  
resistance to slab rollback does not fully hamper trench retreat. For 4800 km wide cases,  
after interaction with the viscosity jump at 660 km, the spherical case continues to re-  
treat without slowing down significantly, but the Cartesian case stops retreating at the  
symmetry plane (Figure 8b). Overall, as slab width increases, the weaker behavior of Carte-



**Figure 13.** 3-D morphology of spherical (top) and Cartesian (bottom) cases at a width of 4800 km: (a/d)  $H = 45$  km,  $\Delta\rho = 40$  kg m<sup>-3</sup>; (b/e)  $H = 70$  km,  $\Delta\rho = 80$  kg m<sup>-3</sup>; (c/f)  $H = 100$  km,  $\Delta\rho = 120$  kg m<sup>-3</sup>. Younger cases (a,d) has relatively uniform morphology along-strike, whereas older cases (b,e,f) have different morphologies: vertically folding at the centre, but horizontally deflect closer to the edge. The spherical cases develop less prominent along-strike variations in morphology than cartesian cases (b,e and c,f)

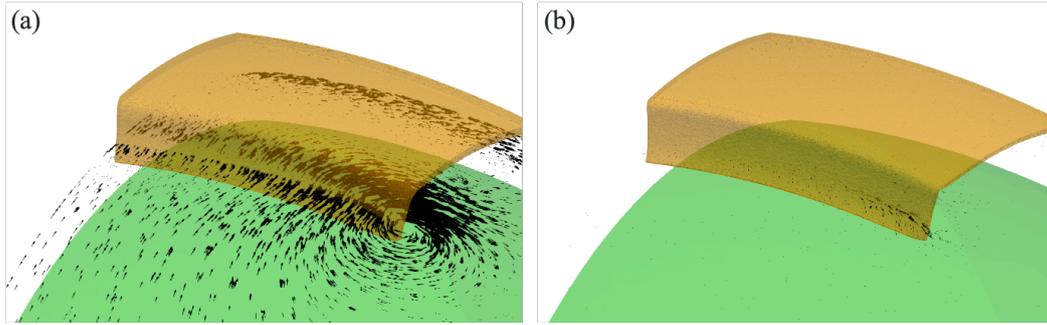
460 sian cases relative to their spherical counterparts becomes more prominent (as outlined  
461 in section 3.3.1).

462 Wider spherical and Cartesian cases also develop significant differences along-strike.  
463 For example, in Case W4800\_ref, the centre of concavity (location of most trench retreat)  
464 continues to retreat after interaction with the transition zone in the spherical model, but  
465 for the Cartesian case, despite initially retreating more than the centre of plate, it switches  
466 to advancing after slab transition-zone interaction (Figure 8b). The differences in retreat  
467 rates and dip angles (Figure 8) lead to different along-strike slab morphologies between  
468 Cartesian and spherical models, as illustrated in Figure 13(b,e). Overall, Cartesian mod-  
469 els display more dramatic along strike variations in morphology than spherical models.

## 470 4 Discussion

### 471 4.1 Role of Subducting Plate Age and Width

472 Our results demonstrate that the evolution of subduction systems is strongly sensi-  
473 tive to slab density and thickness (age), which is consistent with several previous stud-  
474 ies (e.g., Capitanio et al., 2007; Schellart, 2008; Stegman, Farrington, et al., 2010; Garel  
475 et al., 2014). Higher density slabs increase slab pull, which increases upper mantle sink-  
476 ing velocities (e.g., Stegman, Farrington, et al., 2010; Garel et al., 2014; Goes et al., 2017).  
477 Slab thickness determines slab strength (and buoyancy), with thicker slabs possessing  
478 a higher bending resistance (e.g., Conrad & Hager, 1999; Bellahsen et al., 2005; Ribe,  
479 2010; Capitanio & Morra, 2012) and, accordingly, taking longer to bend at the trench.  
480 The regime that a subduction system falls into depends on a delicate balance between  
481 the amount of time taken to bend (larger for thicker slabs) and the sinking time (larger  
482 for younger slabs). Taken together, in younger slabs (low slab pull – longer sinking times;  
483 low slab strength – shorter bending times), bending dominates over trench retreat, with  
484 slabs typically subducting steeply and buckling upon interaction with the mantle tran-  
485 sition zone, owing to the high angle of incidence. Conversely, in older slabs (high slab

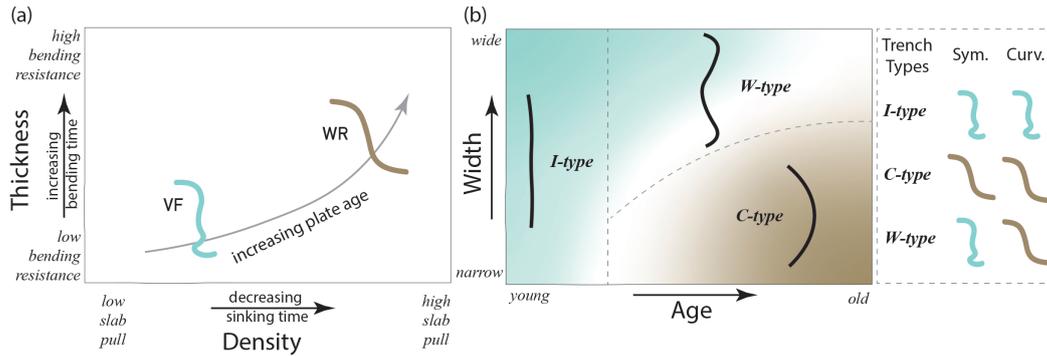


**Figure 14.** Lateral flow patterns at 300 km depth for: (a) case S\_W4800\_ref; and (b) case S\_W4800\_young. The length and direction of the arrows illustrates the magnitude and direction of tangential velocities (i.e. after the radial component has been removed). In both panels, the largest arrow represents a tangential velocity magnitude of 2.5 cm/yr. For the reference age case in (a), a toroidal cell can be identified at the edge of the slab, which has a limited area of influence, and does not affect the centre of the slab. For the younger case in (b), although there is some toroidal flow around the edge of the slab, its magnitude and influence is insignificant when compared to the reference case.

486 pull – shorter sinking times; high slab strength – longer bending times), there is insuf-  
 487 ficient time for substantial bending at the trench, with subduction accommodated prin-  
 488 cipally through trench retreat. As a result, slabs typically exhibit a shallower upper  
 489 mantle dip angle, which prevents slab buckling at the transition zone: the lower the dip  
 490 angle, the more easily slabs can deflect and stagnate at these depths (e.g., Torii & Yosh-  
 491 ioka, 2007; Čížková & Bina, 2013; Garel et al., 2014; Agrusta et al., 2017).

492 The evolution of ‘C’- and ‘W’-trench shapes for our retreating cases are similar to  
 493 results from Schellart et al. (2007), with curvature at slab edges induced by toroidal flow  
 494 into the slab. Interplay between the size and strength of the toroidal cell, the width of  
 495 the slab, and the slab’s tendency to bend, dictate how the trench responds. The size of  
 496 the toroidal cell determines the location along the trench that is experiencing the largest  
 497 force from adjacent mantle flow and, hence, the location of the potential concave cur-  
 498 vature development. The strength of the toroidal cell is determined by slab pull which,  
 499 in turn, determines the strength of forces acting at the trench, whilst the width of the  
 500 plate relative to the size of the toroidal cell determines the distance between the toroidal  
 501 cells at both edges. When these factors are coupled with the the plate’s resistance to bend-  
 502 ing, the evolution of trench shape can be determined. ‘C’-shaped trenches are observed  
 503 for narrow plates, where toroidal cell sizes are almost half of the slab width (Figure 6).  
 504 ‘W’-shaped trenches are observed for wider plates, where the size of the toroidal cell is  
 505 substantially smaller than the width of the plate: the centre of such plates are thus not  
 506 markedly influenced by toroidal flow (Figure 14a). Plates with higher bending resistance,  
 507 which drive trench retreat, can develop enhanced curvature (W4800\_ref, Figure 12b); con-  
 508 versely, plates with a low strength can prevent significant curvature development, remain-  
 509 ing in an ‘I’-shape or elongated ‘C’-shape rather than evolving into a ‘W’-shape (e.g.,  
 510 W2400\_old, Figure 11c).

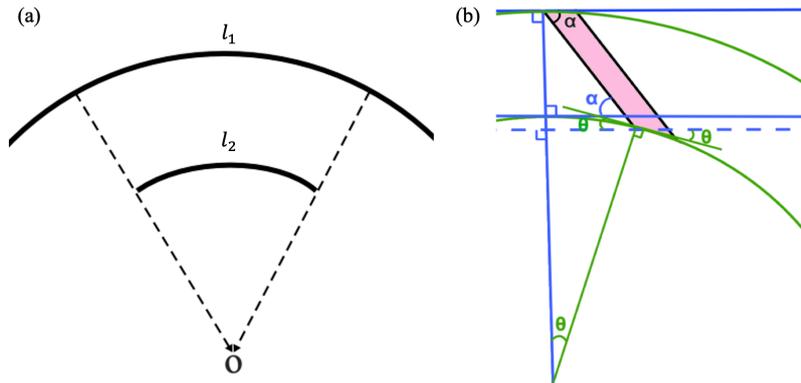
511 While the influence of plate width on subduction dynamics has been carefully stud-  
 512 ied (e.g., Stegman et al., 2006; Schellart et al., 2007; Di Giuseppe et al., 2008), our re-  
 513 sults demonstrate that the important role of width is strongly modulated by the age of  
 514 the plate and its effective strength. The change from a ‘C’-shaped trench to a ‘W’-shaped  
 515 trench with increasing width only occurs for cases that are initially in a retreating regime



**Figure 15.** Schematic diagrams of how (a) density and thickness (adapted from Goes et al. (2017); and (b) age and width, affect subduction styles and trench shape. Regimes: VF - vertical folding; WR - weak retreat. In (a), slabs with higher density have higher slab pull and, accordingly, a reduced upper mantle sinking time. Thicker slabs have more bending resistance and, thus, require more time to bend. Older plates, which are thicker and denser, are able to drive more trench retreat as they have less time to bend. In (b), the three trench types ‘I’, ‘C’ and ‘W’ are separated into three approximate domains, by gray dashed lines, with slabs that lie on domain boundaries at the transition between two trench types. Slab behaviours that are in VF at the symmetry plane are represented by the cyan region, and those in WR are represented by the brown region. The rightmost panel illustrates the slab morphology at the centre of the slab (i.e., the symmetry plane, abbreviated to sym.) and the location of most trench retreat, which is at the centre of the concave curvature (curv.). For young plates, the subduction regime is VF regardless of the width of the plate, and trench shapes are mostly straight, indicated by ‘I’-type. As age increases, it is easier to drive trench retreat and slabs fall into the WR regime; but as width increases, the centre of the plate shifts towards the VF regime. Beyond a certain age, the narrower and/or older plates tend to develop ‘C’-type trenches; wider and/or younger plates tend to develop ‘W’-type trenches.

(i.e., the older plates). For younger plates that are in the vertical folding regime, increasing plate width has little impact on along-strike variability, because the low slab pull and slow upper mantle sinking rates of younger plates are insufficient to generate toroidal cells of the intensity required to induce trench deformation (Figure 14b). Accordingly, the younger cases develop ‘I’-type trench shapes across all widths examined in this study (in both Cartesian and spherical geometries).

Variations in the amount of trench retreat also translate into along-strike morphological variations at depth: the centre of wider slabs are categorised into the vertical folding regime, with steep to overturned upper mantle dips and folding at 660 km depth, whereas in the parts of the slab where the trench retreats most, they subduct in a weak retreat regime with shallower dips and deflect at the transition zone. The lack of toroidal flow influence at the centre of wider slabs reduces the slab’s ability to retreat, which encourages more bending at the trench, leading to steeper slabs that buckle at the transition zone. While all wide slabs display the typical morphology of the vertical folding regime at the centre, the young models have tight buckles, whereas older slabs have open folds with larger bending radii. This difference in bending radii illustrates that older slabs have higher bending resistance and strength than younger slabs, despite falling into the same subduction regime. Overall, as plate width increases, the center of the slab shifts from sinking to bending, due to the lack of toroidal flow and its role in driving trench retreat.



**Figure 16.** Key geometrical features of a spherical geometry that influence subduction evolution when compared to a Cartesian geometry: (a) the spherical geometry concentrates material as it sinks radially towards the centre of the sphere. Bounded by the same radial lines, the length  $l_2$  at depth is shorter than  $l_1$  at the surface. For a 3-D sphere, the tangential area decreases as depth increases (the mantle closes upon itself), concentrating subducting materials; (b) the curvature of the sphere causes the apparent dip of a descending feature to decrease relative to an internal interface. The example is a straight slab of dip  $\alpha$  intersecting the lower mantle in Cartesian and spherical setups (distance is not to scale). The slab forms an angle of  $\alpha$  with the lower mantle in the Cartesian domain (illustrated in blue). In the spherical domain (illustrated in green), the tip of the slab traveled an angular distance of  $\theta$  to reach the lower mantle, and forms an angle of  $(\alpha - \theta)$  with the curved interface at the point of intersection. The angle of difference ( $\theta$ ) due to the curvature is  $\sim 5^\circ$  for plates with an upper mantle dip ( $\alpha$ ) of  $60^\circ$ .

535 The competing role of plate age and plate width in dictating the subduction style  
 536 are summarised via a regime diagram in Figure 15. As plate age (density and thickness)  
 537 increases, the plate behaves more strongly and transitions from a vertical folding regime  
 538 at younger ages to a weak retreat regime regime at older ages. As slab width increases,  
 539 along strike variations in slab morphology can develop due to differences in the amount  
 540 of trench retreat. Younger slabs develop ‘I’-type trenches across all widths. Conversely,  
 541 retreating older cases develop ‘C’-type trenches at narrower widths, and ‘W’-type trenches  
 542 for wider cases, with a transitional ‘C\W’-type at intermediate plate age and width. The  
 543 critical width where trenches transition from ‘C’-type to ‘W’-type depends on plate age:  
 544 the older (stronger) the plate, the greater the slab width required to develop ‘W’ shapes.

## 545 4.2 The Importance of Sphericity

546 One of the most significant differences between spherical and Cartesian geometries  
 547 is the direction of gravity: in spherical domains, gravity acts in the radial direction to-  
 548 wards the centre, whereas in Cartesian simulations, the direction of gravity is constant  
 549 across the entire domain. As illustrated in Figure 16(a), an object of length  $l_1$  sinking  
 550 on a sphere in the direction of gravity must reduce its length according to the reduction  
 551 in radius to maintain the same subtended angle. On Earth, by the time a slab sinks from  
 552 the surface to 660 km depth ( $l_2$ ), its lateral dimensions will be reduced by  $\sim 10\%$ . This  
 553 is a significant amount of shortening that causes buoyancy to concentrate as the man-  
 554 tle closes in upon itself.

555 This concentration of buoyancy increases slab pull and drives faster sinking, thus  
 556 reducing the time available for bending at the trench. As a result, subduction tends to

557 be accommodated through more trench retreat, with the slab descending at a shallower  
558 dip angle. This partially explains why slabs trend towards stronger behavior on a sphere.

559 To fit the curved surface of a sphere, slabs are bent around two orthogonal axes.  
560 This double curvature increases the geometric stiffness of slabs, which are subsequently  
561 able to resist bending and deformation (Mahadevan et al., 2010): the curvature of the  
562 spherical surface therefore increases the stiffness of a subducting slab. As a result, slabs  
563 in a spherical domain require more time to bend at the trench than their Cartesian coun-  
564 terparts. This, combined with the reduced sinking time due to geometrically concentrated  
565 buoyancy, leads to slabs in spherical models having a greater effective strength. The ge-  
566 ometric stiffness and stronger effective strength acts against along-strike deformations,  
567 as exemplified by the less amplified ‘W’-shaped curvature of trenches in wider spheri-  
568 cal cases relative to their Cartesian counterparts.

569 Slab interaction with the transition zone is also influenced by the spherical geom-  
570 etry. The internal interfaces of a sphere, such as the mantle transition zone, are smaller  
571 concentric spheres. As such, the mantle transition zone curves away from the descend-  
572 ing slab at the point of impingement, as illustrated in Figure 16(b). The angle of inter-  
573 action of a slab with the curved mantle transition zone is shallower by the angular dis-  
574 tance,  $\theta$ , travelled by the slab tip, compared to a parallel slab that is in a Cartesian do-  
575 main, where lateral movement of the slab tip does not affect the angle of incidence. This  
576 will enhance trench retreat in a spherical domain, as the lower the dip angle, the more  
577 easily slabs can deflect or stagnate on the transition zone (e.g., Christensen, 2001; Torii  
578 & Yoshioka, 2007; Tagawa et al., 2007).

579 To summarise, both spherical and Cartesian models exhibit a range of slab mor-  
580 phologies and trench curvatures, similar to those predicted in previous studies (e.g., Schel-  
581 lart, 2008; Stegman, Farrington, et al., 2010; Garel et al., 2014). However, plates on a  
582 sphere behave more strongly than plates in a Cartesian domain due to the spherical ge-  
583 ometry. The effect of sphericity becomes more significant for wider plates, being simi-  
584 lar to increasing plate age for Cartesian models. Hence, Cartesian models can capture  
585 the key features of subduction dynamics for narrower plates, but are less suitable for mod-  
586 elling subduction for wider plates. Our results suggest that this limit is approximately  
587 2000 km.

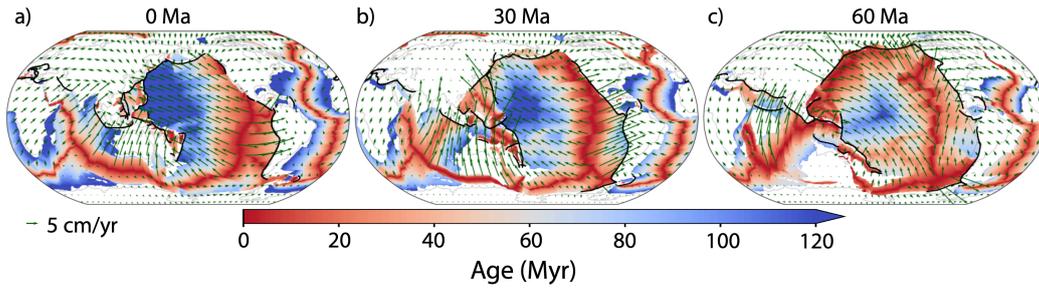
### 588 **4.3 Implications for Subduction on Earth**

589 Our results, across different plate thicknesses, densities and widths, allow us to anal-  
590 yse how plate age and width combine to control the spatio-temporal evolution of trenches  
591 on Earth. While Earth’s subduction zones are substantially more complex than those  
592 considered in our models, due to a multitude of factors including subducting plates of  
593 non-uniform age, the subduction of buoyant anomalies, and the influence of overriding  
594 plates, the ‘I’, ‘C’ and ‘W’ trench shapes predicted by our models are consistent with  
595 present-day trench shapes (e.g. Heuret et al., 2011; Müller et al., 2016) and those in re-  
596 constructions of plate motion histories through the Cenozoic Era (Müller et al., 2019).

#### 597 **4.3.1 ‘I’-type Trenches**

598 Our results demonstrate that ‘I’-shape trenches typically develop when a young plate  
599 subducts with negligible trench motion, regardless of trench width. The tectonic recon-  
600 structions of Müller et al. (2019) provide some examples of young plate subduction dur-  
601 ing the Cenozoic, into the Japan subduction zone at 50–60 Ma and the Farallon (North  
602 American) subduction zone prior to  $\sim 30$  Ma (Figure 17).

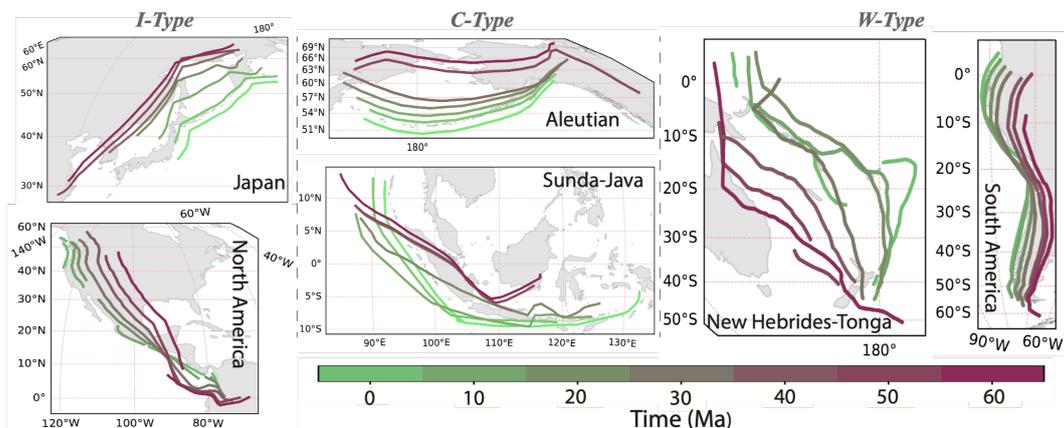
603 As illustrated in Figure 18, the Japan subduction zone was relatively straight (apart  
604 from its northern end) with minimal trench motion between 50 and 60 Ma, character-  
605 istics typical of ‘I’-type trenches. The trench measured  $\sim 5000$  km in width, and a young



**Figure 17.** Maps of ocean-floor age based on plate reconstruction by Müller et al. (2019) at: (a) the present-day; (b) 30 Ma; and (c) 60 Ma. Green arrows represent plate velocity (in the mantle reference frame of the reconstruction). Trenches are shown as black lines and present day coastlines are shown in light grey (Met Office, 2010 - 2015).

606 plate ( $\sim 10$  Myr) was subducted along the whole trench (Figure 17c). As time advanced  
 607 past 30 Ma, the main part of the trench evolved from an ‘I’-type towards a ‘C’-type ex-  
 608 ample, with increasing trench retreat, trench curvature and trench segmentation (Fig-  
 609 ure 18a). This is coincident with an increase in subducting plate age, as shown in Fig-  
 610 ure 17(a,b).

611 Similarly, for Farallon subduction (North America), the reconstructed trench has  
 612 an ‘I’-type shape prior to 30 Ma, when the very young ( $\sim 10$  Myr at 30 Ma) Farallon plate  
 613 was subducting beneath North America, as shown in Figure 17(b). Prior to 30 Ma, the  
 614 trench shape was relatively straight, with very little trench retreat, particularly from 30–50 Ma  
 615 (Figure 18). Following breakup of the Farallon Plate in the mid-Cenozoic, into the Juan  
 616 de Fuca, Cocos and Nazca Plates (e.g., Atwater, 1970; Lonsdale, 2005), the continuity  
 617 of the 5900 km-wide trench was lost, and the strike-slip San Andreas Fault developed on  
 618 the west coast of North America.



**Figure 18.** Examples of ‘I’ shape (a), ‘C’ shape (b) and ‘W’ shape (c) trenches based on the plate reconstruction by Müller et al. (2019), where trenches are drawn at 10 Myr intervals.

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### 4.3.2 ‘C’-type Trenches

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Our results suggest that ‘C’-shape trenches should be associated with moderate to old subducting plate ages and moderate slab widths. ‘C’-shape trenches are the most common trench shape observed on Earth. The Aleutian subduction zone and the Sunda-Java subduction zone are two examples of ‘C’-shape trenches that developed through the Cenozoic (Figure 18). Although both trenches have also been affected by buoyant structures on the incoming plate (such as Yakutat terrane below Alaska and Australian continental crust impinging on the Banda part of the Sunda-Java arc), we propose that the combined width and age of the downgoing plate significantly affected the evolution of trench shape.

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The Aleutian trench extends  $\sim 4000$  km from the south coast of Alaska to Kamchatka (Scholl et al., 1975). At 60 Ma, young material ( $\sim 10$ – $40$  Myr) was subducted along the trench (Figure 17b); as a result, between 50 and 60 Ma, the trench shape remained relatively unchanged, with only a gentle curvature, with a shape between ‘C’- and ‘I’-types (Figure 18, Müller et al., 2019). As time progressed, the age of the subducting plate increased and the Aleutian trench retreated and developed a ‘C’-shape curvature, with enhanced curvature in the west (Figure 18). This is consistent with our modelling predictions, and could be related to the non-uniform subducting plate age at the Aleutian trench: the subducting Pacific plate is younger to the east (currently  $\sim 10$  Myr) and older to the west ( $\sim 120$  Myr at present, Figure 17a,b), with the older part of the plate driving more retreat, generating the asymmetric ‘C’-shaped trench.

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The Sunda-Java trench also has a complex subduction history. Prior to 43 Ma, the active Wharton ridge was subducting beneath Sumatra, but since then the ridge has become inactive (Whittaker et al., 2007). As a result, the majority of material subducted prior to 43 Ma was young ( $\sim 10$  Myr old – Figure 17c), leading to minimal trench motion at the Sunda-Java subduction zone, consistent with ‘I’-type subduction (Figure 18b). As the Wharton ridge ceased spreading and the subducting plate age increased (Figure 17a,b), the trench began to retreat and developed a ‘C’ shape across its  $\sim 5000$  km width, again demonstrating that an ‘I’-type to ‘C’-type transition can occur when subducting plate age increases, consistent with our modelling predictions.

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### 4.3.3 ‘W’-type Trenches

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‘W’-type trenches develop with moderate and older subducting plate age and very wide trenches. The South American trench is the textbook example of a ‘W’-shape (Schellart et al., 2007): it exhibits concave curvature on both edges, with the centre of the trench almost stagnant throughout the Cenozoic (Figure 17c). Subduction in the South Pacific also exhibits ‘W’-type characteristics in the early Cenozoic.

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The South American trench is over 6000 km long, and subducted moderately old material ( $\sim 50$ – $80$  Myr) throughout the Cenozoic at the centre of the trench (Figure 17). Trench evolution shows increasing oroclinal bending through the Cenozoic (Schepers et al., 2017) and, hence, more retreat towards the north and south than in the central part at the Bolivian bend (Figure 18, Müller et al., 2019). The present-day trench shape is typical of our wide plate model predictions, where the Bolivian Orocline protrudes close to the centre of the trench, while sections of the trench close the edges have a concave geometry. The subduction of pre-existing buoyant features on the Nazca Plate likely add complexity to the explanation of its evolution towards the current geometry (e.g., Gutscher, Olivet, et al., 1999; Espurt et al., 2008). The age pattern of the downgoing Nazca Plate and the thickness of the upper plate also potentially influence the orocline, as the topographic symmetry of the Andean mountain belt around Central Andes coincides with the younging of Nazca plate to both north and south directions from Central Andes and the thinning of the upper plate from the center towards the north and south (Capitanio et al., 2011). Thus although it is likely that multiple factors contribute to the shape of

670 the trench at the South America Subduction Zone, our results suggest that the first-order  
671 ‘W’-shape is dictated by its large width and moderate subducting-plate age.

672 The South Pacific region has a more complex tectonic history. In the early Ceno-  
673 zoic, the old Pacific plate was subducting under the South Pacific trench, which had length  
674 exceeding 6000 km (Figure 17c). The trench shape at 60 Ma resembles ‘W’-shape trenches,  
675 where the oldest ( $\sim 100$  Myr) part of the plate was subducting at a region of least trench  
676 retreat, in the northern part of the trench. It had the middle ‘stagnation’ and the south-  
677 ern concave part of the ‘W’-shape, where as the northern concave curvature is not clearly  
678 observed, partly due to the complex tectonic settings to the north (Figure 18, 17c). Here  
679 too, buoyant features, in particular the Ontong-Java plateau (Neal et al., 1997; Mann  
680 & Taira, 2004; Stotz et al., 2017) affected the segmentation of the trench into the New  
681 Hebrides, New Britain and Tonga-Kermadec-Hikurangi trenches (e.g., Pelletier et al., 1998).  
682 However, the shape of the resulting trenches was likely preconditioned by the original  
683 ‘W’ shape.

684 Overall, the examples of ‘I’-, ‘C’- and ‘W’-shape trenches on Earth are in line with  
685 our modeling results. ‘I’-type trenches are associated with very young downgoing plates  
686 of  $\sim 10$  Myr old; and as plate age increases, some transition into ‘C’-shape trenches. ‘W’-  
687 shape trenches are observed in subduction zones exceeding 6000 km width, where older  
688 material (greater than 50 Myr old) is being subducted, thus driving trench retreat. There  
689 is no doubt that the trench shape at each subduction zone is further modulated by ad-  
690 ditional complexities, including variable downgoing plate age along strike, subduction  
691 of buoyant active or bathymetric ridges, and variations in thickness and buoyancy of the  
692 upper plate. Nonetheless, our results demonstrate the key role that both subducting plate  
693 age and width play in controlling the evolution of trench geometry, providing a frame-  
694 work to better understand the evolution of subduction zones.

## 695 5 Conclusions

696 We have presented new 3-D spherical free-subduction models with a composite visco-  
697 plastic plate and viscously layered mantle. We examined the sensitivity of subduction  
698 dynamics and trench evolution to different plate ages (simulated with covarying plate  
699 densities and thicknesses) and plate widths, in both spherical and Cartesian settings.

700 Our models show similar results to previous studies on the effect of age and width  
701 on the evolution of the subduction zone. As plate age increases, plate strength increases  
702 and, as a result, the subduction style transitions from vertically sinking and folding to  
703 retreating with a shallower upper mantle dip angle. Our models produce ‘C’ shaped trenches  
704 for narrower plates and ‘W’ shaped trenches for wider plates, resulting from the toroidal  
705 flow cells at the edge of the retreating subducting plates, consistent with the models of  
706 Schellart et al. (2007). However, we also find that the effect of width is modulated by  
707 the age of the subducting plate. For young plates that are in the vertical folding regime,  
708 the trench does not develop a ‘W’ shape, even for very wide plates. The trench only de-  
709 velops ‘C’ or ‘W’ shapes for retreating cases. Furthermore, for plates that are in the re-  
710 treating regime, a younger plate develops more along-strike variability than an older plate,  
711 due to its lower strength.

712 We find that spherical geometry increases the effective strength of the plate due  
713 to three main factors: (i) the spherical geometry concentrates the buoyancy of subducted  
714 material, leading to faster sinking rates and, accordingly, reducing the time available for  
715 bending at the trench; (ii) the curvature of the mantle transition zone further reduces  
716 the effective dip angle when interacting at 660 km, whereby the slab has more tendency  
717 to retreat and deflect at the transition zone; and (iii) the double curvature of the plate  
718 on a spherical surface adds mechanical strength that resists bending, which is particu-  
719 larly important for wider plates and the evolution of their trenches.

720 Although Cartesian simulations are sufficient to capture the subduction dynam-  
 721 ics of narrow plates (less than approx. 2000 km in width), we now have the means to more  
 722 accurately simulate subduction dynamics on a sphere. This opens up new possibilities  
 723 and will be used in the future to investigate additional factors that affect subduction dy-  
 724 namics and their expression at Earth’s surface.

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 733 [.github.io/](https://fluidityproject.github.io/); the version used for the simulations presented herein has been archived  
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