

# 1 On the rise and fall of Earth's strong clear-sky hemispheric albedo asymmetry

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## 10 11 12 **Abstract**

13  
14 A striking feature of the Earth system is that the Northern and Southern Hemispheres reflect  
15 identical amounts of sunlight. This hemispheric albedo symmetry comprises two asymmetries:  
16 The Northern Hemisphere is more reflective in clear skies, whereas the Southern Hemisphere is  
17 cloudier. The most-cited explanation is that the clear-sky asymmetry is primarily due to the  
18 relatively-bright continents being disproportionately located in the Northern Hemisphere.  
19 However, it is the atmosphere, not the surface, that contributes most to the clear-sky asymmetry.  
20 Here we show that the continent-based component of the clear-sky surface asymmetry is largely  
21 offset by greater reflection from the Southern Hemisphere poles, allowing the clear-sky  
22 asymmetry to be dominated by aerosol. Climate model simulations suggest that aerosol  
23 emissions since the pre-industrial era have driven a large increase in the clear-sky asymmetry  
24 that would reverse in future low-emission scenarios. High-emission scenarios also show a  
25 decrease in asymmetry, but instead driven by declines in Northern Hemisphere ice and snow  
26 cover. Strong clear-sky hemispheric albedo asymmetry is therefore a transient, rather than fixed,  
27 feature of Earth's climate. If all-sky symmetry is maintained despite changes in the clear-sky  
28 asymmetry, compensating cloud changes would have uncertain but important implications for  
29 Earth's energy balance and hydrological cycle.

## 30 31 32 **Introduction**

33  
34 Ever since reliable space-based estimates of Earth's albedo (broadband shortwave reflectivity)  
35 became available in the mid-1960s, it has been observed that the Northern and Southern  
36 Hemispheres (NH and SH, respectively) reflect the same amount of sunlight to within  
37 measurement uncertainty<sup>1-4</sup>. Although this hemispheric albedo symmetry appears to be non-  
38 trivial in a statistical sense<sup>5,6</sup>, at present there exists no generally-accepted physical explanation  
39 for how this symmetry is maintained (if indeed it is maintained, which has not been proven).  
40 State-of-the-art global climate models do not systematically simulate hemispherically symmetric  
41 albedos<sup>3-5,7</sup>, and despite initial findings<sup>1</sup>, Earth's outgoing longwave radiation does not currently  
42 exhibit a similar degree of hemispheric symmetry<sup>3,4</sup>. Prior measurements may have been  
43 inaccurate, or the real situation may have changed as the NH warmed faster than the SH over the  
44 last several decades<sup>8</sup>. The hemispheric imbalance in longwave radiation is balanced by cross-  
45 equatorial heat transport, with a northward oceanic heat transport driven by the Atlantic

46 Meridional Overturning Circulation partially offset by southward atmospheric heat transport  
47 associated with the northward location of the mean Intertropical Convergence Zone (ITCZ)<sup>4,9-11</sup>.  
48

49 Although the all-sky albedo is symmetrical between the hemispheres, its clear-sky and overcast  
50 components are markedly asymmetric, with much greater clear-sky reflection in the NH  
51 balanced by more abundant and brighter clouds in the SH, particularly in the midlatitudes<sup>6,12,13</sup>.  
52 Here we focus on the asymmetric clear-sky component of the all-sky symmetry.  
53

54 The original and most frequently invoked explanation<sup>1-3,6</sup> for the greater NH clear-sky reflection  
55 has to do with the arrangement of the continents: because land surfaces are brighter than the  
56 oceans, and most of the continents are in the NH, the NH clear-sky should be brighter than the  
57 SH. This hypothesis is supported by utilizing the spectral dimension of albedo to attribute the  
58 changes to Earth system properties<sup>14</sup>: the NH is brighter at near-infrared wavelengths associated  
59 with reflection from land surfaces and vegetation, whereas the SH is brighter at the visible  
60 wavelengths associated with reflection from clouds<sup>3</sup>. In this limited view, the NH clear-sky  
61 advantage should be stable on geological (millions of years) timescales.  
62

63 However, the continent-based explanation of Earth's clear-sky hemispheric albedo asymmetry is  
64 substantially incomplete. To start, the atmospheric component of the NH-SH clear-sky  
65 asymmetry is known to be larger than the surface component<sup>3,7</sup>. Here we show that the  
66 atmospheric component dominates the surface component of the clear-sky asymmetry both  
67 because anthropogenic aerosol (airborne particulate matter) enhances atmospheric reflection in  
68 the Northern Hemisphere and because the Antarctic surface is substantially brighter than the  
69 Arctic surface, nearly cancelling the effect of midlatitude and tropical continental surface  
70 reflection. If anthropogenic aerosol and the cryosphere matter for Earth's clear-sky hemispheric  
71 albedo asymmetry as much as the land distribution, the clear-sky asymmetry is more ephemeral  
72 than generally recognized. And if clouds adjust to maintain all-sky albedo symmetry in the face  
73 of clear-sky albedo asymmetry changes (which is not a given), there would be hard-to-predict  
74 ripple effects across the climate system.  
75

## 77 **Results**

### 79 **Atmosphere and cryosphere controls on clear-sky hemispheric albedo asymmetry**

81 To assess the atmospheric and surface contributions to Earth's observed clear-sky albedo, we  
82 separate these components in the Clouds and the Earth's Radiant Energy System (CERES)  
83 Energy Balanced and Filled (EBAF) product<sup>15,16</sup> using a simple single-layer model of shortwave  
84 radiative transfer<sup>17-19</sup> (see Methods). Global maps of total clear-sky reflected shortwave radiation  
85 ( $R_{\text{clr}}$ ) and its atmospheric and surface contributions are shown in Extended Data Figures 1, 2a,  
86 and 3a. Ocean surfaces are extremely dark whereas land is generally brighter, with ice and desert  
87 surfaces in particular reflecting very large quantities of sunlight. Reflection from the atmosphere  
88 is more globally uniform, although clear maxima are evident in areas of high aerosol  
89 concentration (e.g., East Asia, Sahara outflow). Regions of high topography (e.g., the Andes,  
90 Tibet, Antarctica) have minima in atmospheric reflection due to the simple mechanics of there  
91 being a thinner overlying atmosphere than in regions closer to sea-level.

92

93 Figure 1 shows this clear-sky decomposition averaged over each hemisphere, as the average NH  
94 minus SH difference ( $\Delta R_{\text{clr}}$ ), and as the NH-SH difference zonally. In each hemisphere, the  
95 atmosphere contributes approximately 60% of the total clear-sky reflection and the surface  
96 approximately 40%. However, the atmosphere contributes approximately 80% of the  
97 hemispheric contrast, with the surface only contributing 20%, in line with previous findings  
98 using a similar decomposition method<sup>3,7</sup>.

99

100 If the continents are mostly located in the NH, and are brighter than the ocean, what can account  
101 for such a small surface contribution to the clear-sky hemispheric albedo asymmetry? The  
102 answer is: compensation by the cryosphere, specifically by Antarctica. As can be seen from the  
103 markers in Figure 1b and the zonal surface contribution in Figure 1c, the NH continental  
104 advantage is substantial from the tropics to the midlatitudes. At around 60°, however, the  
105 situation reverses dramatically, with the Antarctic reflecting so much more sunlight than the  
106 Arctic that a large portion (~4 W/m<sup>2</sup>) of the tropical and midlatitude continent-based advantage  
107 (~5.5 W/m<sup>2</sup>) is erased. (Tropical, midlatitude, and polar values are defined here as reflection in  
108 each region divided by the full hemispheric area, so the sum of all three regions equals the  
109 hemispheric value.) In contrast, the NH atmosphere is more reflective at all latitudes, with peaks  
110 in the tropics and poles (related in part to the mechanical effect of Antarctica's high topography).

111

112 To better understand the atmospheric component of the clear-sky reflection asymmetry, we  
113 analyze total aerosol optical depth (AOD or  $\tau_a$ ) and its contributions from black carbon (BC),  
114 dust, organic carbon (OC), sea salt, and sulfate (SO<sub>4</sub>) aerosols in the Modern-Era Retrospective  
115 analysis for Research and Applications, Version 2 (MERRA-2) product<sup>20-22</sup> (Fig. 1d; see  
116 Methods).

117

118 The NH dominates AOD in the tropics primarily through a large dust contribution. In the  
119 midlatitudes, sulfate pollution in the NH largely balances sea salt in the SH. Without the (largely  
120 anthropogenic) sulfate contribution, the SH would presumably dominate midlatitude AOD.  
121 Carbonaceous aerosols (mainly OC) slightly favor the SH in the tropics and the NH closer to the  
122 poles. Since dust and sea salt aerosol are largely "natural" in origin, whereas sulfate is largely  
123 due to industrial emissions, it is reasonable to expect that the hemispheric AOD contrast ( $\Delta\tau_a$ ) —  
124 and therefore atmospheric  $\Delta R_{\text{clr}}$  — may have been substantially milder in the pre-industrial  
125 climate.

126

### 127 **From a cleaner past...**

128

129 To test this idea, we analyze output from seven coupled climate models that participated in the  
130 Aerosol Chemistry Model Intercomparison Project (AerChemMIP)<sup>23</sup> "hist-piAer" experiment, in  
131 which aerosol precursor emissions<sup>24</sup> are kept at pre-industrial (PI) values but all else evolves in  
132 the same manner as the "historical" experiment (see Methods). Figure 2 shows the clear-sky  
133 hemispheric albedo asymmetry and its atmospheric and surface contributions from 1850-1865  
134 and 2000-2015 for the historical simulations and from 2000-2015 for the hist-piAer simulations.

135

136 For the present-day (PD) period (2000-2015), the models vary by a few W/m<sup>2</sup> in terms of their  
137 total clear-sky asymmetries, but diverge more radically from the observations in their breakdown

138 between atmospheric and surface reflection (Fig. 2a-c). No model matches the observed  
139 dominance of the atmospheric component over the surface, and some (e.g., MIROC6 and  
140 NorESM2-LM) have the ratio reversed as compared to CERES. Extended Data Figures 2-4 show  
141 the difference in  $R_{\text{clr}}$  (model minus CERES) for the atmosphere ( $R_{\text{clr,atm}}$ ) and surface ( $R_{\text{clr,sfc}}$ )  
142 globally and as averaged over tropical, midlatitude, and polar latitude bands, respectively. The  
143 underestimates in atmospheric  $R_{\text{clr}}$  (Extended Data Fig. 2) are mainly in the tropics (particularly  
144 over South America, Africa, and Arabia) and NH midlatitudes (particularly over eastern North  
145 America and Europe) and the overestimates in surface  $R_{\text{clr}}$  (Extended Data Fig. 3) are  
146 concentrated over the continents. For MIROC6 in particular, dramatic biases in Antarctic sea ice  
147 (Extended Data Fig. 3e) help explain its anomalously low SH surface reflectance (Fig. 2c).

148  
149 These biases notwithstanding, it is apparent that the atmospheric component of the clear-sky  
150 albedo symmetry was much lower (~50-100%) in the PI era (Fig. 2e) or with PI aerosol  
151 precursors (Fig. 2h) in the models. [This is partially offset in the total asymmetry by declining  
152 NH snow and sea ice cover in the historical PD-PI comparisons (Fig. 2d).] Using GISS-E2-1-G  
153 as an example (Fig. 2j-l), timeseries of the full historical and hist-piAer simulations show that the  
154 divergence in the experiments (toward greater asymmetry in the historical total and atmospheric  
155  $R_{\text{clr}}$ ) takes off around the Second World War period (~1935-1950) and is largely complete by the  
156 1960s and 1970s — which happens to be the earliest time period in which we have space-based  
157 observations of Earth's albedo. It is thus possible that we only have reliable observations starting  
158 from a relatively unusual time for Earth's clear-sky albedo asymmetry.

159  
160 The model asymmetry in AOD is highly correlated (Pearson's  $r = 0.93$ ) with the asymmetry in  
161 the atmospheric component of  $R_{\text{clr}}$  (Fig. 3), consistent with a leading role of aerosol in driving  
162 variability in  $R_{\text{clr,atm}}$  over space and time. CERES/MERRA-2 values are a high  $\Delta R_{\text{clr,atm}}$  outlier  
163 compared to the model-based regression fit (see Methods). Relatedly, the fit based on interannual  
164 variability in the asymmetries from CERES and MERRA-2 suggests a much steeper increase in  
165  $\Delta R_{\text{clr,atm}}$  for an increase in  $\Delta\tau_a$  than seen in the model ensemble. This is not likely due to the  
166 different methods of calculating the fit (interannual variability within an O(10 year) period  
167 versus the relationship between different periods): when comparing slopes calculated using only  
168 data from 2000-2015 versus for all years 1850-2015 in each model, there are no systematic  
169 differences (Extended Data Fig. 5). If we therefore assume that the CERES/MERRA-2  
170 interannual slope is representative of PD-PI differences, as it is for the models, we can then  
171 estimate the PI value of the clear-sky atmospheric albedo asymmetry for a given PI aerosol  
172 asymmetry.

173  
174 To estimate the PI aerosol asymmetry, we use the good correlation ( $r = 0.87$ ) between the global  
175 mean AOD in the PD for a given model and its PD-PI change in  $\Delta\tau_a$  (Extended Data Fig. 6) as an  
176 emergent constraint. Via Monte Carlo simulation (see Methods), we find that the PI value of  
177  $\Delta R_{\text{clr,atm}}$  was  $2.4 \text{ W/m}^2$  (95% confidence interval of  $0.7 \text{ W/m}^2$  to  $3.9 \text{ W/m}^2$ ), around half the PD  
178 value of  $4.9 \pm 0.4 \text{ W/m}^2$ . This would represent a substantial decrease in the total clear-sky  
179 hemispheric albedo asymmetry in the PI compared to that observed today.

180  
181  
182 **...toward either a less polluted or less icy future**  
183

184 If the past had weaker contrast than the present, what about the future? Under very different  
185 future scenarios within the Shared Socioeconomic Pathways (SSPs)<sup>25,26</sup> there is one consistent  
186 outcome: the clear-sky hemispheric albedo asymmetry is projected to decline in the coming  
187 century (Fig. 4).

188  
189 For the low-emission SSP1-2.6 ("sustainability") scenario, the atmospheric component of  $R_{\text{clr}}$   
190 drives a decline in overall asymmetry (Fig. 4ab) while the surface plays a more minor role (Fig.  
191 4c), consistent with a decline in co-emitted aerosols and precursor gases. In contrast, the high-  
192 emission SSP3-7.0 ("regional rivalry") scenario gets the same overall result (Fig. 4g) but driven  
193 by the surface (Fig. 4i), rather than the atmosphere (Fig. 4h), consistent with maintained high  
194 emissions of aerosols and their precursors. Results for the intermediate SSP2-4.5 ("middle of the  
195 road") scenario are, fittingly, a blend of those from the two other scenarios (Fig. 4d-f). As  
196 illustrated by the UKESM1-0-LL results (Fig. 4j-l), the divergence in the scenarios is apparent  
197 by midcentury.

198  
199 The surface changes in SSP3-7.0 are largely a story of sea ice (Fig. 5). The hemispheric contrast  
200 in sea ice area (see Methods) has a good correlation ( $r = 0.77$ ) with the hemispheric contrast in  
201 the surface component of  $R_{\text{clr}}$ . All models show a decline in Arctic sea ice and NH snow and ice  
202 cover on land with warming, but the magnitude and even sign of Antarctic sea ice changes are  
203 more variable (Extended Data Fig. 7). This is in part a mean state issue: models like MIROC6  
204 and NorESM2-LM that have small amounts of Antarctic sea ice in the present day (Extended  
205 Data Fig. 3e,g) have limited room for future declines.

206  
207 Changes in sea ice albedo<sup>27</sup> and in snow and ice cover on land<sup>28</sup> may also factor into the surface  
208 asymmetry changes. For instance, despite nearly equally-balanced trends in NH and SH sea ice  
209 area in MRI-ESM2-0 (Fig. 5, Extended Data Fig. 7e,l), there is a modest decline in surface  
210 asymmetry associated with decreased reflection over northern and western North America and  
211 the Himalayan highlands and Tibetan Plateau.

212  
213 Due to large internal variability in sea ice concentration<sup>29-31</sup> and outstanding questions about sea  
214 ice dynamics, particularly in the Antarctic<sup>32-35</sup>, how the surface component of the clear-sky  
215 albedo asymmetry would change in reality in a high-warming scenario is subject to more  
216 uncertainty than the aerosol-driven atmospheric component. However, it is very plausible that  
217 the  $R_{\text{clr,sfc}}$  asymmetry could substantially decline and even reverse in the future if Arctic sea ice  
218 and NH land snow and ice cover are lost more rapidly than Antarctic sea ice in a warming  
219 climate.

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221

## 222 Discussion

223  
224 The changing nature of Earth's clear-sky albedo asymmetry should be observable in the coming  
225 decades under any of the future scenarios considered, especially with the planned launch of  
226 several visible-shortwave infrared spectroradiometers that will allow for spectral decomposition  
227 of reflected sunlight where the signal may be larger<sup>36</sup>. We run simplified radiative transfer  
228 simulations (see Methods) to investigate the spectral signal of a decrease in Northern  
229 Hemisphere aerosol loading (Extended Data Fig. 8). A cleaner Northern Hemisphere becomes

230 less reflective in the visible spectrum (more associated with reflection from the atmosphere) but  
231 has a smaller change in the near-infrared (more associated with reflection from the land surface  
232 and vegetation) for all but the highest solar zenith angles. This will exacerbate the present  
233 hemispheric differences in the visible and near-infrared reflection<sup>2</sup>, which are two portions of the  
234 spectrum that will be directly observed separately and with high accuracy during the upcoming  
235 Earth radiation budget satellite mission, Libera.

236  
237 It is tempting to think that we may have glimpsed a lower aerosol and clear-sky asymmetry  
238 future during 2020 as a result of the societal response to the COVID-19 pandemic<sup>37,38</sup>. Indeed,  
239 2020 featured low outliers in clear-sky hemispheric albedo asymmetry and aerosol contrast  
240 values (Fig. 2j-k, Fig. 3, Fig. 4j-k). However, the situation is more complicated as the aerosol  
241 changes associated with the pandemic lockdowns and economic slowdown were likely too small  
242 to be clearly distinguishable above background variability<sup>39-41</sup> and 2020 also featured  
243 anomalously large aerosol loadings over the Southern Ocean from the 2019-2020 Australian  
244 bushfires<sup>42</sup>. Future work is merited to better understand the unique conditions in 2020 and to  
245 what extent they can be used as an "opportunistic experiment" to constrain aerosol radiative  
246 forcing<sup>43</sup>.

247  
248 If the observed hemispheric all-sky albedo symmetry is merely the result of chance, these results  
249 would remain primarily of academic interest. Indeed, without any compensating mechanisms, we  
250 should observe an asymmetry in all-sky reflection in the next few decades and thus a definitively  
251 negative answer to the question of whether Earth's hemispheric all-sky albedo symmetry is  
252 maintained. However, if clouds respond to the changing clear-sky contrast to maintain all-sky  
253 symmetry over the coming decades, there would be important implications for radiative forcing  
254 and hydrological and circulation changes depending on the (currently unknown) adjustment  
255 mechanism.

256  
257 If SH clouds darken to compensate for the NH trend, the resulting positive radiative forcing  
258 would accelerate global warming. Some worrying evidence that such a hemispheric connection  
259 exists comes from CERES reflected shortwave measurements over the past two decades that  
260 show nearly equal all-sky darkening trends in the NH and SH<sup>6,44,45</sup>. A strong decline in  
261 cloudiness within the northeastern Pacific stratocumulus deck and reductions in aerosol from  
262 eastern North America and eastern Asia offer an explanation for the NH trends<sup>45-47</sup>, but there is  
263 no clear driver for the identical SH trend<sup>6</sup>. Alternatively, if NH clouds brighten to compensate  
264 for the clear-sky darkening, global radiative implications could be minor. Of course, the true  
265 response (if any) may involve some combination of both NH cloud brightening and SH  
266 darkening.

267  
268 One hypothesized adjustment mechanism involves shifts in the ITCZ and thus tropical  
269 cloudiness toward the darker hemisphere<sup>48</sup>. If this were to occur, it would have important  
270 regional implications beyond the global mean precipitation shift<sup>49</sup>, particularly for drought-  
271 vulnerable locations like the Sahel<sup>50-52</sup>. More recently, attention has shifted to the role played by  
272 the extremely cloudy SH midlatitude oceans<sup>6,7</sup>. Changes in Southern Ocean cloudiness could  
273 affect large-scale atmosphere-ocean circulations<sup>13,53-55</sup> and long-term global warming via  
274 changes in ocean heat uptake in the Southern Ocean<sup>56-58</sup>. Given the likelihood of large changes in  
275 the clear-sky hemispheric albedo asymmetry this century under any plausible emissions scenario,

276 determining which of these or other as-yet-unidentified mechanisms would likely operate to  
277 maintain the all-sky hemispheric albedo symmetry should be a research priority.  
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447  
448

449 **Methods**

450

451 **Reflected shortwave radiation data**

452

453 Clear-sky shortwave fluxes from January 2003 to December 2020 come from CERES EBAF  
454 Edition 4.1 and are estimated for the total region (including both cloudy and clear scenes) rather  
455 than for only cloud-free portions of scenes using a regional monthly adjustment factor that  
456 accounts for the difference between computed fluxes with cloud effects removed and those  
457 fluxes when weighted by observed clear-sky fraction<sup>15,16,59</sup>. Clear-sky fluxes estimated in this  
458 manner are more comparable with clear-sky output from climate models. Results using clear-sky  
459 fluxes from cloud-free portions of scenes only are similar to those shown here.

460

461 CERES instruments measure filtered radiances in the shortwave spectrum from 0.3 to 5  $\mu\text{m}$  and  
462 fly aboard NASA's polar-orbiting Terra and Aqua satellites as well as the Suomi National Polar-  
463 Orbiting Partnership and NOAA-20 satellites<sup>15</sup>. We select data from 2003-2020 in which both  
464 Terra and Aqua measurements are available. Data from geostationary satellites are used to  
465 correct for the full diurnal cycle and a one-time adjustment (within the range of observational  
466 and calibration uncertainty) is applied to ensure that the measured net imbalance in top-of-  
467 atmosphere (TOA) radiation matches values from in situ observations of ocean heat  
468 uptake<sup>15,60,61</sup>. Surface irradiances are computed independently using aerosol, cloud, and  
469 thermodynamic properties from satellite observations and reanalysis products and are  
470 constrained by the TOA irradiances<sup>16</sup>.

471

472 Uncertainty in the temporal mean values discussed is quantified using the interannual variability  
473 assuming a red noise process<sup>62</sup>. Measurement uncertainties are neglected. This approach has the  
474 main advantage of allowing us to quantify uncertainty identically between the CERES  
475 observations and the CMIP6 models. It is justified because random measurement errors on the  
476 order of 1-10  $\text{W}/\text{m}^2$  per  $1^\circ \times 1^\circ$  monthly grid box<sup>15,16</sup> rapidly diminish when averaging  
477 hemispherically or globally for long time periods [errors of  $O(0.001-0.01 \text{ W}/\text{m}^2)$  as compared to  
478 errors of  $O(0.1-1 \text{ W}/\text{m}^2)$  for temporal averaging assuming red noise] and while systematic errors  
479 would be more concerning in an absolute sense<sup>17</sup>, they would not affect conclusions drawn on  
480 the atmosphere/surface breakdown or on hemispheric differences.

481

482 Spatiotemporal weighted averaging is performed accounting for the fact that months have  
483 slightly different lengths and that the Earth is oblate, not perfectly spherical. Failure to properly  
484 weight by days per month and area can result in errors of  $O(0.1 \text{ W}/\text{m}^2)$  in globally and  
485 hemispherically averaged values.

486

487 **Aerosol reanalysis data**

488

489 Total AOD at 550 nm from MERRA-2 is constrained by assimilation of AOD as retrieved by the  
490 Moderate Resolution Imaging Spectroradiometer instrument aboard the Terra and Aqua  
491 satellites, in addition to several other satellite instruments and the AERONET ground sites, but  
492 the breakdown into different species is only constrained indirectly through the total AOD  
493 constraint<sup>21</sup>. We therefore place greater emphasis on and have greater confidence in the total  
494 AOD values than their species decomposition. MERRA-2 does compare well overall with

495 unassimilated satellite and aircraft measurements of aerosol column optical properties and  
496 vertical extinction profiles, however, lending some greater confidence<sup>22</sup>. MERRA-2 AOD  
497 behaves similarly to other reanalysis products and generally compares well with various  
498 observational datasets<sup>63</sup>, making it unlikely that the choice to focus on MERRA-2 as opposed to  
499 another equally suitable product has any bearing on our results or conclusions. Uncertainty in  
500 temporal mean values is quantified assuming red noise<sup>62</sup>, as for the reflection data. MERRA-2  
501 data is analyzed from January 2003 to December 2020 to match the CERES record.

502

### 503 **Sea ice concentration data**

504

505 Sea ice area data from passive microwave remote sensing observations from January 2003 to  
506 December 2020 come from the National Snow and Ice Data Center (NSIDC) Sea Ice Index  
507 Version 3 product<sup>64</sup>. Weighting sea ice area by insolation improves its correlation with  $R_{\text{clr,sfc}}$  for  
508 each hemisphere separately but has a negligible impact on the hemispheric difference.

509

### 510 **Climate model data**

511

512 Seven state-of-the-art global climate models (abbreviated names in parentheses) from the  
513 Coupled Model Intercomparison Project Phase 6 (CMIP6) archive<sup>65</sup> are selected based on their  
514 participation in the Aerosol Chemistry Model Intercomparison Project (AerChemMIP) hist-  
515 piAer experiment<sup>23</sup> and the Scenario Model Intercomparison Project (ScenarioMIP) SSP1-2.6,  
516 SSP2-4.5, and SSP3-7.0 experiments<sup>26</sup>: NOAA Geophysical Fluid Dynamics Laboratories  
517 GFDL-ESM4 (GFDL)<sup>66-68</sup>; NASA Goddard Institute for Space Studies GISS-E2-1-G (GISS)<sup>69-  
518 71</sup>; Institut Pierre-Simon Laplace IPSL-CM6A-LR (IPSL)<sup>72-74</sup>; University of Tokyo, National  
519 Institute for Environmental Studies, and Japan Agency for Marine-Earth Science and  
520 Technology MIROC6 (MIROC)<sup>75-77</sup>; Japan Meteorological Agency Meteorological Research  
521 Institute MRI-ESM2-0 (MRI)<sup>78-80</sup>; Norwegian Earth System Model Climate Modeling  
522 Consortium NorESM2-LM (NorESM)<sup>81-83</sup>; and the UK Met Office Hadley Centre-Natural  
523 Environment Research Council UKESM1-0-LL (UKESM)<sup>84-86</sup>.

524

525 For models with multiple variants, only one is selected for analysis per model: r1i1p1f1 (GFDL,  
526 IPSL, MIROC6, MRI, NorESM); r1i1p3f1 (GISS); and r1i1p1f2 (UKESM).

527

528 Temporal averaging accounts for the different calendars used by each model (Gregorian for  
529 IPSL, MIROC, and MRI; Gregorian without leap years for GFDL, GISS, and NorESM; and  
530 uniform 30-day months for UKESM) and spatial averaging uses atmospheric grid box area for  
531 the radiation and aerosol (AOD at 550 nm) fields and either the atmospheric or oceanic grid box  
532 area for sea ice depending on the model and its archived output. Not weighting by days per  
533 model month can result in errors of  $O(0.01-0.1 \text{ W/m}^2)$  in globally and hemispherically averaged  
534 values. Uncertainty in temporal means is calculated assuming a red noise process, as in the  
535 observations.

536

### 537 **Decomposition of top-of-atmosphere reflection into atmospheric and surface components**

538

539 Following Donohoe & Battisti<sup>17</sup>, we calculate the atmospheric component of the top-of-  
540 atmosphere (TOA) planetary albedo using the relation:

541

$$542 \quad A = \alpha_{\text{atm}} + \alpha_{\text{sfc}} \frac{\mathcal{T}^2}{(1 - \alpha_{\text{atm}} \alpha_{\text{sfc}})}, \quad (1)$$

543

544 where  $A$  is the planetary albedo (calculated as the ratio of upwelling to downwelling shortwave  
545 radiation at TOA),  $\alpha_{\text{atm}}$  is the atmospheric component of the planetary albedo,  $\alpha_{\text{sfc}}$  is the surface  
546 albedo (calculated as the ratio of upwelling to downwelling shortwave radiation at the surface),  
547 and  $\mathcal{T}$  is the atmospheric transmissivity (calculated as the ratio of downwelling radiation at the  
548 surface to that at TOA). We then calculate  $R_{\text{clr}}$  and its atmospheric and surface components  
549 ( $R_{\text{clr,atm}}$  and  $R_{\text{clr,sfc}}$ , respectively) by multiplying by the incoming solar radiation flux,  $F_{\odot}$ :

550

$$551 \quad R_{\text{clr}} = R_{\text{clr,atm}} + R_{\text{clr,sfc}} = F_{\odot} \alpha_{\text{atm}} + F_{\odot} \alpha_{\text{sfc}} \frac{\mathcal{T}^2}{(1 - \alpha_{\text{atm}} \alpha_{\text{sfc}})}. \quad (2)$$

552

553 The atmosphere/surface reflection decomposition method is identical between CERES and the  
554 CMIP6 models.

555

### 556 **Pre-industrial aerosol contrast estimate**

557

558 In order to calculate the PI value of the atmospheric component of the hemispheric asymmetry in  
559  $R_{\text{clr}}$ , we need to estimate the PI value of the hemispheric AOD asymmetry and be able to relate  
560 the AOD and  $R_{\text{clr,atm}}$  asymmetries.

561

562 To estimate the PI aerosol contrast, we use the ordinary least squares (OLS) regression between  
563 the PD value (defined as the 2000-2015 average) of global mean AOD and the difference  
564 between the hemispheric AOD asymmetry in the PD from the PI (defined as the 1850-1865  
565 average) in the seven CMIP6 models as an emergent constraint (Extended Data Fig. 6). All  
566 averages are inclusive of the starting year and exclusive of the ending year (e.g., the 2000-2015  
567 average includes all months from January 2000 to December 2014). Using the emergent  
568 constraint, we then estimate the real PI aerosol contrast using the PD global mean value from  
569 MERRA-2.

570

571 We use the OLS regression between  $\Delta R_{\text{clr,atm}}$  from CERES and  $\Delta \tau_{\text{a}}$  from MERRA-2 to estimate  
572 the PI value of the atmospheric component of the  $R_{\text{clr}}$  asymmetry. Based on the similarity  
573 between the 2000-2015 and 1850-2015 regressions from the CMIP6 models (Extended Data Fig.  
574 5), the modern CERES-MERRA-2 relationship should be valid for extrapolation back to the PI  
575 era.

576

577 To quantify uncertainty, we use Monte Carlo simulation to generate 10,000 estimates by  
578 randomly drawing from t-distributions of the PD global mean AOD value from MERRA-2 (error  
579 calculated assuming red noise), the PD-PI difference in hemispheric AOD contrast (error  
580 calculated from the OLS regression uncertainty), the PD hemispheric contrast in AOD from  
581 MERRA-2 (error calculated assuming red noise) to calculate the PI hemispheric AOD contrast,  
582 and finally the PI value of the  $R_{\text{clr,atm}}$  asymmetry (error calculated from the OLS regression  
583 uncertainty). Kernel density estimation of the Monte Carlo results is used for presentation  
584 purposes in Figure 3.

585

586 **Spectral albedo calculations**

587

588 Spectrally-resolved radiative transfer output used in Extended Data Figure 8 is calculated with  
589 1D DIScrete Ordinate Radiative Transfer (DISORT) using the Santa Barbara DISORT  
590 Atmospheric Radiative Transfer (SBDART) program<sup>87</sup>. Calculations are performed from 0.2-3.0  
591  $\mu\text{m}$  at 0.005- $\mu\text{m}$  spectral sampling. For simplicity, the NH and SH results are derived from a  
592 weighted combination of three calculations that each use standard surface and aerosol properties  
593 built into SBDART: (1) snow surface and tropospheric aerosol, (2) ocean surface and oceanic  
594 aerosol, and (3) vegetated surface and rural aerosol. The averaging weights for the NH/SH are  
595 (1) – 10%/10%, (2) – 58%/77%, and (3) – 32%/13%, respectively<sup>6</sup>. AOD is set to 0.2 for present  
596 day NH, 0.1 for present day SH, and 0.15 for clean NH. All calculations use the US Standard  
597 atmosphere.

598

599 **Data availability**

600

601 CERES data are available from the NASA Langley Research Center  
602 (<https://ceres.larc.nasa.gov/data/>). MERRA-2 data are available from the NASA Goddard Earth  
603 Sciences Data and Information Services Center  
604 (<https://disc.gsfc.nasa.gov/datasets?project=MERRA-2>). The Sea Ice Index is available from the  
605 NSIDC (<https://nsidc.org/data/G02135/versions/3>). CMIP6 data are available from the Earth  
606 System Grid Federation (ESGF) and were downloaded from the US Department of  
607 Energy/Lawrence Livermore National Laboratory node ([https://esgf-  
608 node.llnl.gov/projects/cmip6/](https://esgf-node.llnl.gov/projects/cmip6/)).

609

610 **Code availability**

611

612 All python libraries used in the analysis (cartopy<sup>88</sup>, matplotlib<sup>89</sup>, numpy<sup>90</sup>, scipy<sup>91</sup>, and xarray<sup>92</sup>)  
613 are freely available. The SBDART code is available from Paul Ricchiazzi  
614 (<https://github.com/paulricchiazzi/SBDART>).

615

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632

633 **Author contributions**

634

635 M.S.D. conceived and designed the study with input from J.J.G., J.E.K., and G.F. M.S.D.  
636 performed all analyses except for the radiative transfer calculations, which were performed by  
637 J.J.G. M.S.D. wrote the manuscript with input and editing from all coauthors.

638

639 **Competing interests**

640

641 The authors declare no competing interests.

642

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644

645

646 **Methods References**

647

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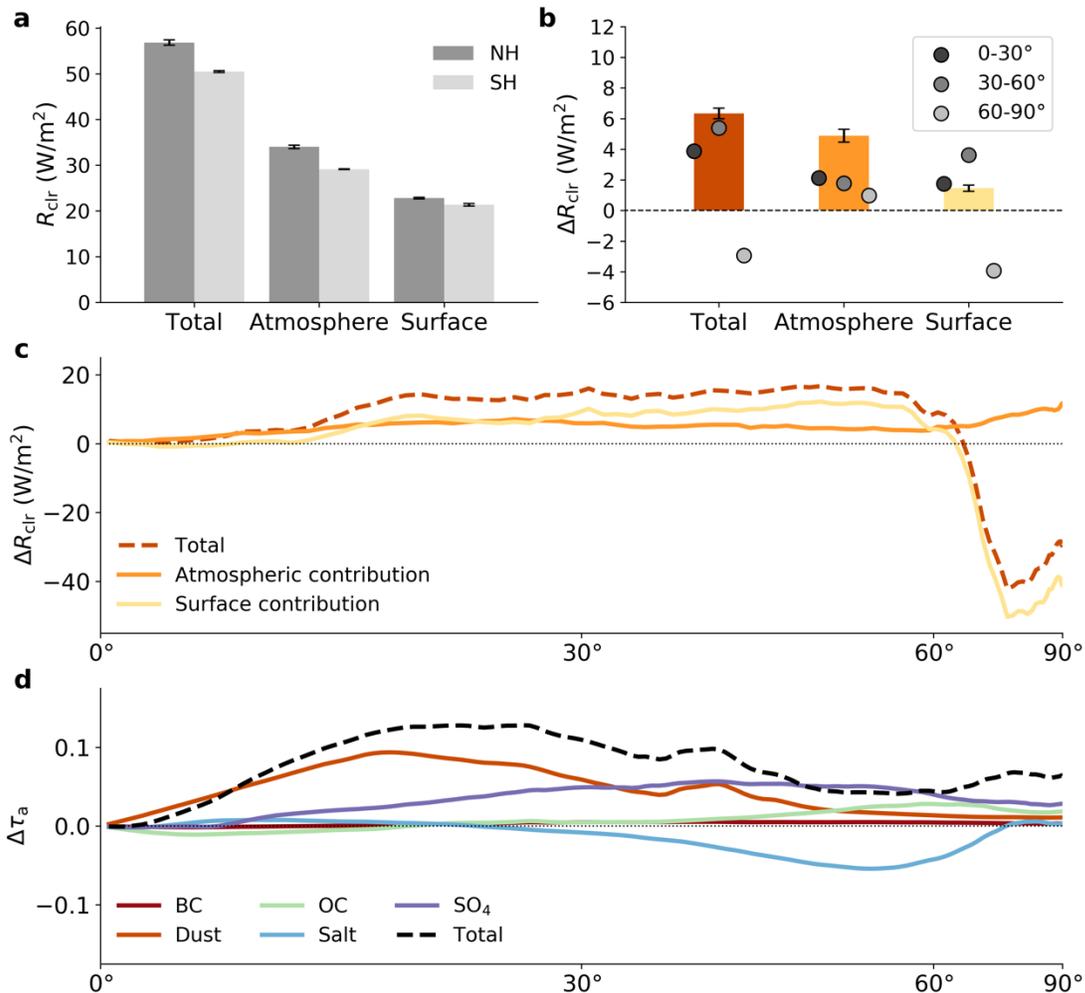
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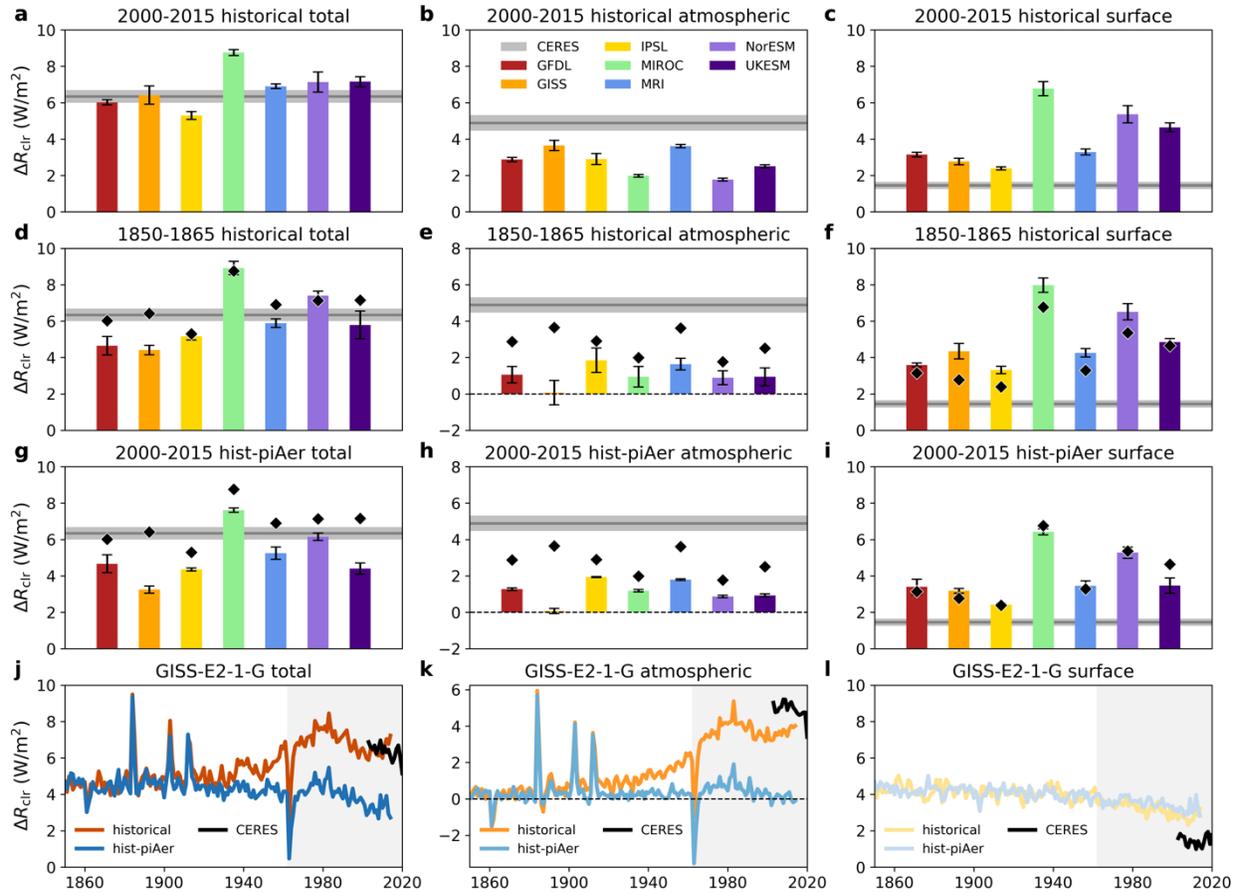
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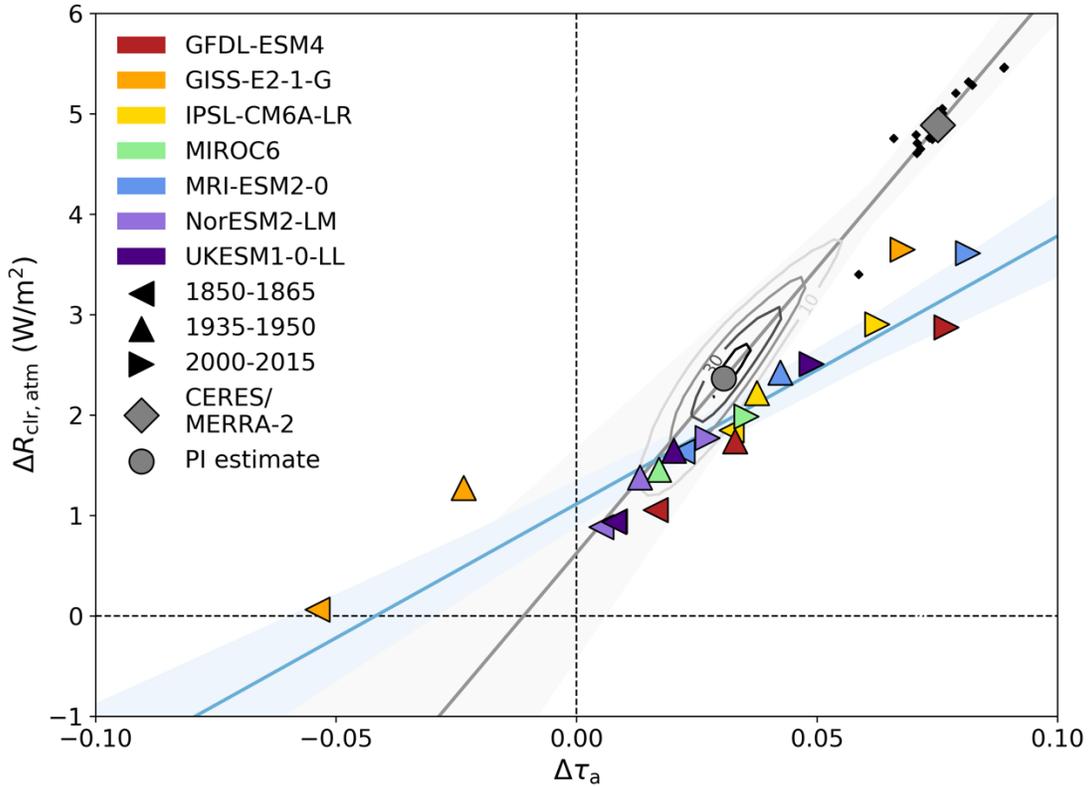


788 **Fig. 1 | The atmosphere, not the surface, contributes most to the observed clear-sky**  
 789 **hemispheric albedo asymmetry. a-c**, Hemispherically-averaged clear-sky reflected solar  
 790 radiation (a), the hemispheric difference (Northern Hemisphere minus Southern Hemisphere)  
 791 (b), and zonal differences (c). Markers in b indicate the component of the total hemispheric  
 792 difference attributable to the tropics (0-30°), midlatitudes (30-60°), and poles (60-90°). d, Zonal  
 793 hemispheric difference for total AOD and each species. Error bars in a-b represent 95%  
 794 confidence in the mean value. For c-d, the abscissa is area-weighted (plotted as sine of latitude).  
 795 All averages are for 2003-2020, inclusive.  
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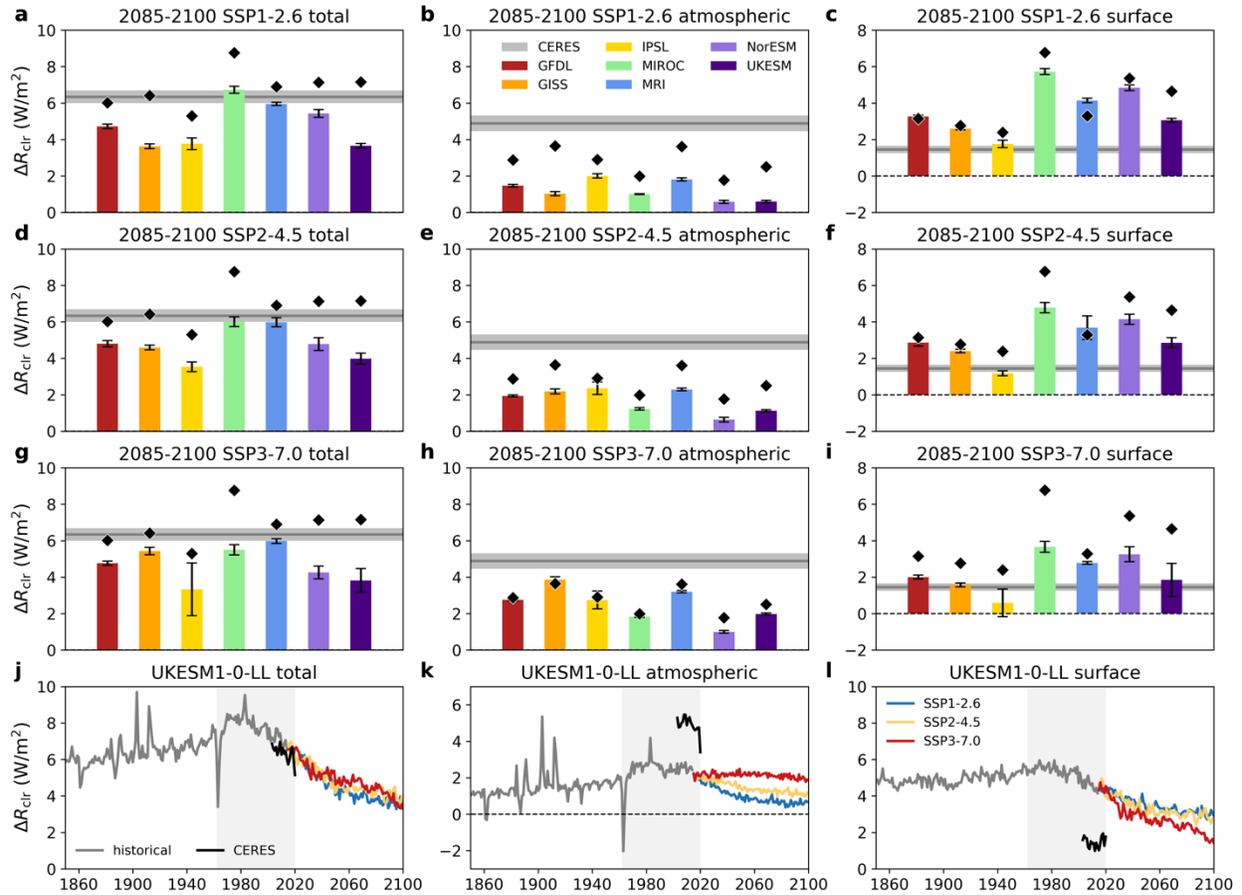


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800 **Fig. 2 | Clear-sky hemispheric albedo asymmetry changes in historical CMIP6 runs.**  
 801 Average of the Northern Hemisphere minus Southern Hemisphere total clear-sky reflection and  
 802 its atmospheric and surface contributions for the 2000-2015 period in the historical runs (a-c),  
 803 1850-1865 period in the historical runs (d-f), and 2000-2015 in the hist-piAer runs (g-i). Error  
 804 bars for each model represent 95% confidence in the mean value. Diamond markers in d-i  
 805 represent each model's 2000-2015 historical run mean value for reference. CERES mean values  
 806 are shown as gray lines with shading representing the 95% confidence interval. j-l, Example time  
 807 series of historical and hist-piAer total, atmospheric, and surface reflection asymmetries from the  
 808 GISS-E2-1-G model. Shading represents the time period in which reliable space-based estimates  
 809 of Earth's albedo have been available.  
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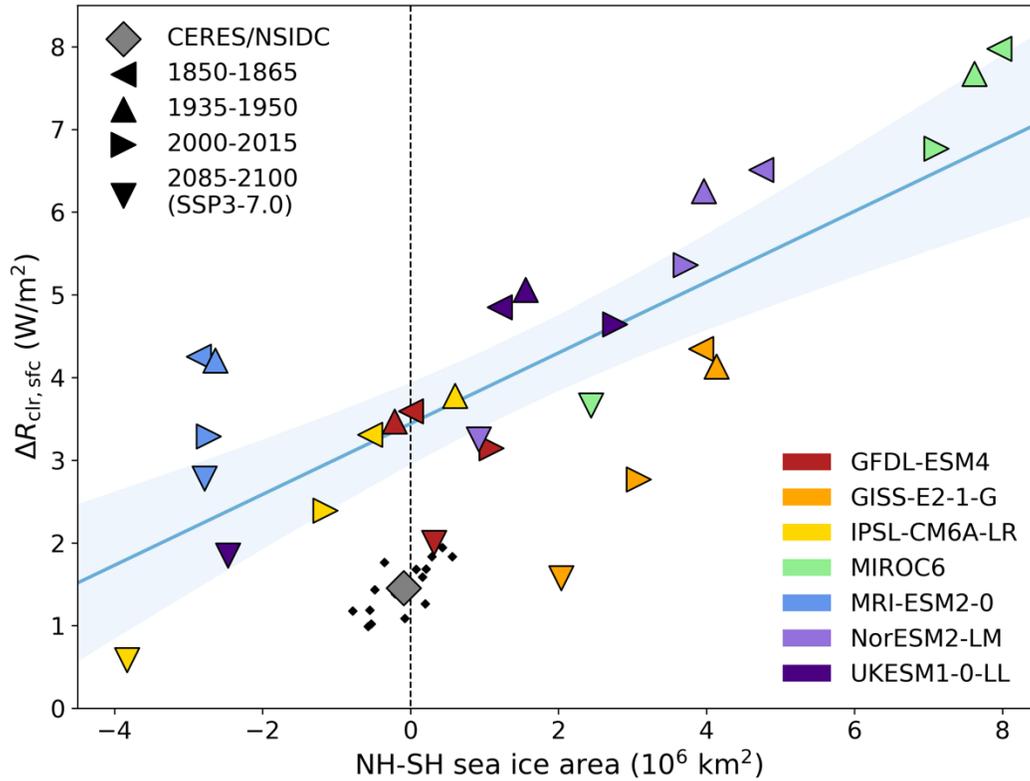


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 812 **Fig. 3 | Relationship between the hemispheric aerosol and atmospheric reflection**  
 813 **asymmetries in the pre-industrial, midcentury, and present-day.** CMIP6 model values from  
 814 the historical runs are represented as colored triangles (facing left for the 1850-1865 mean, up for  
 815 1935-1950, right for 2000-2015) and their regression fit and its 95% confidence interval are  
 816 represented by the blue line and shading. CERES/MERRA-2 data are represented as a gray  
 817 diamond for the modern mean and as smaller black diamonds for individual years. The  
 818 regression fit between individual CERES/MERRA-2 years and its 95% confidence interval is  
 819 represented by the gray line and shading. Contours represent kernel density estimates of the  
 820 Monte Carlo probabilities (shown every 10 counts) for calculating the pre-industrial value of  
 821  $\Delta R_{\text{clr,atm}}$ , with the gray circle representing the mean value.  
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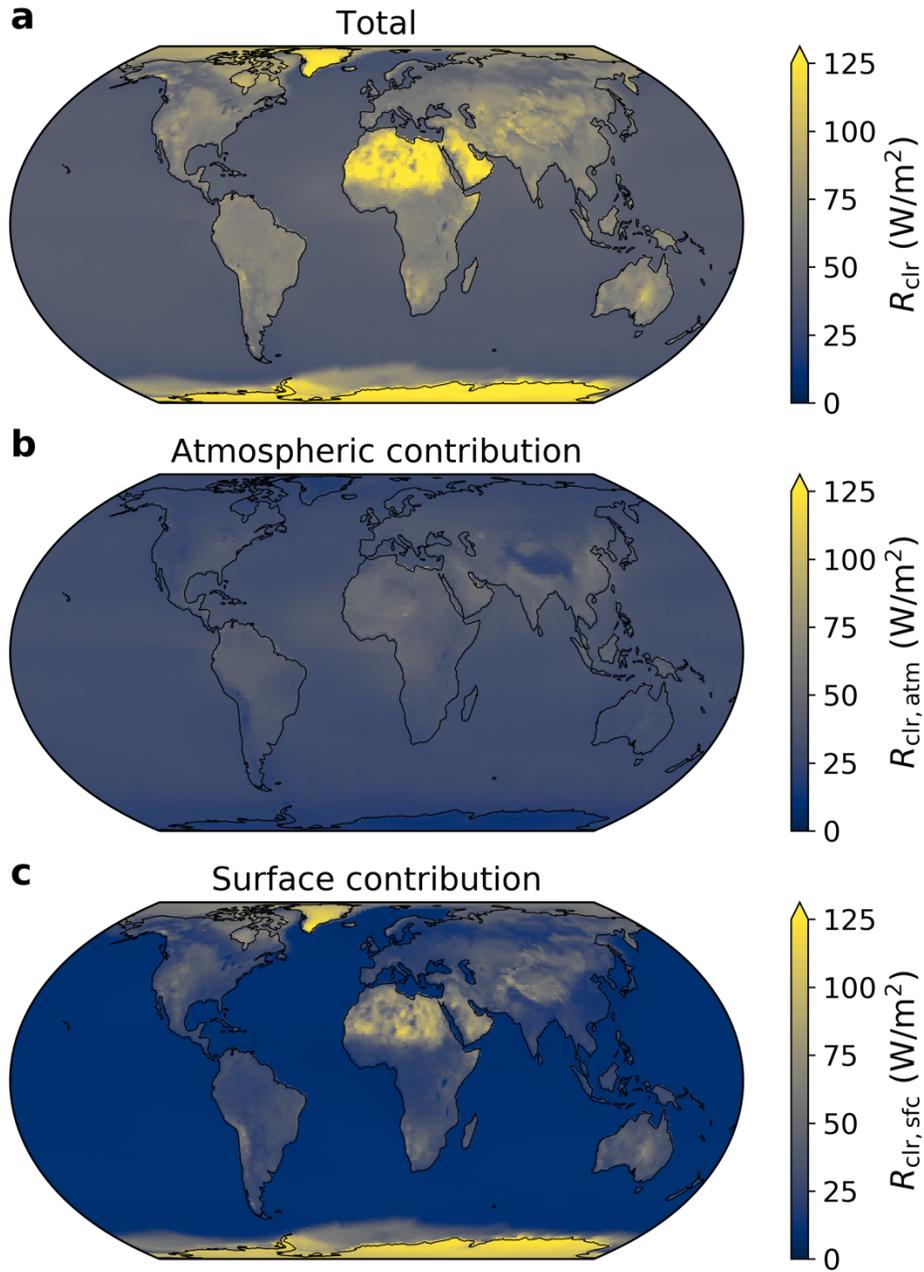
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**Fig. 4 | Clear-sky hemispheric albedo asymmetry changes in CMIP6 runs of future scenarios. a-i**, as in Fig. 2 d-i, but for the SSP1-2.6 (a-c), SSP2-4.5 (d-f), and SSP3-7.0 (g-i) runs. **j-l**, as in Fig. 2 j-l, but for the UKESM1-0-LL model.



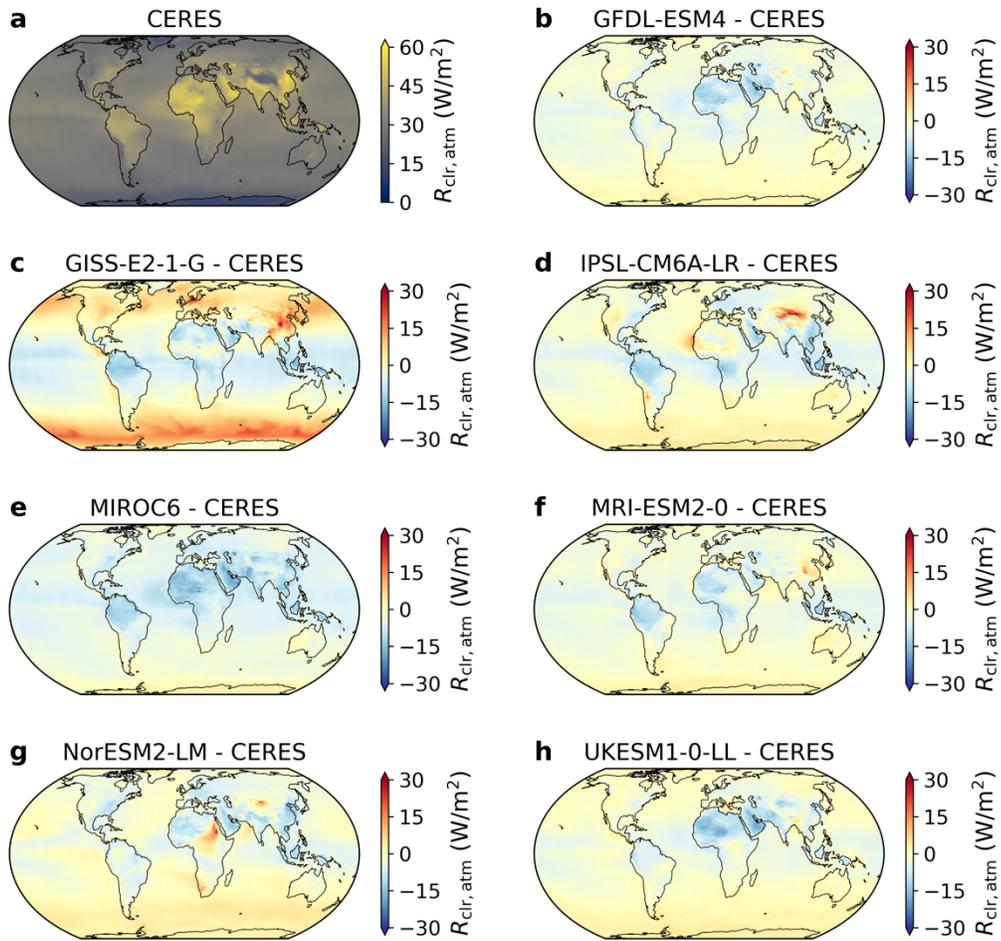
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**Fig. 5 | Relationship between the hemispheric sea ice and surface reflection asymmetries in the pre-industrial, midcentury, present-day, and high-emissions future.** CMIP6 model values from the historical and SSP3-7.0 runs are represented as colored triangles (facing left for the 1850-1865 mean, up for 1935-1950, right for 2000-2015, down for 2085-2100) and their regression fit and its 95% confidence interval are represented by the blue line and shading. CERES/National Snow and Ice Data Center (NSIDC) data are represented as a gray diamond for the modern mean and as smaller black diamonds for individual years.

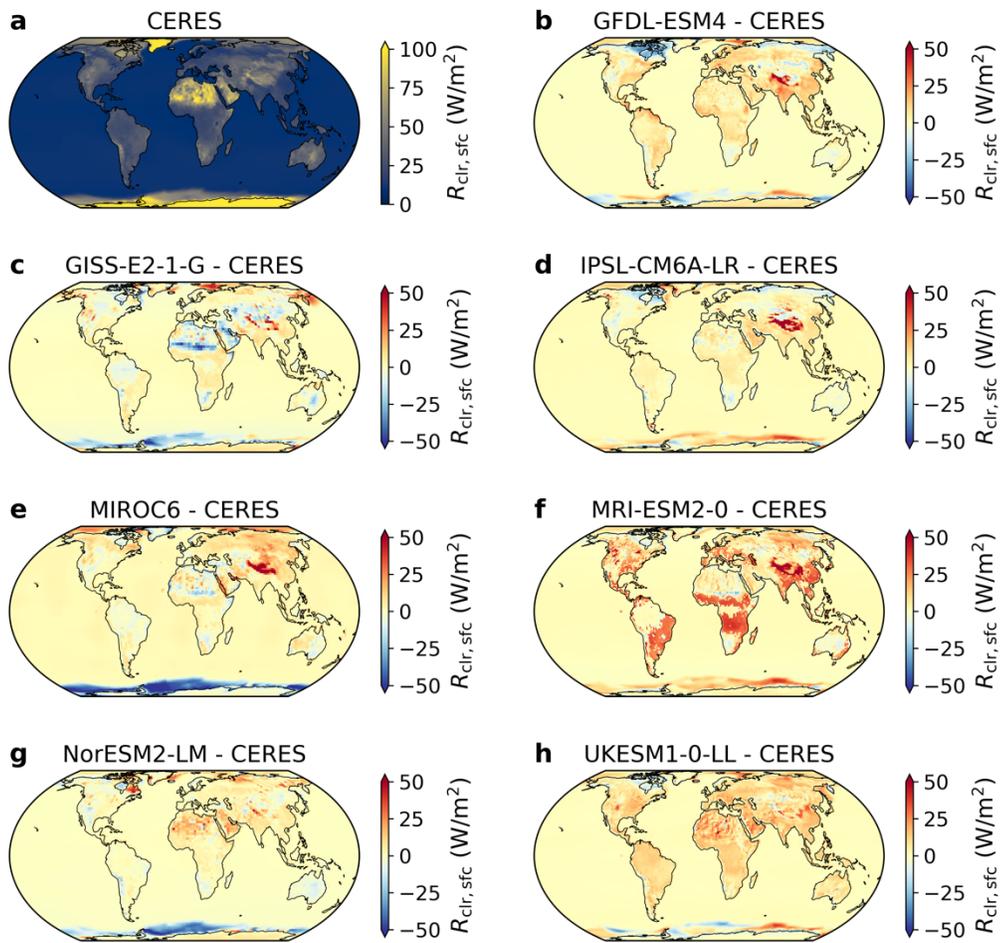


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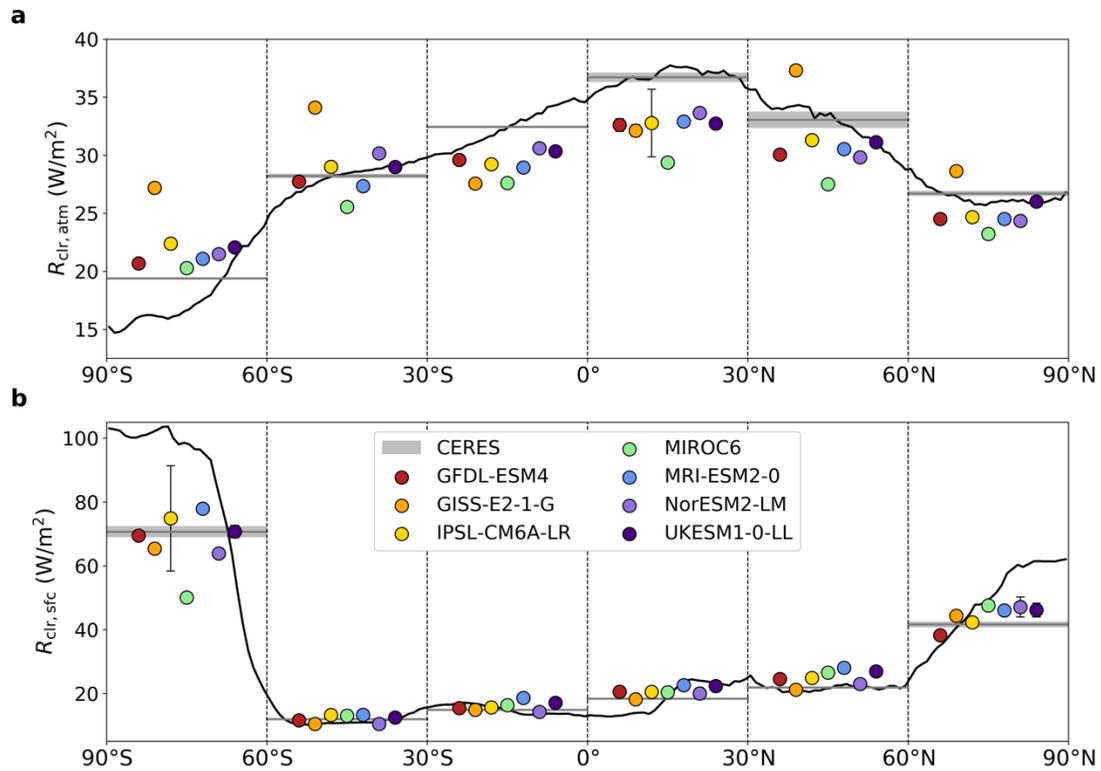
**Extended Data Fig. 1 | Maps of clear-sky reflection.** Total  $R_{\text{clr}}$  (a) and its atmospheric (b) and surface (c) components are shown globally on an equal-area projection<sup>93</sup>.



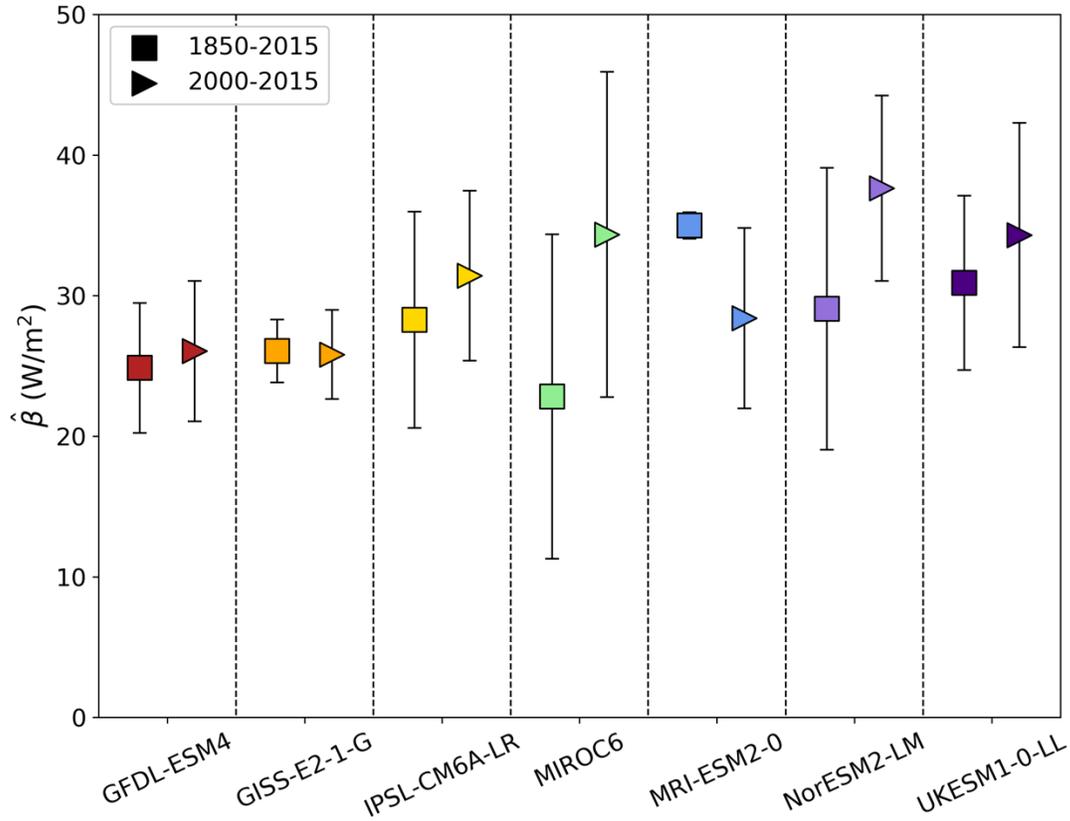
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 846 **Extended Data Fig. 2 | Maps of the atmospheric contribution to clear-sky reflection for**  
 847 **CERES and the CMIP6 models.** Observed  $R_{\text{clr,atm}}$  from CERES (a) and the difference between  
 848 the observed value and each of the CMIP6 models analyzed (b-h) are shown globally on an  
 849 equal-area projection<sup>93</sup>.  
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851  
 852 **Extended Data Fig. 3 | Maps of the surface contribution to clear-sky reflection for CERES**  
 853 **and the CMIP6 models.** Observed  $R_{\text{clr},\text{sfc}}$  from CERES (a) and the difference between the  
 854 observed value and each of the CMIP6 models analyzed (b-h) are shown globally on an equal-  
 855 area projection<sup>93</sup>.  
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 858 **Extended Data Fig. 4 | Zonal differences between CERES and the CMIP6 observations. a,**  
 859 **mean CERES value and 95% confidence interval are represented by the gray line and shading**  
 860 **and mean CMIP6 model value and 95% confidence interval are represented by the circular**  
 861 **markers and error bars for the atmospheric component of the clear-sky reflection for the**  
 862 **Southern Hemisphere poles (90°-60° S), midlatitudes (60°-30° S), and tropics (30° S-0°) and the**  
 863 **Northern Hemisphere tropics (0°-30° N), midlatitudes (30°-60° N), and poles (60°-90° N). Zonal**  
 864 **mean CERES observations are shown as a dark gray line for reference. b, as in a, but for the**  
 865 **surface component of the clear-sky reflection. Large errors for IPSL-CM6A-LR in the Northern**  
 866 **Hemisphere tropics and Southern Hemisphere poles are primarily due to a very high degree of**  
 867 **temporal autocorrelation as opposed to large standard deviations.**  
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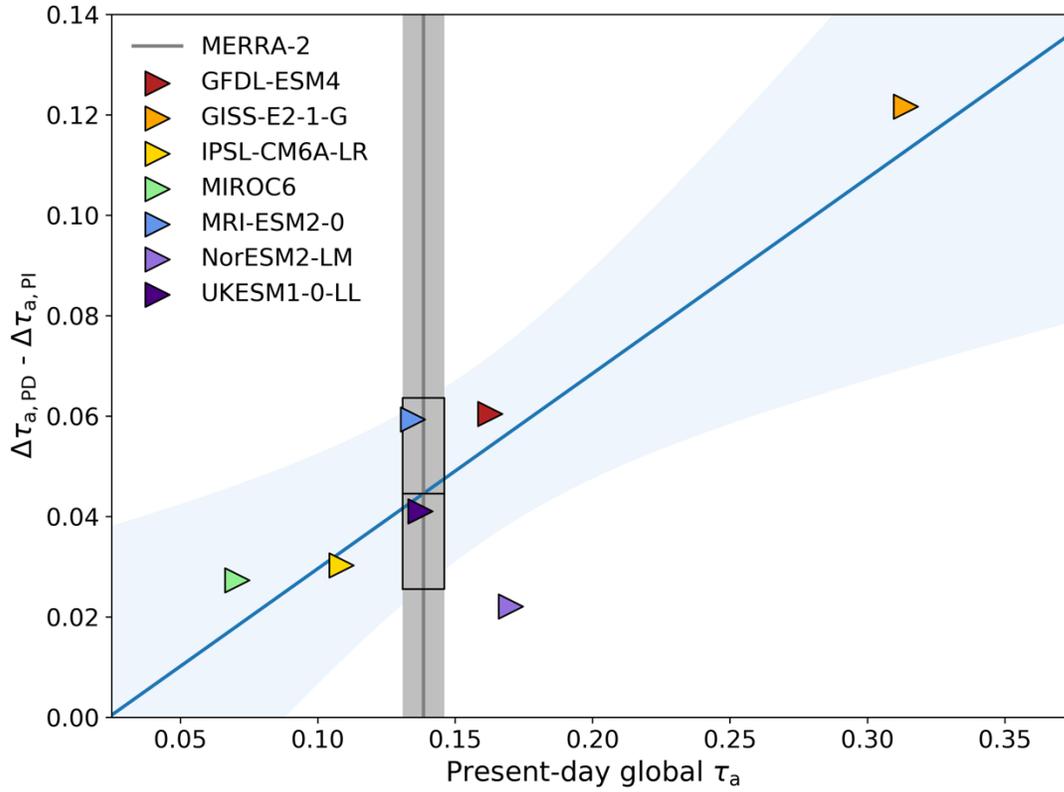


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871 **Extended Data Fig. 5 | Modern and full historical regression slopes for each CMIP6 model.**

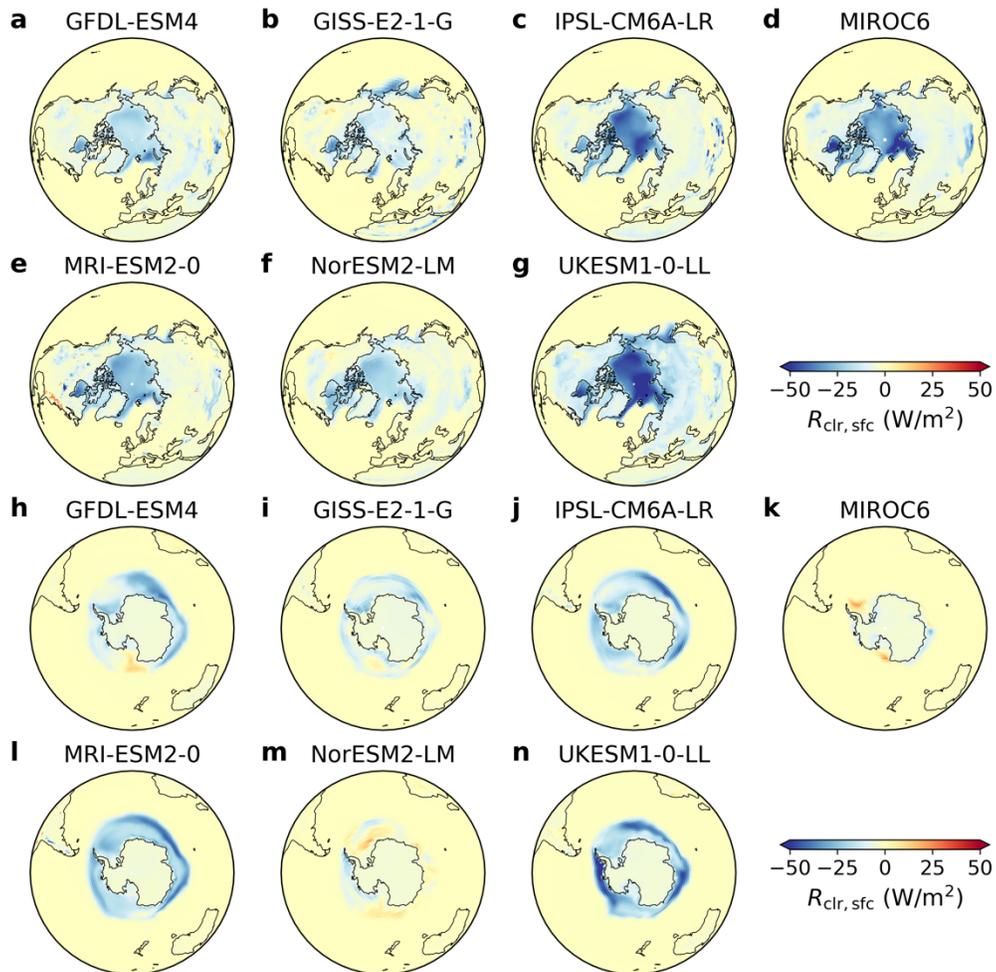
872 Regression slopes ( $\hat{\beta}$ , units of W/m<sup>2</sup> in  $\Delta R_{\text{clr,atm}}$  per unit  $\Delta\tau_a$ ) for each CMIP6 model and their  
873 95% confidence intervals are represented by colored markers (square for the 1850-2015  
874 regression, triangle for 2000-2015 only) and error bars.

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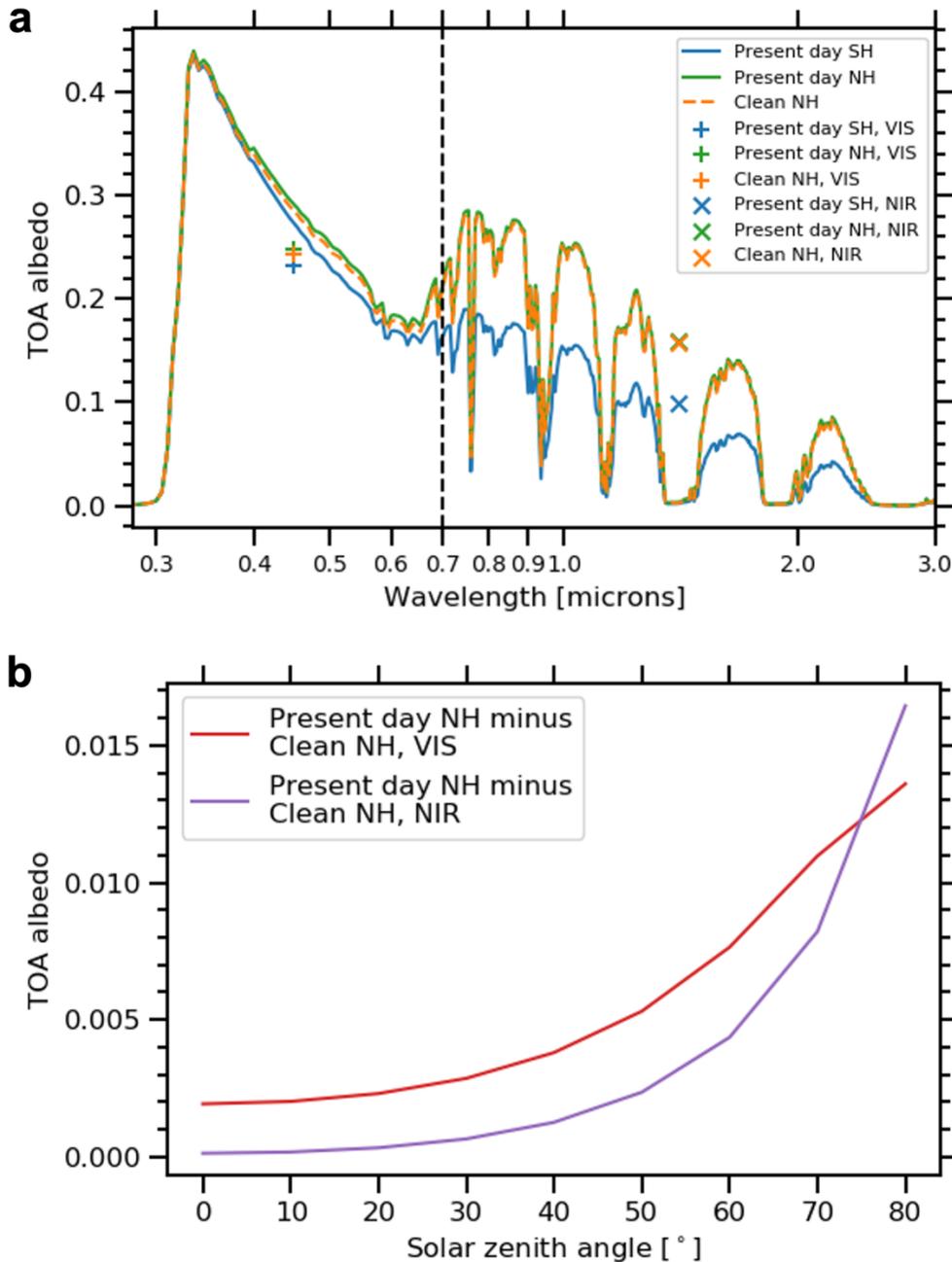


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 877 **Extended Data Fig. 6 | Emergent constraint for the change in hemispheric aerosol contrast**  
 878 **from present-day to pre-industrial based on present-day global mean aerosol optical depth.**  
 879 CMIP6 models are represented by the colored triangles and their regression slope and its 95%  
 880 confidence interval by the blue line and shading. MERRA-2 values for the present-day global  
 881 mean AOD and its 95% confidence interval are represented by the gray line and shading. The  
 882 constraint on the present-day to pre-industrial change in  $\Delta\tau_a$  is represented by the black box, with  
 883 a center line at the mean value and extent based on the 95% confidence interval.  
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SSP3-7.0 2085-2100 minus historical 2000-2015



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 886 **Extended Data Fig. 7 | Change in surface reflection over the poles in the SSP3-7.0 high-**  
 887 **emissions scenario.** Difference in surface reflection between the SSP3-7.0 end-of-century  
 888 (2085-2100 mean) and historical present-day (2000-2015 mean) for each CMIP6 model centered  
 889 around the Arctic (**a-g**) and Antarctic (**h-n**) using an orthographic map projection.  
 890



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 892 **Extended Data Fig. 8 | Radiative transfer calculations of the clear-sky albedo for a**  
 893 **hypothetical "cleaner" Northern Hemisphere.** **a**, spectral top-of-atmosphere albedo at an  
 894 example solar zenith angle of 50° for present-day Northern Hemisphere, present-day Southern  
 895 Hemisphere, and clean Northern Hemisphere. Spectrally-integrated albedos for the ultraviolet  
 896 and visible portion (VIS) of the spectrum (0.2-0.7  $\mu\text{m}$ ) are represented by the colored pluses and  
 897 the near-infrared portion (NIR) of the spectrum (0.7-3.0  $\mu\text{m}$ ) by crosses. **b**, differences between  
 898 present-day and clean NH for VIS and NIR albedo as a function of solar zenith angle.  
 899