

1 **Megathrust Heterogeneity, Crustal Accretion, and a Topographic Embayment in**
2 **Western Nepal Himalaya: Insights from the Inversion of Thermochronological Data**

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17 **Key Points:**

- 18 • Thermochronologic data and thermokinematic modeling results suggest a complex 3D
19 geometry of the megathrust in the western Nepal Himalaya
- 20 • Crustal accretion along the mid-lower crustal ramp in the megathrust is a mechanism of
21 maintaining the edge of a plateau-like landscape
- 22 • The development of the 3D megathrust geometry may have controlled the first-order
23 evolution of an embayment and regional drainage system

35 Abstract

36 Between 81°30' E and 83°E the Himalayan range's "perfect" arcuate shape is interrupted by an
37 embayment. We hypothesize that thrust geometry and duplexing along the megathrust at mid-
38 lower crustal depths plays a leading role in growth of the embayment as well the southern margin
39 of the Tibetan plateau. To test this hypothesis, we conducted thermokinematic modeling of
40 published thermochronologic data from the topographic and structural embayment in the western
41 Nepal Himalaya to investigate the three-dimensional geometry and kinematics of the megathrust
42 at mid-lower crustal depths. Models that can best reproduce observed cooling ages suggest that
43 the megathrust in the western Nepal Himalaya is best described as two ramps connected by a
44 long flat that extends further north than in segments to the east and west. These models suggest
45 that the high-slope zone along the embayment lies above the foreland limb of an antiformal
46 crustal accretion zone on the megathrust with lateral and oblique ramps at mid-lower crustal
47 depths. The lateral and oblique ramps may have initiated by ca. 10 Ma. This process may have
48 controlled along-strike variation in Himalayan-plateau growth and therefore development of the
49 topographic embayment. Finally, we analyze geological and morphologic features and propose
50 an evolution model in which landscape and drainage systems across the central-western
51 Himalaya evolve in response to crustal accretion at depth and the three-dimensional geometry of
52 the megathrust. Our work highlights the importance of crustal accretion at different depths in
53 orogenic-wedge growth and that the mid-lower crustal accretion determines the location of
54 plateau edge.

55 1 Introduction

56 For large orogens, the roles of the geometry and kinematics of the megathrust in
57 controlling strain accumulation and the feedback mechanisms between deformation and erosion
58 are keys to understanding the growth of the orogenic wedge and the evolution of its morphology.
59 The Himalayan orogenic wedge is a classical natural laboratory to study this, because it exhibits
60 both along-strike consistency in first-order tectonic-morphologic framework and pronounced
61 along-strike segmentation. Early studies show that the tectonostratigraphy, major shear zones,
62 deformation styles, and morphologic features are generally continuous along the strike of the >
63 1500 km orogen (Heim and Gansser, 1939; Gansser, 1964; Le Fort, 1975; Burg and Chen, 1984;
64 Pêcher, 1989; Yin, 2006) (Fig. 1). Along with these features, concentrated micro-seismicity,
65 maximum horizontal strain rate, zones of high river gradient, and zones of high relief are
66 interpreted to be collocated along a small circle or "perfect arc" (Seeber and Gornitz, 1983;
67 Bendick and Bilham, 2001). This arc is interpreted to represent the active-uplift front of the high
68 Himalaya based on geological, geophysical, and geomorphological observations and to mark the
69 plateau margin, the transition between a taper-shaped outer wedge and inner-wedge plateau (Fan
70 and Murphy, 2021 and references therein). It is thought to form through deformation associated
71 with a mid-lower crustal ramp in the megathrust, which connects an upper-crustal flat below the
72 outer wedge and a flat seismic reflector extending beneath Tibet (Hauck et al., 1998; Nábělek et
73 al., 2009).

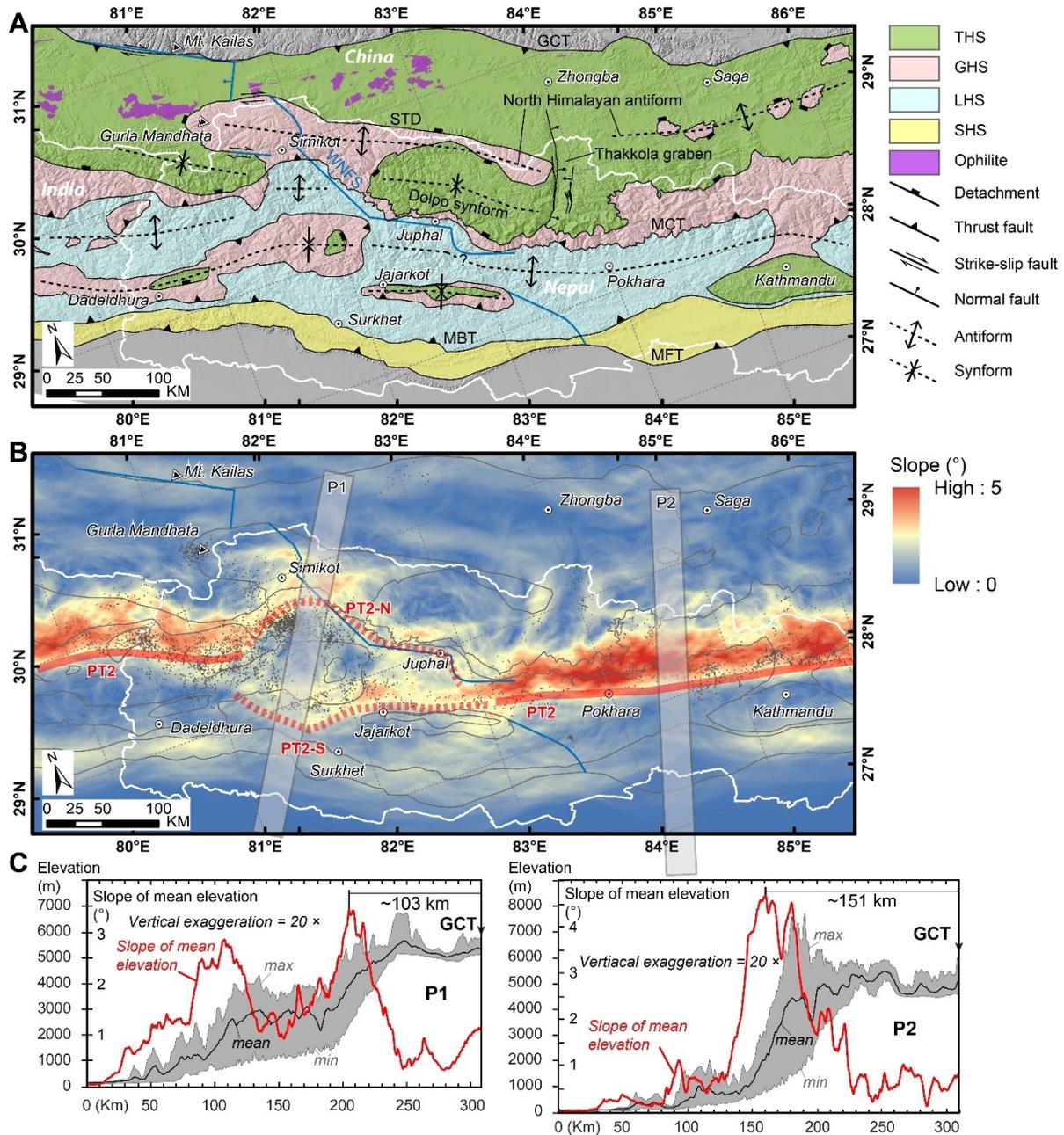


Figure 1 Geological map (A) and slope of mean elevation map (B) of central-western Nepal and adjacent areas and (C) topographic swath profiles and slope of mean elevation profiles of western and central Nepal Himalaya. Grey dots in (B) show earthquake hypocenters from 1995 to 2004, relocated by Ader et al. (2012). Gray lines in (B) are major faults shown in (A). Slope of mean elevation was calculated by first smoothing the topography by taking the mean within a 25 km moving window and then calculating the slope of the resulting grid. THS – Tethyan Himalayan Sequence; GHS – Greater Himalayan Sequence; LHS – Lesser Himalayan Sequence; SHS – Sub-Himalayan Sequence; GCT – Great Counter Thrust; STD – South Tibet Detachment; MCT – Main Central Thrust; MBT – Main Boundary Thrust; MFT – Main Frontal Thrust; WNFS – Western Nepal Fault System.

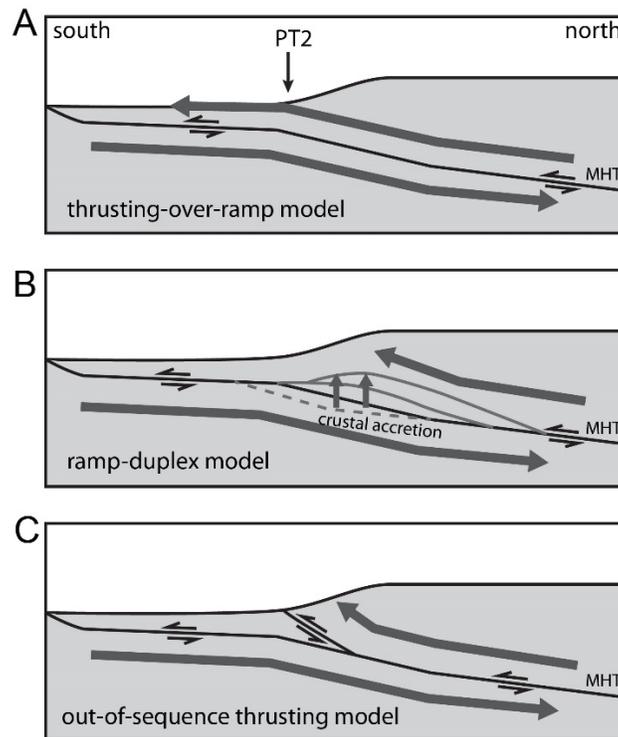
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76 Megathrust kinematic models can be classified into three main groups (Fig. 2); (1)
77 thrusting of the orogenic wedge over a crustal ramp (Gansser, 1964; Jackson and Bilham, 1994;
78 Cattin and Avouac, 2000; Lavé and Avouac, 2001; Robert et al., 2009; Robert et al., 2011;
79 Coutand et al., 2014; Elliott et al., 2016), (2) thrusting of the orogenic wedge over a crustal ramp
80 accompanied by localized transfer of material from the footwall to hanging wall via duplexing or
81 accretion (Avouac, 2003; Bollinger et al., 2004; Bollinger et al., 2006; Herman et al., 2010;
82 Grandin et al., 2012; Cannon and Murphy, 2014; Adams et al., 2016; Landry et al., 2016;
83 Stübner et al., 2018; Fan and Murphy, 2021), and (3) out-of-sequence thrusting (Harrison et al.,
84 1997; Wobus et al., 2003; Hodges et al., 2004; Thiede et al., 2004; Thiede et al., 2005; Wobus et
85 al., 2005; Wobus et al., 2006; Whipple et al., 2016) or a hybrid of out-of-sequence and duplex
86 model (Morell et al., 2017). Besides the dispute over the kinematics controlling the active-uplift
87 zone, an increasing number of studies challenge the notion of a “perfect arc” orogenic system
88 with evidences of non-cylindricity in surface geology, wedge morphology, and the deep
89 megathrust geometry (Duncan et al., 2003; Yin, 2006; Hetenyi et al., 2016; Cannon et al., 2018;
90 Eugster et al., 2018; Dal Zilio et al., 2020a; Fan and Murphy, 2021). One pronounced deviation
91 from the “perfect arc” shape of the Himalaya is the bifurcation of several features defining the
92 active-uplift front in western Nepal Himalaya (Fig. 1) (Harvey et al., 2015). The northern branch
93 of the bifurcation has been interpreted to represent the active-uplift zone, implying that there is a
94 recession or embayment in the active-uplift front of the wedge (Cannon et al., 2018; Fan and
95 Murphy, 2021). Some studies have ascribed this along-strike anomaly to an along-strike
96 difference in the number and location of ramps in the Himalayan megathrust, the Main
97 Himalayan Thrust (MHT) (Robert et al., 2011; Harvey et al., 2015; van der Beek et al., 2016;
98 Fan and Murphy, 2021). This conceptual interpretation is consistent with the observation of a
99 longer upper-crustal flat in the megathrust in western Nepal than in some other sectors (Subedi et
100 al., 2018) and the pattern of microseismicity (Ader et al., 2012; Hoste-Colomer et al., 2018).
101 However, different from the well-studied MHT upper-crustal flat and the structures above it in
102 the outer wedge (DeCelles et al., 2020 and references therein), the geometry and kinematics of
103 the MHT at aseismic slip depths in western Nepal are not well understood.

104 A better understanding of the geometry of the megathrust in western Nepal Himalaya is
105 also crucial to assess seismic hazard in the region. Beneath a thickened continental crust, the
106 downdip limit of the seismogenic zone along a megathrust should be located where the
107 megathrust intersects the temperature-controlled transition between seismic and aseismic slip
108 depths (Hyndman et al., 1997). The geometry of the megathrust therefore should be the
109 determining factor for the along-strike variation in the location of the downdip limit of the
110 seismogenic zone. The along-strike complexity of this limit in the Himalaya is implied by the
111 heterogeneity shown in coupling models derived from geodetic data (Ader et al., 2012; Stevens
112 and Avouac, 2015; Marechal et al., 2016; Jouanne et al., 2017; Li et al., 2018; Sreejith et al.,
113 2018; Yadav et al., 2019; Dal Zilio et al., 2020a). The potential along-strike changes in the
114 geometry of the MHT require the existence of lateral or oblique ramps. These ramps, including
115 frontal ramps at the front of the wedge could act as barriers to rupture propagation during large
116 earthquakes as suggested by studies on the 2015 Mw 7.8 Gorkha earthquake in central Nepal
117 (Duputel et al., 2016; Elliott et al., 2016; Hubbard et al., 2016; Kumar et al., 2017; Wang et al.,
118 2017; Zhang et al., 2017; Bai et al., 2019; Wang et al., 2019). In western Nepal Himalaya, the
119 along-strike extent of the embayment approximately corresponds to the region devastated in the
120 1505 AD earthquake (Bollinger et al., 2016), and other large historical earthquakes have been
121 reported (Murphy et al., 2014; Hossler et al., 2016; Ghazoui et al., 2019). These observations

122 suggest significant seismic hazard for western Nepal and that the megathrust lateral ramps may
 123 control rupture propagation.



124

Figure 2 Simplified sketch of the kinematic models of MHT explains the formation of physiographic transition (PT2) between outer wedge and inner wedge. (A) Overthrusting and underthrusting along the MHT ramp. (B) Crustal accretion in the form of duplexes along MHT ramp. (C) Out-of-sequence thrusting at the physiographic transition.

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126 To address these issues, we investigate the geometry and kinematics of the megathrust
 127 from the perspective of thermokinematics. In this contribution we report new zircon (U-Th)/He
 128 ages. Most of our samples are located in the eastern limb of the embayment. The rest of the
 129 samples are located in the hinterland of the Himalayan plateau. Previously published
 130 thermochronologic data are used to inform inversion models with different megathrust
 131 kinematics. We also take this opportunity to discuss the general evolution of crustal ramps and
 132 topography in western Nepal by comparing our study with observations in adjacent areas. We
 133 conclude by summarizing sedimentary studies across the Himalaya orogen and propose a
 134 conceptual model for the landscape and drainage system evolution that highlights the role of
 135 deep crustal accretion in the development of the Himalayan orogenic plateau.

136 2 Geological Setting

137 2.1 Tectonostratigraphy and Structures

138 The Himalayan orogen is characterized by several rock sequences juxtaposed by major
 139 east-west-striking shear zones (Gansser, 1964; Le Fort, 1975; Yin, 2006). From structurally high
 140 to low levels in the thrust wedge, these rock sequences include the Tethyan Himalayan Sequence

141 (THS), the Greater Himalayan Sequence (GHS), the Lesser Himalayan Sequence (LHS), and the
142 Siwalik Group (SG) (Fig. 1). The THS is a deformed package of unmetamorphosed to low-grade
143 metasedimentary rocks that experienced greenschist to amphibolite facies metamorphism
144 (Gaetani and Garzanti, 1991; Burchfiel et al., 1992; Liu and Einsele, 1994; Murphy and Yin,
145 2003; Myrow et al., 2009). The THS primarily crops out from the Indus-Yarlung suture zone to
146 the South Tibet Detachment (STD), a north-directed normal-sense shear zone or passive roof
147 thrust that juxtaposes the THS against the GHS (Burg et al., 1984; Burchfiel et al., 1992; Hodges
148 et al., 1992; Coleman, 1996; Edwards et al., 1996; Carosi et al., 1998). The GHS is composed of
149 middle amphibolite facies to lower granulite facies meta-sedimentary and meta-igneous rocks
150 (Pêcher, 1989; Vannay and Hodges, 1996; Murphy et al., 2002; Searle and Godin, 2003; Martin
151 et al., 2009; Kohn, 2014; Iaccarino et al., 2017). It is juxtaposed against the LHS along the Main
152 Central Thrust (MCT), a top-to-south shear sense thrust fault (Burg and Chen, 1984; Brunel,
153 1986; Schelling, 1992; Robinson et al., 2003). The MCT is primarily active during the Late
154 Oligocene - Middle Miocene (e.g. Hubbard and Harrison, 1989; Kohn et al., 2005; Yin, 2006;
155 Montomoli et al., 2013; Cottle et al., 2015a; Larson et al., 2015; Carosi et al., 2018; Catlos et al.,
156 2018; Braden et al., 2020). Some segments of the MCT are known to have been reactivated in
157 the Late Miocene - Pliocene (e.g. Catlos et al., 2001; Braden et al., 2018; Catlos et al., 2018;
158 Montemagni et al., 2019; Braden et al., 2020). The STD is usually interpreted to be coeval with
159 the main active stage of the MCT and ceased moving in the Early Miocene in the western Nepal
160 Himalaya (Hodges et al., 1992; Hodges et al., 1998; Murphy and Harrison, 1999; Searle and
161 Godin, 2003; Godin et al., 2006; Cottle et al., 2015b). Between the STD and MCT, within the
162 GHS, some studies report tectonic or metamorphic discontinuities, which have been interpreted
163 to accommodate in-sequence thrusting prior to the initiation of the MCT (e.g. Larson et al., 2015;
164 Montomoli et al., 2015; Carosi et al., 2018). Toward the end of the time that the STD-MCT
165 system was active (ca. 19–13 Ma), a series of mid-crustal culminations referred to as the North
166 Himalayan antiform developed in the central Himalayan hinterland and expose the high-grade
167 orogenic core (e.g. Lee et al., 2000; Godin et al., 2006; Lee and Whitehouse, 2007; Larson et al.,
168 2010) (Fig. 1). In the western Nepal Himalaya, the Northern Himalayan antiform is characterized
169 by an east-west-trending belt of dome-shaped outcrops of GHS rocks from 81°E to 90°E (Fig. 1).
170 The Thakkhola graben is bounded by several steeply dipping north-south striking normal faults
171 (Hurtado et al., 2001; Baltz, 2012), and the Gurla Mandhata area is a metamorphic core complex
172 bounded by the Gurla Mandhata-Humla fault system (Murphy et al., 2002; Murphy and
173 Copeland, 2005). They both developed mainly starting from middle-late Miocene and
174 accommodated orogen-parallel extension (Coleman and Hodges, 1995; Murphy et al., 2002;
175 Murphy and Copeland, 2005; McCallister et al., 2014; Brubacher et al., 2020). Fan and Murphy
176 (2021) reconciled the coeval orogen-normal shortening and orogen-parallel extension and the
177 difference in the amount of extension of these two features in a 3D conceptual model of oblique
178 convergence in an arcuate orogen. The LHS consists of lower greenschist- to lower-amphibolite-
179 facies metasedimentary rocks and is deformed by a duplex system that formed in the Late
180 Miocene after ca. 10 Ma within the outer wedge (e.g. Srivastava and Mitra, 1994; DeCelles et al.,
181 2001; Robinson et al., 2006; Webb, 2013). Growth of the duplex led to folding of the MCT
182 thrust sheet structurally above, which is presently exposed in synformal klippen in the outer
183 wedge (DeCelles et al., 2001; Pearson and DeCelles, 2005). The antiformal LHS duplex and
184 Northern Himalayan antiform usually make the region between them form synformal half-klippe,
185 for example the Dolpo THS synform (Fig. 1). The Main Boundary Thrust (MBT) juxtaposes the
186 LHS against the SG, which are foreland basin deposits of the wedge since middle Miocene and

187 then incorporated into the thrust wedge (Mugnier et al., 1999; Bernet et al., 2006; van der Beek
188 et al., 2006; Baral et al., 2015). It is separated from the undeformed foreland by the Main Frontal
189 Thrust (MFT). The MFT is the surface trace of the MHT. The MBT and MCT sole into the MHT
190 at depth. In central-western Himalaya, the surface geology suggests the existence of MHT lateral
191 or oblique ramps that affect foreland-ward propagation. Between the Thakkhola graben and the
192 Gurla Mandhata area, the trace of the MCT, LHS duplex, and the axis of the folded Almora-
193 Dadeldhura klippe, all shift toward the hinterland (north) in western Nepal. This results in the
194 alignment of antiformal features, such as LHS duplexes, and synformal features, such as GHS
195 klippen and THS half-klippen, along their strike (DeCelles et al., 2020; Fan and Murphy, 2021)
196 (Fig. 1).

197 2.2 Topography and Landscape

198 The topography of the Himalayan wedge is characterized by a sharp physiographic
199 transition characterized by an abrupt transition from high slopes between the high-elevation
200 hinterland plateau and the lower slopes that mark the low-elevation frontal wedge (PT2 of
201 Hodges et al., 2001). It is located at the base of a narrow high-slope zone connecting the inner
202 wedge characterized by high-elevation, low-relief plateau landscape, and the outer wedge
203 characterized by a taper-shaped regional slope (Fig. 1). PT2 exhibits along-strike variations and
204 is usually collocated with features that define the active-uplift front of the Himalaya (e.g. Duncan
205 et al., 2003; Harvey et al., 2015; Morell et al., 2015). The young cooling ages along this zone
206 suggest rapid exhumation (e.g. Copeland et al., 1991; Wobus et al., 2003; Blythe et al., 2007;
207 Johnston et al., 2020), which is thought to facilitate extrusion of high-grade rocks from lower
208 crust in some models (e.g. Beaumont et al., 2001; Hodges et al., 2001). The above-mentioned
209 bifurcation of the high slope zone in western Nepal Himalaya obscure the plateau-taper pattern
210 of the topography there (Fig. 1). The northern and south branches of PT2 in western Nepal are
211 referred to as PT2-N and PT2-S (Harvey et al., 2015).

212 2.3 Megathrust Models

213 Although some MHT models explain the specific location of the active-uplift front of the
214 high Himalaya and supporting datasets in other areas in Himalaya, no such model exists for
215 western Nepal. Several conceptual models are proposed suggesting an upper-crustal flat extends
216 further to the north compared along-strike were proposed based on low-temperature
217 thermochronologic ages and geomorphology (Harvey et al., 2015; van der Beek et al., 2016).
218 Based on the better understanding of the MHT in the area ruptured by the Gorkha earthquake,
219 Hubbard et al. (2016) proposed a three-dimensional model for the MHT across Nepal. Similarly,
220 by integrating various datasets and a thickness model of the Himalayan high grade core, Fan and
221 Murphy (2021) introduced lateral and oblique ramps to the 2D conceptual model of the MHT in
222 western Nepal and proposed a 3D conceptual model explaining the embayment of the active-
223 uplift front and the along-strike change in the surface geology. Several studies focusing on the
224 fold and thrust belt in the outer wedge of the Himalaya propose MHT models based on a
225 structural restoration (DeCelles and DeCelles, 2001; Robinson, 2008; Robinson and McQuarrie,
226 2012; Olsen et al., 2019; DeCelles et al., 2020) (Fig. 3). These models can well explain the
227 detailed structures in the outer wedge to the south of the PT2 but cannot resolve the structures in
228 the inner wedge and the mid-lower crustal ramp in the MHT. They also fail to explain the first-
229 order landscape because the northern MHT ramp in these models is usually located below the

230 area between the PT2-S and PT2-N, where the landscape is relatively flat and of low relief (Fig.
 231 3).

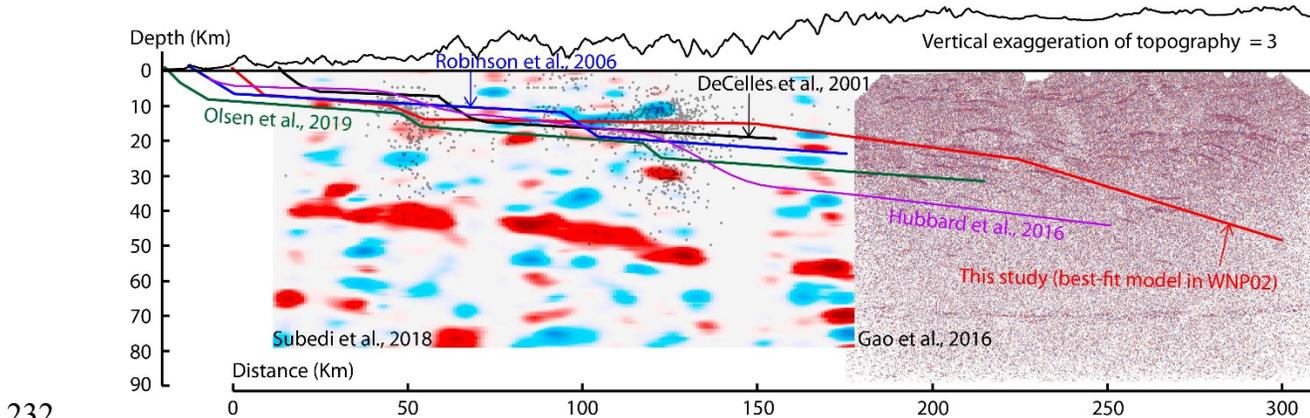


Figure 3 A comparison between different models of the MHT in the western Nepal Himalaya. The red line shows the geometry of preferred model in this study. Gray dots show the earthquake hypocenters projected to the profile from within 20 km. The data includes earthquakes from December 2014 to November 2015 reported by Hoste-Colomer et al. (2018) and earthquakes from 1995 to 1999 recorded by National Seismic Center and relocated by Ader et al. (2012). Image in the left is P-to-S receiver function migration image reported by Subedi et al. (2018) in which the red and blue represent high and low PS/P amplitude respectively. The seismic reflection profile image in the right is reported by Gao et al. (2016). For original interpretation of the images, refer to the original papers. Notice the reflector along the upper crustal flat of the MHT and the duplexes in the hinterland.

233
 234 Receiver function analysis of teleseismic waveforms in western Nepal suggests a longer
 235 upper-crustal flat in MHT compared to central Nepal. This flat in the MHT connects to a long,
 236 gently dipping ramp that extends northward to lower-crustal depths (Subedi et al., 2018) (Fig. 3).
 237 However, using the available data, the geometry of this ramp is not precisely located. Subedi et
 238 al. (2018) interpret the ramp geometry based on the change in depth between the upper-crustal
 239 reflectors and several short reflectors at the northern end of the profile. The reflectors at the
 240 northern end are consistent with lower-crustal reflectors on the other receiver function profile
 241 image that extends from to the west of the Gurla Mandhata metamorphic core complex to the
 242 south of South Tibet (Xu et al., 2017). A deep seismic reflection profile along a transect close to
 243 the profile in Xu et al. (2017) also shows a middle-lower crustal duplex structurally above the
 244 mid-lower crustal reflector (Gao et al., 2016). The locations of foci of earthquakes in western
 245 Nepal also suggest the seismogenic part of the MHT extends further to the north compared to
 246 areas along strike (Ader et al., 2012; Hoste-Colomer et al., 2018). However, none of these
 247 models are based on a geophysical data along a complete transect from the lower Himalaya to
 248 the higher Himalaya of the embayment in western Nepal area and none have been tested with
 249 geologic data.

250 3 Thermochronologic Data

251 3.1 Sample and Methods

252 In order to investigate the geometry and kinematics of the MHT in western Nepal and the
 253 correlation between 3D geometry of the MHT (e.g. lateral and oblique ramp locations) with the

254 embayment, we carried out thermokinematic modeling along a transect that extends across the
255 apex of the embayment. We then compare the modeling results with our new thermochronologic
256 data along the eastern limb of the embayment. The data used in the modeling is from published
257 studies and is described later.

258 Zircon (U-Th)/He (ZHe) dating was carried out because of the sensitivity of ZHe ages to
259 active and new tectonics over a time span of several million years. The closure temperature of
260 the ZHe thermochronometer is affected by several factors, such as chemical composition of the
261 crystals, concentration of radiation damage, grain size, and cooling rate (Reiners et al., 2004;
262 Reiners, 2005). For typical plutonic cooling rates and crystal sizes, the closure temperature is
263 about 170-190 °C (Reiners et al., 2004). Therefore, it can record the exhumation history through
264 the upper about 5-8 km of the crust. In an active region, such as the Himalaya, where the highest
265 exhumation rate close to the PT2 can reach several mm/yr, the ZHe ages can record cooling
266 history in the past less than 3 million years.

267 We analyzed 12 samples. Seven are from the northern branch of PT2 (Fig. 4). These data
268 fill a data gap between far-western Nepal along the Karnali river transect and western Nepal. The
269 other samples are from the inner wedge. The locations, lithology, and units of the samples are
270 listed in Table S1.

271 Samples were processed for ZHe dating at thermochronology lab in University of Texas,
272 Austin (refer to Text S1 for analytical procedure). For each sample, we usually analyzed three
273 grains, but from DG-10 and DF-3 we analyzed four grains and two grains respectively. Some
274 aliquots were discarded on the basis of anomalous isotopic values and others because they
275 yielded outlier ages defined as being more than 30% older or younger than the rest of the grains
276 in their respective samples.

277 3.2 Results

278 Analytic results of our samples are reported in Table 1. Samples from south of the Dolpo
279 synform have ZHe ages that range from 3.1 – 7.4 Ma. Two samples among them are 3.1 and 3.3
280 Ma. These young ages are consistent with or slightly older than the ages reported at the active-
281 uplift front, though they are closer to the front of the orogen compared with the young cooling
282 ages reported at the apex of the topographic embayment in western Nepal Himalaya (Fig. 4). The
283 samples from the inner wedge have ZHe ages that range from 9.5 – 12.2 Ma, significantly older
284 than the samples from the southern Dolpo area and from the apex of the topographic embayment,
285 though they approximately align well with the samples of young cooling ages at the apex of the
286 topographic embayment along the strike of the northern Himalayan anticline (Fig. 4). These old
287 ages are consistent with recently reported ages close to the northern Himalayan antiform in the
288 inner wedge along the Thakkhola graben (Brubacher et al., 2020).

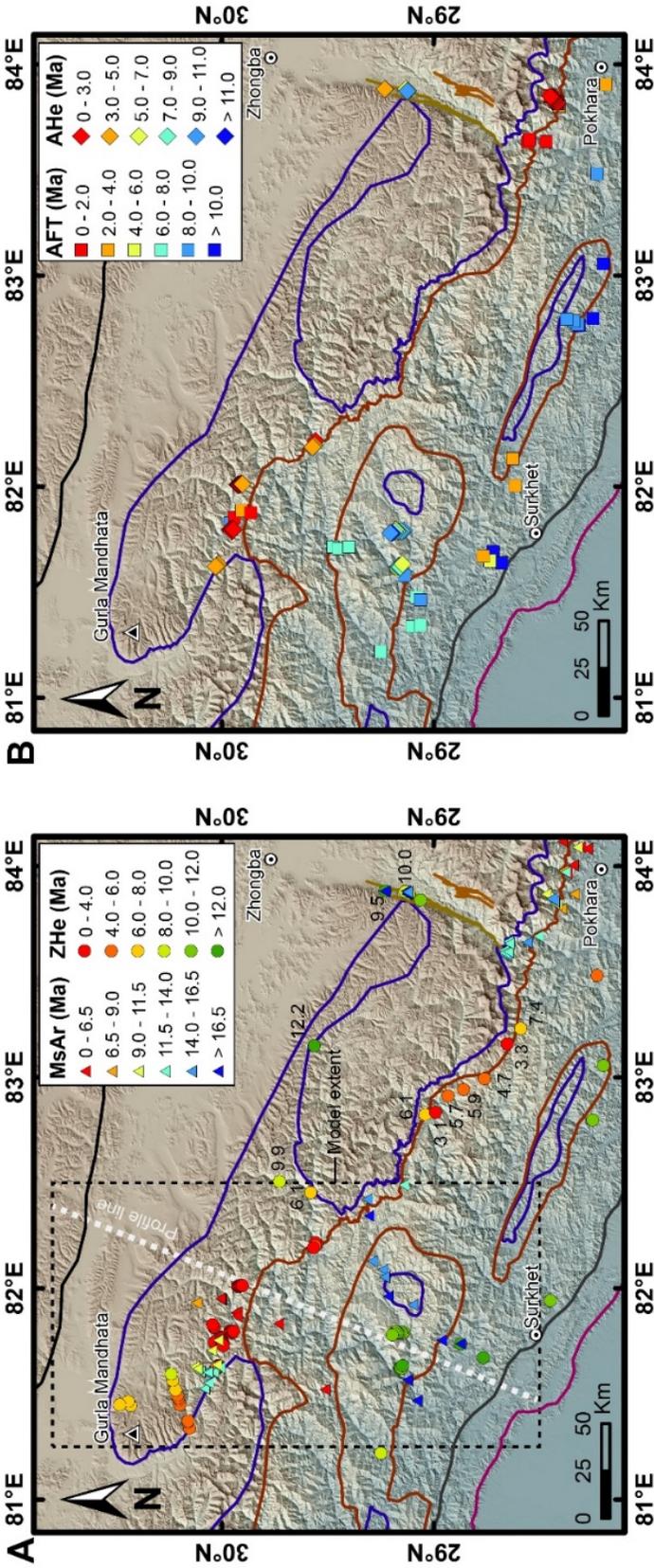


Figure 4 Previously reported and new thermochronologic ages in the western Nepal Himalayan area. The dash box in (A) shows the extent of the thermochronologic models in this study. The ZHe ages labeled with age values are newly reported ages in this study. Previously published data are from Vannay and Hodges (1996), Godin et al. (2001), Robert et al. (2011), Nadin and Martin (2012), Sakai et al. (2013), Mercier (2014), McCallister et al. (2014), Harvey (2015), Martin et al. (2015), Nagy et al. (2015), Gibson et al. (2016), van der Beek et al. (2016), Soucy La Roche et al. (2018), Braden et al. (2020), Brubacher et al. (2020), and DeCelles et al. (2020). MsAr – Muscovite ⁴⁰Ar/³⁹Ar age; ZHe – Zircon (U-Th)/He age; AFT – Apatite fission track age; AHe – Apatite (U-Th)/He age.

290

Table 1 Zircon (U-Th)/He Dating Results^a

Sample	Age (Ma)	U (ppm)	Th (ppm)	¹⁴⁷ Sm (ppm)	He (nmol/g)	mass (ug)	Ft	ESR	Mean Age (Ma)	Error (1σ)
DG-10-2	3.3	505.5	28.1	0.0	6.3	1.99	0.70	36.31		
DG-10-3	3.3	309.0	9.2	-3.0	4.1	3.18	0.74	42.97		
DG-10-4	3.4	703.7	39.8	0.8	10.7	11.07	0.82	65.07		
DG-10-5	3.1	551.7	10.1	-1.7	7.2	5.67	0.78	52.54	3.27	0.06
DG-38-2	2.9	124.9	83.6	0.0	1.6	3.16	0.72	41.28		
DG-38-3	3.2	240.7	105.4	0.0	3.1	2.04	0.68	35.25		
DG-38-4	3.1	673.9	182.9	0.0	8.3	1.90	0.69	36.24	3.08	0.07
DH-10-1	5.7	481.3	203.7	2.6	13.4	14.34	0.83	69.80		
DH-10-2	5.5	300.0	75.5	0.0	6.9	3.83	0.74	43.37		
DH-10-3	7.3	518.3	85.7	1.4	17.5	13.10	0.83	68.71	6.14	0.46
<i>Dolpo-04-1</i>	<i>16.3</i>	<i>253.9</i>	<i>87.0</i>	<i>0.8</i>	<i>19.7</i>	<i>11.78</i>	<i>0.82</i>	<i>66.09</i>		
Dolpo-04-2	12.1	278.5	98.3	0.0	15.2	6.24	0.77	50.57		
Dolpo-04-3	12.2	171.6	61.7	0.8	10.0	12.01	0.82	64.83	12.20	0.03
DF-3-2	10.0	1141.8	194.6	0.0	46.4	2.91	0.72	40.72		
<i>DF-3-3</i>	<i>4.7</i>	<i>190.4</i>	<i>59.5</i>	<i>1.8</i>	<i>3.9</i>	<i>5.13</i>	<i>0.76</i>	<i>49.33</i>	10.0	0.80
DF-4-3	10.5	1649.4	30.2	1.3	74.1	7.23	0.79	54.31		
DF-4-1	9.7	1720.5	55.4	1.3	71.2	7.23	0.78	52.79		
DF-4-2	8.4	1551.4	27.8	0.0	41.8	0.86	0.60	26.03	9.53	0.51
DG-12-1	8.0	689.6	81.1	0.0	21.4	2.64	0.70	37.73		
<i>DG-12-2</i>	<i>14.1</i>	<i>1937.3</i>	<i>46.7</i>	<i>0.0</i>	<i>93.0</i>	<i>1.37</i>	<i>0.63</i>	<i>28.95</i>		
DG-12-3	6.8	342.1	54.0	0.0	7.8	0.94	0.60	26.66	7.40	0.41
DG-22-1	6.2	321.2	88.3	2.2	9.5	16.75	0.84	72.47		
DG-22-2	4.6	439.6	34.1	0.0	8.3	4.78	0.76	46.31		
DG-22-3	3.3	262.8	37.6	0.0	3.4	2.30	0.70	37.83	4.68	0.67
DG-29-1	6.7	1143.0	35.4	0.0	29.8	2.65	0.71	38.82		
DG-29-2	5.2	136.1	74.8	1.9	3.6	15.10	0.83	72.20		
DG-29-3	5.8	514.7	83.8	0.0	13.1	6.78	0.79	55.59	5.91	0.36
DG-30-1	5.5	281.9	30.1	1.5	7.5	30.41	0.87	91.30		
DG-30-2	6.4	301.1	44.1	1.2	9.8	136.73	0.92	151.44		
DG-30-3	5.2	166.8	41.4	1.5	4.2	18.85	0.84	75.56	5.69	0.29
TB10-12-1	5.4	294.9	97.9	1.9	7.5	14.49	0.82	66.64		
TB10-12-2	6.8	322.2	75.5	0.0	10.4	17.24	0.84	73.28		
<i>TB10-12-3</i>	<i>16.5</i>	<i>248.9</i>	<i>97.5</i>	<i>1.0</i>	<i>19.3</i>	<i>8.97</i>	<i>0.80</i>	<i>57.83</i>	6.06	0.50
TB10-9-1	11.7	1054.0	90.3	21.2	57.8	24.09	0.85	80.46		
TB10-9-2	10.2	794.7	106.4	10.7	39.7	45.06	0.88	99.95		
TB10-9-3	7.7	542.8	86.0	5.1	20.2	31.07	0.86	86.76	9.87	0.95

^aFt is the alpha ejection correction factor. ESR is equivalent sphere radius. Data in italic indicate outliers that were discarded in mean age calculation. The age error is the standard deviation between aliquots divided by the square root of the number of aliquots.

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292 4 Thermokinematic Modeling

293 To test the geometric and kinematic models of the MHT, we conducted thermokinematic
 294 modeling by using a modified version of the software Pecube (Braun, 2003; Braun et al., 2012).
 295 Models are evaluated by comparing the observed ages obtained from sample analyses with ages
 296 predicted by the software. To quantitatively do this, the software calculates the misfit using a
 297 goodness-of-fit statistic Φ ,

298

$$\Phi = \frac{1}{n} \sqrt{\sum_{i=1}^n \frac{(PredAge_i - ObsAge_i)^2}{\sigma_i^2}}$$

299 where n is the number of thermochronometer ages, *PredAge* and *ObsAge* are the ages
 300 predicted by Pecube and obtained from sample analyses, respectively, and σ is the 1-sigma
 301 uncertainty in the observed age.

302 4.1 Modeling Extent and Inversion Input Data

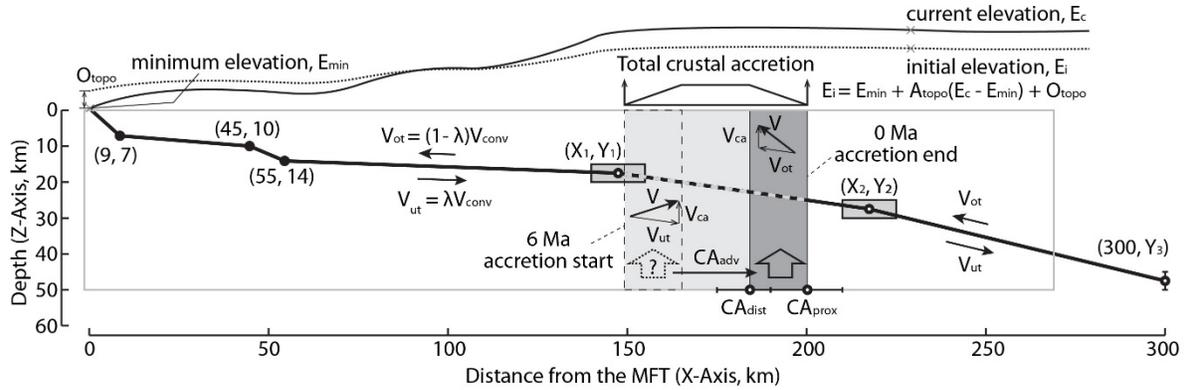
303 Conceptual models describe a network of oblique ramps to explain along strike changes
 304 in the geometry of the MHT (Hubbard et al., 2016; Fan and Murphy, 2021). In this study, we
 305 focus on explaining the most representative sector of the embayment along the Karnali river
 306 transect, where it reaches its largest recession (Fig. 4). Because faults in Pecube can only be
 307 defined by a constant strike (i.e., no lateral ramp can be incorporated) we use the MHT of the
 308 Karnali transect to define the megathrust of the entire model and only use the ages that are
 309 representative of the Karnali transect in the model.

310 Thermochronologic ages used in the modeling are reported from multiple dating systems
 311 including muscovite $^{40}\text{Ar}/^{39}\text{Ar}$ (MsAr) ages, zircon and apatite (U-Th)/He (ZHe and AHe
 312 respectively) ages, and apatite fission track (AFT) ages (McCallister et al., 2014; Mercier, 2014;
 313 Harvey, 2015; Nagy et al., 2015; van der Beek et al., 2016; Soucy La Roche et al., 2018; Braden
 314 et al., 2020). We selected the data for modeling from published datasets following criteria as
 315 follows: (1) we do not use the data out of the swath of our interested section because those
 316 samples might have significantly different thermal histories from the rocks within the swath due
 317 to the existence of lateral and oblique ramps of the MHT; (2) we do not use data that may be
 318 affected by the upper-crustal ramp of the MHT in the outer wedge close to the PT2-S because
 319 this is outside the area of interest; (3) we do not use ages older than 16 Ma because several
 320 studies suggest that the dynamics of the Himalayan orogen and thus the exhumation pattern
 321 changed significantly during the Middle Miocene and resolving the complex evolution history in
 322 this transition is not a goal of this study (e.g. Mugnier and Huyghe, 2006; DeCelles et al., 2011;
 323 Husson et al., 2014; Webb et al., 2017). For some data close to the interpreted lateral ramps, we
 324 do not use them in the inversion modeling or calculating the misfit values, but we still use the
 325 algorithm to estimate their ages in the forward model and incorporate them in the age-
 326 comparison plot for discussion. Close to the PT2-N, both ca. 6 Ma and 9-10 Ma MsAr ages are
 327 reported along Karnali transect, but MsAr ages reported along the active-uplift front in other
 328 sectors in Himalaya are mostly ca. 6 Ma or younger. The reason for the large difference in the
 329 MsAr cooling ages over a short distance along the modeled transect is unknown, therefore we
 330 used all the ages. The distribution of the ages used in the modeling is shown in the **Figure S1**.
 331 Some of the reported MsAr ages have much smaller uncertainties than other dating systems.
 332 Because we are focused on the current geometry and kinematics of the MHT and the small
 333 uncertainties of the potential outliers of the 9-10 Ma MsAr ages close to the PT2-N may mislead
 334 the inversion process, we change the uncertainties of the MsAr ages younger than 9 Ma to 0.5
 335 Ma and the uncertainties of the MsAr ages older than 9 Ma to 1.0 Ma when conducting inversion
 336 modeling, such that the inversion does not over-value the MsAr ages, which are usually much
 337 older than the ages of other dating systems. However, in the forward model, we used true
 338 uncertainties for all ages. The calculated misfit values of the inversion model and forward model
 339 are notated as misfit (Φ) and corrected misfit (Φ_{corr}) later.

340 4.2 Forward Model in Pecube and Model Parameters

341 Pecube calculates the location and temperature of points in the model using inputs of
 342 fault geometry, fault slip rate, flexural isostasy, radiogenic heat production, frictional heat
 343 production, and the boundary temperatures at the surface and the bottom of the model. These
 344 calculations are repeated at every timestep. The time-temperature history for each point is then
 345 used to calculate the expected cooling ages for each thermochronologic system, using known
 346 kinetic parameters for the system in question.

347 In Pecube, for a model without a localized zone of enhanced rock uplift, mass transport is
 348 parallel to each fault dip panel, and the fault geometries are defined by coordinate pairs (X, Y)
 349 defining the locations of fault dip change in a transect parallel with the mass-transport direction.
 350 In this study, the X and Y are the horizontal distance from the MFT fault trace, and vertical
 351 distances below sea level, respectively (Fig. 5). The slip rate along the MHT is defined by the
 352 convergence rate (V_{conv}) between the hanging wall and footwall and a partitioning factor (λ)
 353 partitioning the convergence into hanging-wall overthrusting, $V_o = (1 - \lambda)V_{\text{conv}}$, and footwall
 354 underthrusting, $V_u = \lambda V_{\text{conv}}$, with respect to the MHT (Fig. 5). This modified version of Pecube
 355 used in this study is the same as that used in Landry et al. (2016), and it deals with the change in
 356 the velocity in the hanging wall between different dip panels using a kink-band style fault
 357 kinematics described by Suppe (1983) — the velocity vectors change across the calculated
 358 planes bisecting the angle between any two neighboring fault dip panels, rather than using the
 359 velocity averaging approach in the original version. The other modification of the software is
 360 adding a function of applying a localized zone of enhanced rock uplift, similar to the version
 361 used in Herman et al. (2010). It simplifies the effect of localized crustal accretion by adding an
 362 extra vertical velocity component (V_{CA}) within the region defined by two specified distances
 363 (CA_{dist} and CA_{prox} for the southern and northern points respectively in this study) from the fault
 364 trace, MFT (Fig. 5). Rather than using the strategy that applies an enhanced uplift zone at a
 365 constant location with respect to the topography in the simulations conducted by Herman et al.
 366 (2010) and Landry et al. (2016), we apply a narrower enhanced uplift zone and make it translate
 367 hinterland-ward at an advection rate (CA_{adv}) to its present location which is defined by CA_{dist} and
 368 CA_{prox} in a certain amount of time (CA_{T}) to simulate one cycle of accretion (Fig. 5). Our strategy
 369 is more consistent with the kinematics for the generation of a new horse in a duplex, i.e. a newly
 370 generated ramp underthrusts beneath the slice of rock in hanging wall of the new ramp. The total
 371 uplift caused by the enhanced uplift zone in one cycle of crustal accretion on a cross-section is an
 372 isosceles trapezoid (Fig. 5), which is more consistent with geometry of a horse than the
 373 rectangle-shaped uplift generated by an enhanced uplift zone at a constant location. We only
 374 simulate the last cycle of accretion and set $CA_{\text{T}} = 6$ Ma to test the thermal effects since the
 375 specific kinematics of the entire multi-cycle accretion process in the mid-lower crust is unknown.
 376 The time span of one cycle of accretion is also unknown but we consider 6 Ma to be reasonable
 377 compared with the balanced cross-section coupled with thermokinematic modeling study in
 378 central Nepal Himalaya, the inversion result of the time span of the last accretion below the LHS
 379 duplex in northwestern Himalaya, and the geodynamic simulation of the crustal accretion in
 380 Himalaya (Mercier et al., 2017; Stübner et al., 2018; Ghoshal et al., 2020). We enable fault
 381 translation in all the models to simulate lateral advection of the model topography, such that all
 382 samples can translate laterally with respect to the fault but only move vertically relative to the
 383 topography.



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Figure 5 Illustration shows the thermokinematic model boundary conditions and parameters. The kinematic model has the convergence rate (V_{conv}) partitioned on either side of the megathrust into upper plate overthrusting component (V_{ot}) and lower plate underthrusting component (V_{ut}) using a partitioning factor (λ). The geometry of the megathrust is defined using a series of points along its length and the inverted free points, such as (X_1, Y_1) , (X_2, Y_2) , and $(300, Y_3)$, may occupy any position within each search box or on the error bar. Crustal accretion is carried out by adding a vertical uplift component (V_{ca}) to the velocity field within the range defined by two distance coordinates (CA_{dist} , and CA_{prox}). The range can advect from a position represented by dash box hinterland-ward to the present position defined by CA_{dist} and CA_{prox} in the past 6 Ma at an advection velocity of CA_{adv} . The effect of topographic evolution is simulated by assuming a linear evolution from initial topography (E_i) to the current topography (E_c) and assuming an initial topography by compress the relief of current topography ($E_c - E_{min}$) by an amplification factor (A_{topo}) and make it offset vertically by O_{topo} . For details of the model and other parameters refer to the text and Table 3.2 and 3.3.

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The dynamic thermal field of the model is calculated using an iterative solution to the finite-element formulation of the 3-D thermal advection-diffusion equation (Braun et al., 2012),

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$$\rho c \left(\frac{\partial T}{\partial t} + V \nabla T \right) = k \nabla^2 T + H$$

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where ρ is density, c is heat capacity, T is temperature, t is time, V is the velocity, k is thermal conductivity, and H is volumetric radiogenic heat production (for units and symbols, see Table 2). In our models, the temperature at the base of the model is set at 750 °C and does not change over time. Surface temperature decreases with elevation at an atmospheric lapse rate of 6 °C/km from 25 °C at 0 km. Thermal diffusivity is 35.0 km²/Myr. The bottom of each model is 50 km below the lowest point of the modeled area. Although we only use ages younger than 16 Ma, we run the models from 20 Ma to allow the samples in the model to accumulate enough displacement to be thermally reset at the beginning and to help the model obtain a natural, kinematic-affected, thermal field at 16 Ma. The default age of rocks in the model is 20 Ma.

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Table 2 Parameters of Pecube Models^a

Parameter Name	Parameter Range	Units	Parameter Symbol	Reference
<i>Material Properties</i>				
Thermal conductivity	2.5	W/m/K	k	Whipp et al. [2007]
Specific heat capacity	800	J/kg/K	c	Whipp et al. [2007]
Crustal density	2700	kg/m ³	ρ_c	
Upper mantle density	3200	kg/m ³	ρ_m	
Thermal diffusivity	35.0	km ² /My	α	
Volumetric radiogenic heat production		μ W/m ³	H	
Radiogenic heat production	15-25	°C/Myr	A	
Effective elastic thickness of the India plate	25.0	km		Jordan and Watts [2005] Berthet et al. [2013]
Young's modulus	100.0	GPa	E	
Poisson's ratio	0.3	n/a	ν	
<i>Pecube Model Parameters</i>				
Mean annual surface temperature in the foreland	25.0	°C	T _s	
Atmospheric lapse rate	6	°C/km	L	Landry et al. [2016]
Basal temperature	750	°C	T _b	
India-Eurasia convergence rate	Variable	mm/yr	V _{conv}	Bilham et al. [1997] Jouanne et al. [1999] Larson et al. [1999]
Convergence partitioning	0.5-0.7	n/a	λ	
Model time step	Optimal	years		
Horizontal node spacing	0.9	km		
Vertical node spacing (0-5 km)	0.9	km		
Vertical node spacing (5-15 km)	2.7	km		
Vertical node spacing (15-50 km)	8.1	km		
Model domain	255×122×50	km		
Fault geometry	Variable	km	(X _n , Y _n)	
Crustal accretion (south boundary)	Variable	km	CA _{dist}	
Crustal accretion (north boundary)	Variable	km	CA _{prox}	
Crustal accretion vertical rate	Variable	mm/yr	V _{ca}	
Crustal accretion window advection rate	Variable	mm/yr	CA _{adv}	
Amplification factor for evolving topography	Variable	n/a	A _{topo}	
Offset factor for evolving topography	Variable	km	O _{topo}	

^aThe bold entries indicate the free parameters that were inverted for in the simulations.

404

405 Previous studies dispute whether the Himalayan orogen had a similar topography in the
 406 Miocene to the present topography in High Himalaya (e.g. Masek et al., 1994; Wang et al., 2010;
 407 Gébelin et al., 2013; McDermott et al., 2013; Carrapa et al., 2016). Moreover, recent
 408 geomorphologic analysis in the Dolpo area suggests that glacial and fluvial erosion along with
 409 tectonics have significantly reshaped the extent of the orogenic plateau (Buceta et al., 2020).
 410 Therefore, we incorporate the effect of evolving topography in one inversion. Instead of coupling
 411 Pecube with a landscape evolution modeling software, as Herman et al. (2010) did, we simply
 412 assume an initial topography of the same shape but different relief and minimum elevation than
 413 the present topography, and assume a linear evolution from the initial topography over time (Fig.
 414 5). To make a synthetic initial topography, the software first keeps the elevation of the lowest
 415 point (E_{\min}) in the current topography unchanged and linearly compresses the current topography
 416 by an amplification factor (A_{topo}). After the compression, the elevation of any point on the
 417 surface changes from current elevation decreases to $E_{\min} + A_{\text{topo}}(E_c - E_{\min})$. The software then
 418 vertically moves the compressed topography by an offset distance (O_{topo}). This results in a
 419 synthetic initial topography, in which every point has an initial elevation, $E_i = E_{\min} + A_{\text{topo}}(E_c -$
 420 $E_{\min}) + O_{\text{topo}}$.

421 In models with evolving topography, flexural isostasy is enabled to compute the effect of
 422 isostatic adjustment due to evolving topography. The algorithm deals with the subducting plate
 423 as a thin elastic plate. We use 25.0 km for the effective elastic thickness of the Indian plate based
 424 on the reported Bouguer gravity anomaly data (Jordan and Watts, 2005; Berthet et al., 2013). We
 425 use 2700 kg/m^3 and 3200 kg/m^3 for crustal and upper mantle density respectively. Young's
 426 modulus (E) is 100 GPa and Poisson's ratio (ν) is 0.25 in the models.

427 We designed three inversions to simulate exhumation: (1) exhumation caused by the flat-
 428 ramp geometry of the MHT with a steady state topography ("ramp model", WNP01); (2)
 429 exhumation caused by both the flat-ramp geometry of the MHT and localized enhanced crustal
 430 accretion with a steady state topography ("ramp + duplex model", WNP02); (3) exhumation
 431 caused by both the flat-ramp geometry of the MHT and localized enhanced crustal accretion, and
 432 affected by the evolving topography ("ramp + duplex + evolving topography model", WNP03).
 433 We did not design a model for the out-of-sequence thrusting hypothesis, but discuss it based on
 434 our modeling results. All models have a dimension of $255 \text{ km} \times 122 \text{ km} \times 50 \text{ km}$. The
 435 topography data is down sampled to a resolution of approximately 900 m. To save computing
 436 time, we used a varying node spacing at different depths in the models; 0.9 km spacing for the
 437 upper 0-5 km, 2.7 km spacing for the 5-15 km part, and 8.1 km spacing for the 15-50 km part.
 438 For a list of the model parameters, including their values, units, symbols and important
 439 references, refer to the Table 2. We use the inversion algorithm of Pecube to invert for
 440 parameters specifically characterizing the three tectonic scenarios. For the geometry of the MHT
 441 in scenarios WNP01 and WNP02, because we focus on the exhumation caused by the deep
 442 structures in the hinterland of the wedge, we only invert for the coordinates of the northernmost
 443 three nodes among the nodes defining the geometry of the MHT, notated as $(X1, Y1)$, $(X2, Y2)$,
 444 and $(X3, Y3)$ (Fig. 5). For the northernmost dip panel of the MHT, we only invert for $Y3$ in the
 445 models and set $X3 = 300 \text{ km}$ to find a proper dip angle. For all other nodes to the south of these
 446 three nodes, primarily defining the upper-crustal flat, the southern small ramp, and the MFT, we
 447 set their coordinates based on seismic images, balanced cross-sections, and microseismicity, and
 448 do not invert for them. Previous studies suggest that the basal temperature and radiogenic heat
 449 production are correlated and can compensate each other to some extent. Because we define the

450 basal temperature at 750 °C, we invert for the radiogenic heat production, A (in °C/Myr). We
 451 also inverted for the convergence rate and partitioning factor for several reasons; (1) Current
 452 GPS data suggests convergence rate is in a wide range of 14-21 mm/yr; (2) the convergence rate
 453 and the partitioning between overthrusting and underthrusting over the past 20 Ma are not well
 454 known; (3) Internal shortening within the upper plate is not simulated in the simplified models,
 455 which may cause inconsistency between the models and the real-world geology even if using the
 456 correct values. For scenario WNP02, we also invert for the four parameters defining the crustal
 457 accretion (CA_{dist} , CA_{prox} , CA_{adv} , and V_{CA}). The inverted parameters for scenario WNP03 are
 458 slightly different from scenario 2; (1) We invert for the two parameters defining the evolving
 459 topography, A_{topo} and O_{topo} ; (2) To reduce the dimension of the parameter space, we set the Y1
 460 and Y3 as constants respectively based on the inversion result of scenario WNP02. **Table 3**
 461 shows the parameter values and the ranges of the values used in each inversion set.

Table 3 Inversion Results^a

462

	Inversion Name		
	WNP01	WNP02	WNP03
Number of models	22220	28820	28820
Φ/Φ_{corr}	0.26/0.78	0.22/0.55	0.22/0.55
A_{topo}	n/a	n/a	0.96 (0.5:1.0)
O_{topo} (km)	n/a	n/a	0.14 (0.0:2.0)
Tb (°C)	750	750	750
A (°C/Myr)	24.89 (15:25)	24.94 (15:25)	24.63 (15:25)
Y3 (km)	49.9 (45:50)	48.5 (45:50)	50
Y2 (Km)	25.0 (25:30)	25.4 (25:30)	25.9 (25:30)
X2 (Km)	220.3 (215:230)	224.6 (210:225)	224.5 (210:225)
Y1 (Km)	15.2 (15:20)	15.2 (15:20)	15
X1 (Km)	167.3 (160:175)	149.2 (140:155)	147.2 (140:155)
V_{conv} (mm/yr)	17.1 (15:20)	15.5 (14:20)	14.5 (14:20)
λ	0.52 (0.5:0.7)	0.52 (0.5:0.7)	0.50 (0.5:0.7)
V_{ca} (mm/yr)	n/a	4.0 (3.0:6.0)	5.0 (3.0:6.0)
CA_{dist} (km)	n/a	189.0 (175:190)	189.3 (175:190)
CA_{prox} (km)	n/a	203.2 (190:210)	203.3 (190:210)
CA_{adv} (mm/yr)	n/a	5.2 (4.0:6.0)	5.2 (4.0:6.0)

^aValues in brackets indicate the investigated range of the free parameters and values in bold indicate lowest misfit values for each parameter.

463

464 4.3 Inversion Algorithm: Neighborhood Algorithm

465 When combined with the Neighborhood Algorithm (NA) (Sambridge, 1999a, b;
 466 Rickwood and Sambridge, 2006), Pecube can perform inversion modeling that searches for a
 467 combination of parameters depicting a forward model that produce predicted cooling ages

468 consistent with the observed ages. The algorithm searches within the multidimensional parameter
469 space defined by given ranges of free parameters.

470 The NA deals with the inversion in two stages. In the first stage, the software selects the
471 first set of forward models by randomly selecting parameters from their defined ranges and uses
472 the selected parameters to divide the model parameter space into Voronoi cells (Sambridge,
473 1999a). The parameters for subsequent forward models are selected from within the subset of
474 Voronoi cells with a low misfit (Φ) to the observed ages. During the inversion, each subsequent
475 forward model is selected from 80-85% (resampling ratio, varies with different scenarios) of the
476 current Voronoi cells with the lowest misfits.

477 The second stage of the NA is an appraisal of the search results to define statistical limits
478 on the ranges of input parameters that provide a good fit to the observed age data (Sambridge,
479 1999b). Bayesian inference is used to produce posterior probability density functions (PPDFs)
480 for each model parameter using a likelihood function L ,

$$481 \quad L = \exp\left(-\frac{n}{2} \sqrt{\sum_{i=1}^n \frac{(\text{PredAge}_i - \text{ObsAge}_i)^2}{\sigma_i^2}}\right)$$

482 the appraisal yields 1-D and 2-D PPDFs for the model parameters that are presented for
483 each set of model parameters.

484 5 Modeling Results

485 For the three inversion sets presented below (Fig. 6-8), the parameter ranges are
486 determined based on multiple trials. To evaluate the defined ranges of the inverted parameters in
487 each trial, we use reasonably small ranges of parameters and compute more than 20,000 forward
488 models in each inversion. If the inversion converges at the end of the defined ranges for many
489 inverted parameters, we change the ranges of the corresponding parameters in the next inversion
490 trial. A summary of the parameter ranges of the presented inversions, parameter values of the
491 best-fit model in each inversion, and the misfit value of the best-fit model for the presented
492 inversion for each scenario is shown in Table 3. As introduced above, we use synthetic errors for
493 some ages in the inversion and calculate the corresponding misfit, Φ . We then select the best
494 model found in each inversion and calculate the corrected misfit, Φ_{corr} , using true errors of all
495 ages.

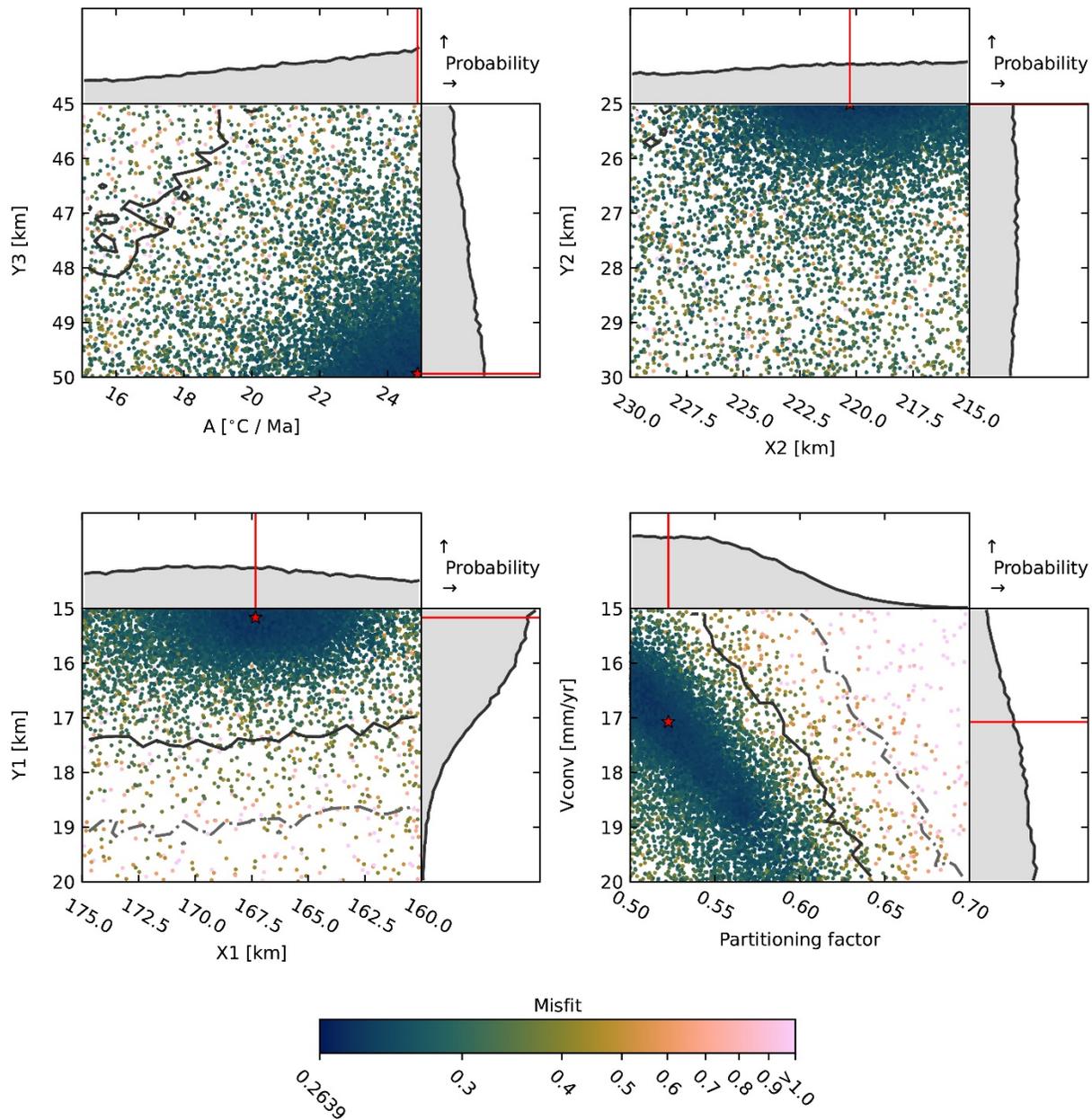
496 5.1 Inversion Set WNP01: Slip on the MHT with Steady State Topography

497 The models in the inversion set WNP01 simulate a scenario where exhumation is caused
498 only by slip along the MHT with steady topography (ramp model). The inversion set comprises
499 22,220 forward models. Using a resampling ratio of 85%, the inversion converges to a misfit $\Phi <$
500 0.3 after approximately 8,000 models. For the best-fit model, the misfit is $\Phi = 0.26$ and the
501 corrected misfit $\Phi_{\text{corr}} = 0.78$.

502 The inversion result indicates that the misfit of model in this scenario is sensitive to the
503 fault geometry (Fig. 6). The distance between the mid-lower crustal ramp and the MFT (X1 and
504 X2) is well constrained. The depths of the inverted nodes of the fault in low-misfit models
505 suggest a shallow mid-crustal ramp indicated by the low Y1 and Y2 and a steep hinterland sector
506 indicated by the large Y3 (Fig. 9). The best-fit model has a mid-crustal ramp starting from 167.3

507 km to 220.3 km from the MFT and has a dip of 10.5°. The hinterland sector of the MHT in the
508 best-fit model has a dip of 17.4°. These three depth parameters converge to the end of the given
509 ranges in the inversion but we didn't test new ranges because the ranges of these parameters are
510 consistent with other geophysical and geologic data. The convergence rate and partitioning factor
511 for low-misfit models are correlated in a narrow zone on the 2D PPDF plot (Fig. 6), suggesting
512 that the misfit is very sensitive to the slip rate along the MHT at the first order but the effect of
513 convergence and the partitioning between the underthrusting and overthrusting can compensate
514 each other to some extent. The best-fit model has a convergence rate of 17.1 mm/yr with 52% of
515 it accommodated by footwall underthrusting. Both of these values are reasonable compared with
516 previous studies (Bilham et al., 1997; Jouanne et al., 1999; Larson et al., 1999; Coutand et al.,
517 2014; Landry et al., 2016). Because we used a constant temperature at the bottom of the model,
518 we inverted for the heat production over a wide range, 15-25 °C/Myr, and the inversion
519 converges to the high end of the range, suggesting a hot orogen.

520 The comparison between the predicted ages of the best-fit model in this inversion set and
521 the observed ages is shown in Figure 9. Some ages not used in the inversion or misfit calculation
522 are also included in the plot. Generally, for the low-temperature-system ages used in the
523 inversion, including AHe, ZHe, and AFT ages, most predicted ages match the observed ages
524 within 1 standard uncertainty. Age cluster of AHe and ZHe between 29.5°-29.6° N (132-135 km
525 from the MFT) are not used in the inversion, they have a worse match than the ages used in the
526 inversion. Among the four ZHe and four AHe ages in the cluster, only two predicted AHe ages
527 overlap with the observed ages within 1 standard uncertainty, but the discrepancy is usually
528 within 2 standard uncertainties. For the MsAr ages, there is a large discrepancy between the
529 predicted and observed ages for rocks in the hinterland close to the PT2-N. Among the two
530 previously introduced groups of MsAr ages close to the high-slope zone, the predicted ages are
531 closer to the group of 9-10 Ma data than the ca. 6 Ma ages. However, ages of ca. 6 Ma or
532 younger are common along the high-slope zone in other segments of the Himalaya. For example,
533 one MsAr age of 5.85 ± 0.67 Ma is reported at 29.73°N (139 km from the MFT) in this
534 region (Mercier, 2014), close to the high-slope zone to the west of the modeled profile across a
535 proposed lateral ramp in the MHT (Fig. 4). Thus, we interpret that this model fails to produce
536 MsAr ages consistent with observation close to the high-slope zone. The predicted MsAr ages of
537 the samples from the southernmost part of the model are consistent with observed ages (age
538 difference < 1 Ma).



539

Figure 6 Inversion results of model WNP01 solving for 8 free parameters (see Table 3.3) with Y_3 versus A , Y_2 versus X_2 , Y_1 versus X_1 , and V_{conv} versus λ . Each dot represents a single forward model, and its color corresponds the goodness of fit to the data. The red star represents the parameter values obtained from the forward model with the lowest misfit. One-dimensional posterior probability density functions (1-D PPDFs) derived from the NA appraisal are shown adjacent to the axes for each parameter. The red lines indicate parameter values for the lowest misfit forward model. Two-dimensional PPDFs (2-D PPDFs) are represented by lines overlying the scatter diagram where the solid black line is the 1 σ confidence interval and the dashed line the 2 σ confidence interval.

540

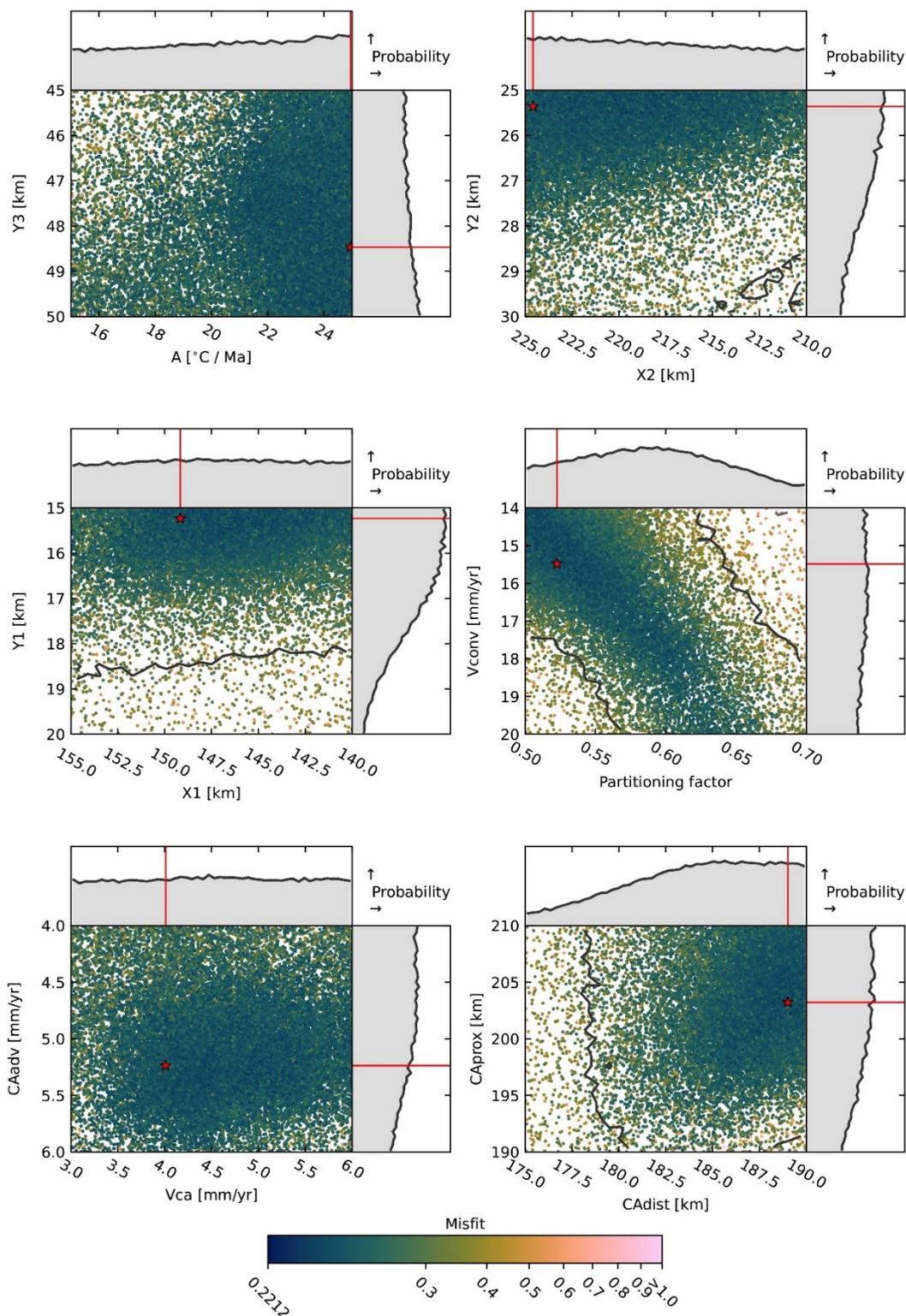
541 5.2 Inversion Set WNP02: Slip on the MHT and Localized Enhanced Rock Uplift at Steady
542 Topography

543 The models in the inversion set WNP02 simulate a scenario where exhumation is caused
544 by a combination of slip along the MHT and localized enhanced uplift by crustal accretion with
545 steady topography (ramp + duplex model). The inversion set comprises 28,820 forward models
546 (Fig. 7). Using a resampling ratio of 80%, the inversion converges to a misfit $\Phi < 0.35$ after
547 approximately 2,000 models. For the best-fit model, the misfit is $\Phi = 0.22$ and the corrected
548 misfit $\Phi_{\text{corr}} = 0.55$.

549 Similar to the inversion set WNP01, the misfit values are sensitive to the convergence
550 rate and underthrusting-overthrusting partitioning factor as suggested by the strongly correlated
551 pattern for the low-misfit forward models on the 2D PPDF plot (Fig. 7). The best-fit model in
552 this set has a convergence rate of 15.48 mm/yr with 0.52% of it accommodated by
553 underthrusting. Both values are broadly consistent with previous studies (Bilham et al., 1997;
554 Jouanne et al., 1999; Larson et al., 1999; Coutand et al., 2014; Landry et al., 2016). The locations
555 of distal and proximal ends of the crustal accretion zone are well-constrained. The rates of
556 localized enhanced uplift and accretion-zone advection are less well-constrained probably due to
557 lack of age constraints in the inversion for the northernmost area. The best-fit mode has a
558 localized enhanced uplifting zone translated at a rate of 5.24 mm/yr to its present location at
559 189.0 – 203.2 km from the MFT with a vertical crustal accretion rate of 4.01 mm/yr. Due to the
560 existence of an enhanced uplift zone close to the high-slope zone in the model, the dip of the
561 northernmost sector of the MHT, different from the inversion set WNP01, has more freedom in
562 this inversion as suggested by the diffuse distribution of the Y3 values of low-misfit models (Fig.
563 7). This suggests that the existence and kinematics of a crustal accretion zone plays a more
564 determining role than the dip of the hinterland sector of the MHT. In this inversion, the mid-
565 crustal ramp in the MHT is well-constrained. The best-fit mode has a mid-crustal ramp starting
566 from 149.2 km to 224.6 km from the MFT and has a dip of 7.6° (Fig. 9). The hinterland sector of
567 the MHT in the best-fit model has a dip of 17.0°. The inversion of the heat production parameter
568 also converges to the high end of the given range, with a best-fit value of 24.94 °C/Myr,
569 suggesting a hot orogen.

570 The comparison between the estimated ages in the best-fit model and the observed ages is
571 shown in Figure 9. Similar to the best-fit model in WNP01 inversion, for the low-temperature
572 systems, the estimated ages also can generally match the observed ages, including the ages not
573 used in the inversion. The large decrease in corrected misfit of the best-fit model compared with
574 the one in WNP01 is mainly contributed by a significantly improved fit for the data cluster close
575 to the high-slope zone (160-185 km). The difference between the observed MsAr ages and
576 predicted ages there are mostly less than 2 Ma. The estimated MsAr ages are younger than the
577 estimation in the best-fit model of the WNP01 inversion, and therefore are more consistent with
578 the young MsAr ages (as young as < 5 Ma) close to the high-slope zone along-strike in other
579 segments (e.g. Copeland et al., 1991; Huntington and Hodges, 2006; Martin et al., 2015; Gibson
580 et al., 2016). The estimated MsAr ages of the samples from the southernmost part of the model
581 also have a good fit with the observed ages (age difference < 1.5 Ma). We interpret that this
582 model is overall successful in producing ages consistent with the observed ages for all age
583 systems discussed.

584



585

Figure 7 Inversion results for model WNP02 solving for 12 parameters (see Table 3.3) with Y_3 versus A , Y_2 versus X_2 , Y_1 versus X_1 , V_{conv} versus λ , CA_{adv} versus V_{ca} , and CA_{prox} versus CA_{dist} . Symbols are the same as for Figure 6.

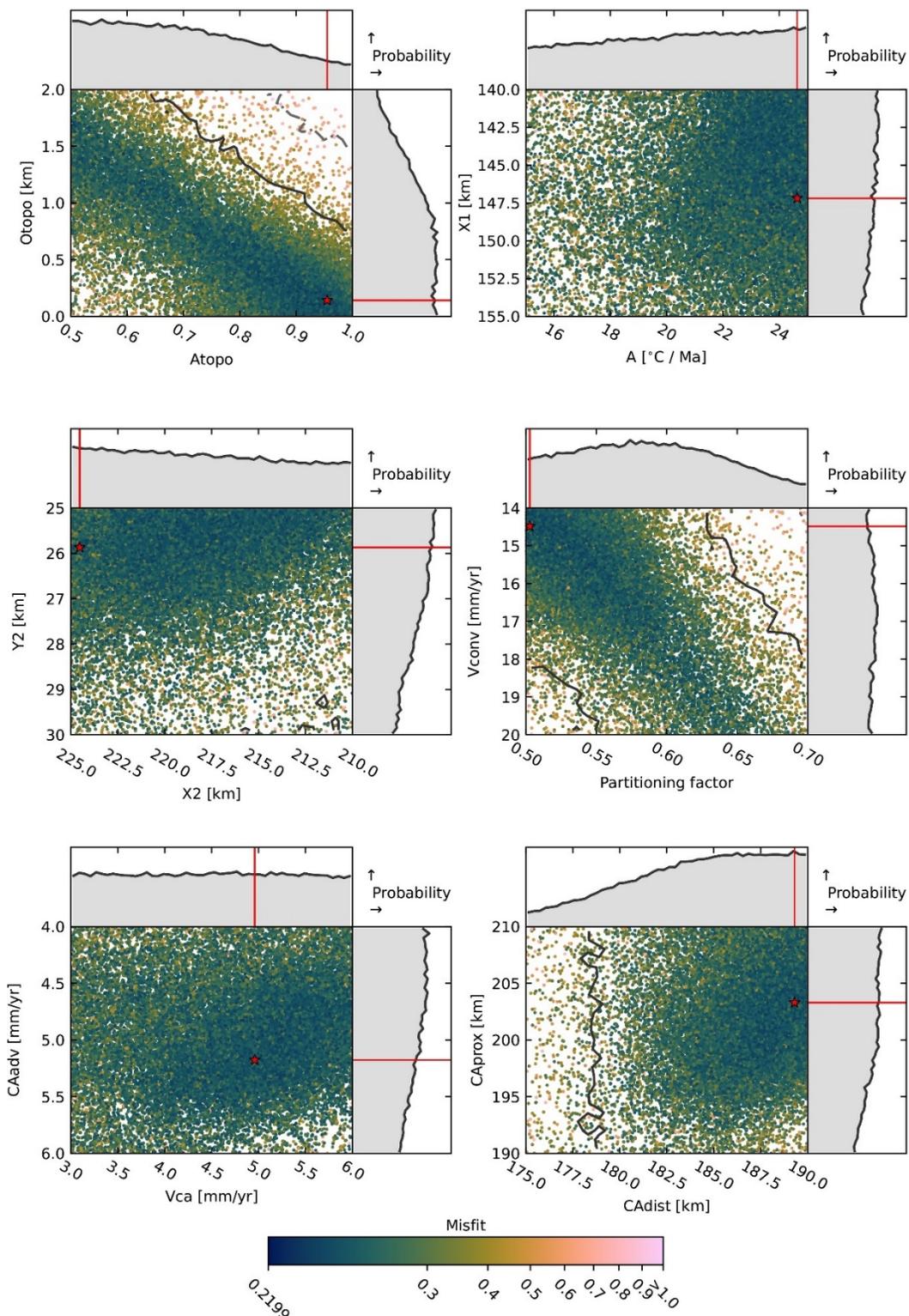
586

587 5.3 Inversion Set WNP03: Slip on the MHT, Localized Enhanced Rock Uplift, and Evolving
588 Topography

589 The models for inversion set WNP03 simulate a scenario similar to WNP02 except that it
590 has a dynamically evolving topography (ramp + duplex + evolving topography model). Because
591 of the introduction of two parameters defining the evolving topography, to keep the number of
592 parameter-space dimensions unchanged, we use constant values for the vertical coordinates of
593 the hinterland most node of the MHT ($Y_3 = 50.0$ km) and the upper end of the ramp ($Y_1 = 15.0$
594 km). The value of Y_1 is selected based on the consistent Y_1 values of the best models in the
595 inversion sets WNP01 and WNP02. We also change Y_3 to a constant parameter because the
596 inversion result of WNP02 shows, when crustal accretion is incorporated, Y_3 cannot be well-
597 constrained. This inversion set comprises 28,820 forward models. Using a resampling ratio of
598 80%, the inversion has a worse convergence than the WNP02 inversion set, but it still converges
599 to a misfit $\Phi < 0.35$ after approximately 28,000 models (with only one model exception). The
600 lowest misfit is $\Phi = 0.22$, and the corrected misfit $\Phi_{\text{corr}} = 0.55$.

601 The inversion result is similar to WNP02 in relatively less well-constrained fault-bend
602 locations than WNP01 (related parameters X_1 , X_2 , Y_2), highly correlated convergence rate
603 (V_{conv}) and underthrusting-overthrusting partitioning factor (λ), well-constrained distal and
604 proximal ends of the crustal accretion zone and uplift rate (CA_{dist} and CA_{prox}), not well-
605 constrained uplift rate and advection velocity of the crustal accretion zone (V_{CA} and CA_{adv}) (Fig.
606 8). The introduction of inversion for evolving topography makes the distribution of good-fit X_1 ,
607 X_2 , and Y_2 more diffuse than the result of WNP02 in their 2D PPDF plots. The best-fit model
608 has a convergence rate of 14.48 mm/yr. with 50% of it accommodated by underthrusting. Both of
609 these values are broadly consistent with previous studies (Bilham et al., 1997; Jouanne et al.,
610 1999; Larson et al., 1999; Coutand et al., 2014; Landry et al., 2016). The best-fit geometry and
611 kinematics of the MHT, including crustal accretion is approximately similar to the best model of
612 WNP02. The best-fit MHT has a shallow-dipping middle crustal ramp of an 8.0° dip, starting
613 from 147.2 km to 224.5 km from the MFT, and a hinterland sector of 17.7° . The best-fit mode
614 has a localized enhanced uplift zone translated at a rate of 5.18 mm/yr to its present location at
615 189.3 – 203.3 km from the MFT with a vertical crustal accretion rate of 5.00 mm/yr during the
616 last 6 Ma in the model. The inversion of the heat production parameter converges to 24.63
617 $^\circ\text{C}/\text{Myr}$, also indicating a hot orogen. Like the correlation between the convergence rate and
618 partitioning factor of the low-misfit models, the parameters defining the initial topography, A_{topo}
619 and O_{topo} , are also well correlated in a narrow zone on the 2D PPDF plot. The best-fit A_{topo} and
620 O_{topo} are 0.955 and 0.14 km, indicating no significant topographic evolution in the best-fit model.

621 Between the inversion sets WNP02 and WNP03, the parameters of the best-fit models are
622 not significantly different, and they have the same level of observation-prediction fit (Fig. 9).
623 The pattern in the estimated ages is similar to that of the best-fit model in WNP02. We conclude
624 that this model is overall successful in producing ages consistent with the observed ages for all
625 age systems discussed and that whether the topography significantly evolved over time cannot be
626 determined with the methods used in this study.



627

Figure 8 Inversion results for model WNP03 solving for 12 parameters (see Table 3.3) with O_{topo} versus A_{topo} , $X1$ versus A , $Y2$ versus $X2$, V_{conv} versus λ , CA_{adv} versus V_{ca} , and CA_{prox} versus CA_{dist} . Symbols are the same as for Figure 6.

628

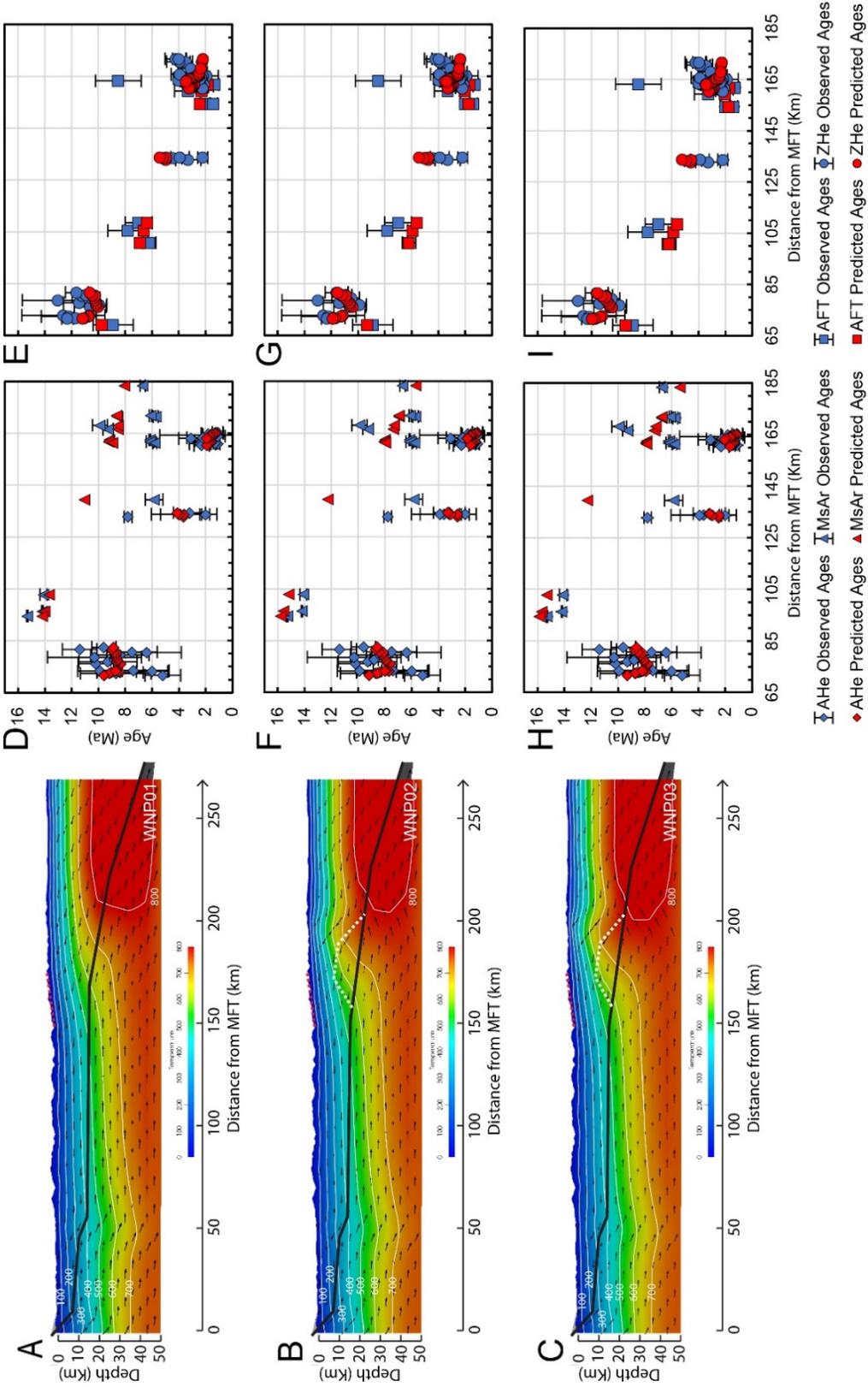


Figure 9 Thermal structures of the best-fit models of the 3 modeled scenarios (A-C) and comparison between the predicted and observed cooling ages for each model (D-I). White dash lines show the amount and location of the crustal accretion along the megathrust. The red dash lines show the high-slope zone on the topography. Black arrows show present snapshot of the velocity field for each model.

630 6 Discussion

631 6.1 Mid-lower Crustal Duplex Causing Hinterland Exhumation

632 Comparing the results of the models with and without the involvement of crustal creation
633 shows that mid-lower crustal accretion is required in a model that can produce observation-
634 fitting cooling ages. While each of the three best-fit models can produce good-fit cooling ages of
635 low-temperature systems, only the two best-fit models involving crustal accretion can also
636 produce observation-fitting MsAr ages.

637 Thermo-kinematic modeling has been used in many studies to test the crustal-accretion
638 model of the exhumation, tectonics, and metamorphism in the Himalayan wedge. Similar work
639 in the central Nepal Himalaya also suggests that a model with crustal accretion along a mid-
640 crustal ramp in the MHT can produce good-fit cooling ages of MsAr, ZHe, and AFT (Herman et
641 al., 2010). The other study in the central Nepal Himalaya used a different version of Pecube
642 which can be coupled with detailed kinematics based on balanced cross-sections to assess
643 different structural reconstructions (Ghoshal et al., 2020). They also found that only models
644 involving hinterland-dipping duplexing and the thrusting of the duplex over a mid-crustal ramp
645 can produce sufficiently rapid exhumation to produce the observed young MsAr ages as well as
646 low-temperature cooling ages. An equivalent study in Sikkim Himalaya by Landry et al. (2016)
647 did not investigate MsAr ages, but, by comparing the model-produced and observed ZHe and
648 AHe ages they also favor a model with crustal accretion. Equivalent studies in Eastern Bhutan,
649 Kumaun, and Garwhal Himalaya did not test competing kinematic models, but confirmed that
650 models involving crustal accretion can produce observation-fitting cooling ages of various
651 geochronometers including MsAr, ZHe, and AFT (C  lerier et al., 2009; Grujic et al., 2020).
652 Similar thermokinematic models also invoked duplexing or complex kinematics requiring matter
653 accretion from the subducting plate to the hanging wall of the MCT to produce both the inverted
654 metamorphic gradient recorded in the rocks across the MCT and the exhumation of the LHS
655 (Harrison et al., 1997; Bollinger et al., 2004; Grujic et al., 2020).

656 Some studies using other approaches also suggest that crustal accretion is a main
657 mechanism accommodating mid-lower crustal strain. Within the MCT thrust sheet, many thrust-
658 sense tectonic or metamorphic discontinuities have been documented (e.g. Montomoli et al.,
659 2013; Larson et al., 2015; Montomoli et al., 2015; Carosi et al., 2016; Carosi et al., 2018). These
660 discontinuities approximately mimic the PT2 on the map and usually correspond to the areas on
661 the foreland limb of the antiformal crustal accretion in the thermokinematic models discussed
662 above, indicating mid-lower crustal strain accumulation. Underplating of the material from the
663 Indian plate to the upper plate is supported by isotopic signatures of the high-grade crystalline
664 rocks in the Gurla Mandhata area along the Northern Himalayan antiform (Murphy, 2007; Godin
665 et al., 2021). Direct observation of a crustal-scale duplex structure on seismic-reflection profiles
666 across the Yarlung-Zangbo suture were documented (Gao et al., 2016; Guo et al., 2018) or
667 reinterpreted (Laskowski et al., 2018). The similar inclinations recorded in the GHS and LHS
668 indicated by the secondary remnant magnetization formed during metamorphism contradict to
669 MCT ramping and rather favors the interpretation of duplex structures (Schill et al., 2004).
670 Landscape evolution simulations in the Bhutan and central Nepal Himalaya suggest that crustal
671 accretion at depth can facilitate in-situ formation of low-relief landscape and a physiographic
672 transition at the front of it (Herman et al., 2010; Adams et al., 2016). Similarly, Fan and Murphy
673 (2021) proposed that strain accumulation in mid-lower crust through duplexing is a mechanism

674 of plateau outward growth based on the observation that the thick high-grade metamorphic core
675 of the orogen correlates well with the high-elevation, low-relief landscape to the north of the
676 present active-uplift front. Grandin et al. (2012) addressed that the interseismic transient-uplift
677 peak indicated by interferometric synthetic aperture radar data spatially matches the long-term
678 uplift peak indicated by the study on trans-Himalayan river incision in central Nepal. This
679 pattern requires formation of a new crustal ramp via footwall-to-hanging-wall accretion as the
680 old ramp translated towards the hinterland. Similarly, a steady position of the active uplift with
681 respect to the PT2 since ca. 1.5 Ma in northwestern Himalaya is suggested by the spatial
682 similarity of the long- (Ma) and short-term (ka) erosion rates, which lead Morell et al. (2017) to
683 draw the same conclusion.

684 There are two main groups of studies invoking the ramp model without the involvement
685 of crustal accretion to explain observations from a variety of disciplines, but their data usually
686 cannot represent long-term orogen-building process. The first group used mechanical modeling
687 methods to simulate the deformation depicted by geodetic data within the time span of seismic
688 cycles, but these data record deformation no older than decades (e.g. Jackson and Bilham, 1994;
689 Bilham et al., 1997; Jouanne et al., 1999; Larson et al., 1999; Cattin and Avouac, 2000; Berger et
690 al., 2004; Godard et al., 2004; Elliott et al., 2016). Some of these studies also considered the
691 effects of rheology change and erosion in their models, and compared the model predictions with
692 results of erosion and denudation studies which usually do not account for processes older than
693 Quaternary (Cattin and Avouac, 2000; Lavé and Avouac, 2001; Godard et al., 2004, 2009). The
694 other group of studies that support the ramp model conducted thermokinematic modeling similar
695 to this study, but they usually did not test effect of crustal accretion (Robert et al., 2009; Robert
696 et al., 2011; Coutand et al., 2014). They also used only ages of low-temperature chronometers to
697 test the models which may not be able to determine the feasibility of continuous crustal accretion
698 (this study, Whipp et al., 2007; Ghoshal et al., 2020). Some other studies invoked the ramp
699 model to conceptually explain some geological observations without quantitative test, and
700 therefore do not conflict with the models invoking crustal accretion (e.g. Seeber et al., 1981;
701 Molnar, 1984; Ni and Barazangi, 1984; Beaumont, 2004; van der Beek et al., 2016).

702 The other competing kinematic model to explain the PT2 and the rapid exhumation along
703 it is the out-of-sequence model (Harrison et al., 1997; Wobus et al., 2003; Hodges et al., 2004;
704 Thiede et al., 2004; Thiede et al., 2005; Wobus et al., 2005; Wobus et al., 2006; Whipple et al.,
705 2016). We did not directly test this model in this study because this model contradicts several
706 independent lines of observations. If there is an active thrust fault accommodating the hinterland
707 active uplift and maintaining the PT2 and rapid exhumation, there should be an abrupt change in
708 the cooling ages across the active fault. However, age profiles along most of the investigated
709 transects do not show an abrupt change, especially for the ages of low-temperature dating
710 systems (e.g. Thiede and Ehlers, 2013; Coutand et al., 2014; Landry et al., 2016; Stübner et al.,
711 2018; McQuarrie et al., 2019; Ghoshal et al., 2020). Although several lines of evidence for active
712 out-of-sequence thrusting or reactivation of the MCT have been reported at several places (e.g.
713 Harrison et al., 1997; Catlos et al., 2001; Wobus et al., 2003; Hodges et al., 2004; Wobus et al.,
714 2005; Wobus et al., 2006; Whipple et al., 2016; Thiede et al., 2017; Braden et al., 2018), no
715 continuous active thrust fault along the generally orogen-wide continuous (> 1,500 km long) PT2
716 has yet been mapped. Finally, thermokinematic modeling in central Nepal shows that an out-of-
717 sequence model producing observation-fitting exhumation pattern requires either unrealistic fast

718 movement along the thrust fault and the STD or mechanically unfeasible high-dip angle of the
719 thrust fault (Herman et al., 2010).

720 6.2 3D Megathrust Ramps

721 Many studies suggest that the MHT has complex along-strike heterogeneity regarding the
722 dip, location and number of the fault ramps (Hauck et al., 1998; Larson et al., 1999; Robert et al.,
723 2011; Mugnier et al., 2017; Fan and Murphy, 2021). For the western Nepal Himalaya, different
724 MHT geometries from the central Nepal Himalaya have been proposed based on observations
725 from a variety of disciplines, including structural geology (Robinson et al., 2001; Hubbard et al.,
726 2016; DeCelles et al., 2020), thermochronology (Gibson et al., 2016; van der Beek et al., 2016),
727 metamorphism (Soucy La Roche and Godin, 2019), seismicity (Hoste-Colomer et al., 2018), and
728 landscape analysis (Harvey et al., 2015). However, most of these models are conceptual and are
729 not quantitatively constrained. Some models based on balanced cross-sections focus on the
730 structures primarily to the south of the PT2 and do not deal with the mid-lower crustal structures
731 or MHT crustal ramps (e.g. DeCelles et al., 2001; Robinson et al., 2006; Robinson, 2008; Olsen
732 et al., 2019) (Fig. 3). Geophysical data used in interpreting the MHT geometry are usually of
733 poor quality for deep structures in the hinterland (e.g. Caldwell et al., 2013; Subedi et al., 2018).
734 The numerical models in this study provide the first assessment of the MHT hinterland geometry
735 in western Nepal Himalaya from the perspective of thermokinematics.

736 The best models of inversion sets WNP02 and WNP03 are very similar and we consider
737 both to be the preferred models based on their successful prediction of the calculated cooling
738 ages (Fig. 9). The upper-crustal portion of the MHT (shallower than 14 km) are not investigated
739 in this study, and in our model the geometry of this part is adopted from the two-ramp model
740 based on geomorphology, thermochronology, microseismicity, and balanced cross-sections
741 (DeCelles et al., 2001; Harvey et al., 2015; van der Beek et al., 2016; Hoste-Colomer et al.,
742 2018; Olsen et al., 2019). The upper-crustal ramp connects to a long flat at approximately 14-15
743 km deep in our preferred models. This flat is imaged as a midcrustal low-velocity zone from
744 receiver function analysis on teleseismic waveforms (Subedi et al., 2018), and its southern extent
745 is consistent with the distribution of microseismicity (Ader et al., 2012; Hoste-Colomer et al.,
746 2018). At about 149 km to the north of the MFT, this midcrustal flat connects to a not-well-
747 expressed mid-lower crustal ramp, which is consistent with the northward lack of intense
748 microseismicity as the ramp extends beneath seismogenic depths (Fig. 3, 9). A sector of low-
749 velocity zone to the north but discontinuous from the midcrustal flat on the receiver function
750 image also in agreement with the gently dipping mid-lower crustal ramp geometry in our
751 preferred models (Fig. 3). The crustal accretion in our preferred models is on this mid-lower
752 crustal ramp, indicating that PT2-N in western Nepal Himalaya represents the active-uplift front
753 of the high Himalaya plateau, which is consistent with previous interpretations (e.g. Cannon and
754 Murphy, 2014; Fan and Murphy, 2021) (will be discussed in the next section) (Fig. 9). The mid-
755 lower crustal ramp connects to a steeper hinterland ramp at ca. 25 km depth in our preferred
756 models. But this hinterland ramp should be interpreted with caution because no data, including
757 cooling ages, are available to test it. It is possible that it is a trade-off product of the inversion
758 caused by the limitation of the model design: The northern Himalaya anticline is thought to start
759 its thickening process as early as Early-Middle Miocene (e.g. Murphy et al., 2002; Murphy and
760 Copeland, 2005; Godin et al., 2006; Fan and Murphy, 2021) and thus should have thickened via
761 multiple cycles of crustal accretion. However, in the numerical models, we only simulate one

762 accretion cycle. Therefore, the thermal effect of a hot hinterland due to multiple cycles of
763 accretion prior to the simulated last accretion cycle can be compensated by fast exhumation
764 along a steep hinterland ramp in the simulation.

765 Our preferred models are different from the MHT geometry in the central Nepal
766 Himalaya which is characterized by one large mid-lower crustal ramp at approximately 100 km
767 from the MFT. The differences in the number and location of MHT ramps between the western
768 and central Nepal Himalaya require lateral or oblique ramps in the MHT. A 3D conceptual
769 model with a lateral ramp has been invoked to explain the different P-T-t paths of metamorphic
770 rocks from approximately the same structural position but different along-strike segments in
771 western Nepal Himalaya (Soucy La Roche and Godin, 2019). Based on a balanced cross-section
772 in western Nepal and the assumption of an along-strike constant amount of shortening, DeCelles
773 et al. (2020) discussed the 3D evolution of the upper-crustal structures in midwestern Nepal area,
774 but the model does not address the deeper structures in the hinterland. Based on the finding that a
775 balanced cross-section with LHS duplex on a mid-crustal ramp can fit well with the MHT
776 geometry revealed by seismic observations from the 2015 M_w 7.8 Gorkha earthquake, Hubbard
777 et al. (2016) assumed the constant kinematics along-strike and proposed a 3D MHT model for
778 the Nepal Himalaya using the axis of the LHS anticlinorium as a proxy of the top of the mid-
779 lower crustal ramp. Fan and Murphy (2021) synthesized different datasets to propose a 3D
780 evolution model of the MHT in the midwestern Nepal Himalaya and discussed the mechanism of
781 orogenic wedge growth. Their results support the idea that the inner wedge or the orogenic
782 plateau grow outward through crustal accretion at the bottom of the brittle-ductile transition
783 zone, the location of which is mainly controlled by the geometry, especially the mid-lower
784 crustal ramp, of the MHT. This hypothesis predicts that the along-strike variation in the location
785 of active-uplift front of the plateau or the high slope zone is the surface expression of the along-
786 strike variation in MHT geometry.

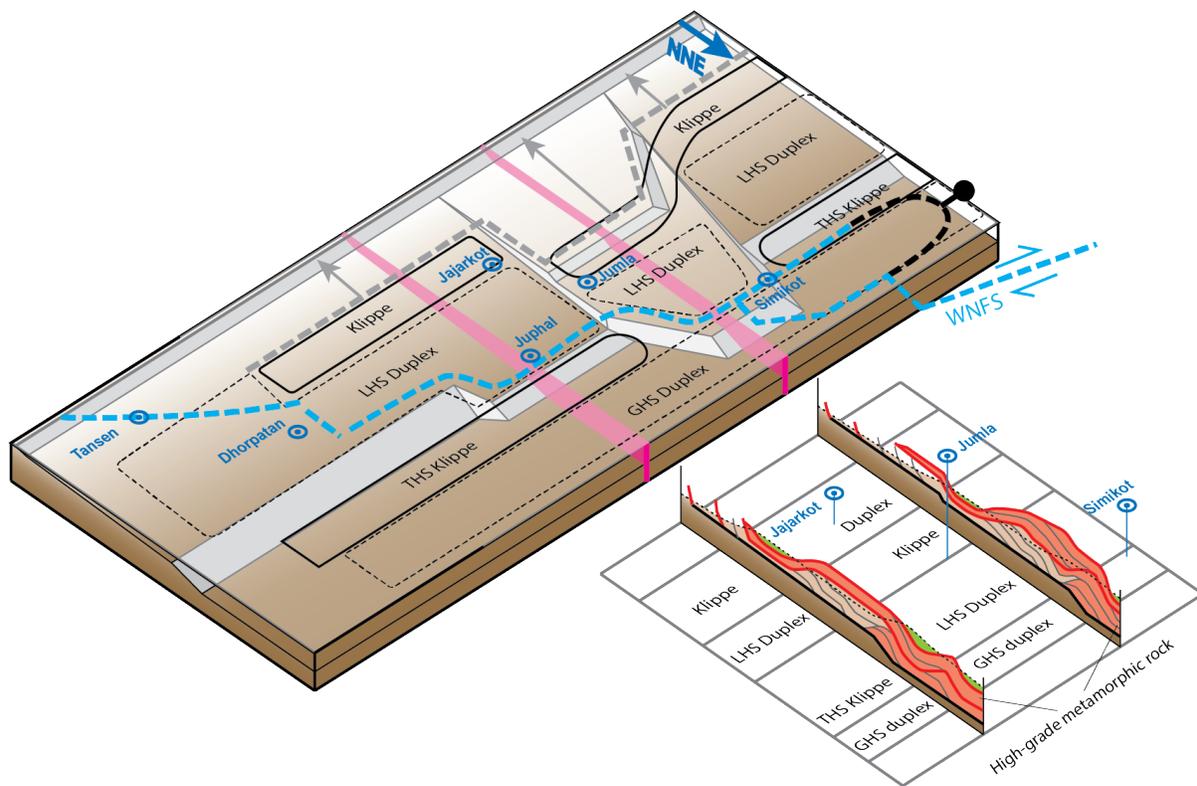
787 To assess the conceptual 3D MHT model for the mid-western Nepal Himalaya proposed
788 by Fan and Murphy (2021), three questions need to be addressed before assessing the details of
789 the geometry: First, is it true that the mid-lower crustal ramp always controls where the active-
790 uplift front location along the Himalaya? Second, can the high-slope zone represent the active-
791 uplift front of the orogen? Third, which high-slope zone among the two branches in western
792 Nepal Himalaya represent the active-uplift front of that segment?

793 The northern branch of the high-slope zone where the young cooling ages cluster along
794 the modeled transect, is above the upper part of the mid-lower crustal ramp and the southern
795 limb of the antiformal crustal accretion in our preferred models (Fig. 9). Along a transect across
796 the Kathmandu and Anapurna areas, the cooling ages have a similar spatial pattern showing the
797 youngest age cluster located in the high-slope zone is associated with the rapid exhumation
798 caused by crustal accretion on a mid-lower crustal ramp in good-fit thermokinematic models
799 (Herman et al., 2010; Ghoshal et al., 2020). This pattern is not affected by the ca. 60-70 km
800 difference in the distance between the MFT and the ramp or the active-uplift front between the
801 western and central Nepal Himalaya (e.g. Herman et al., 2010; Elliott et al., 2016; Hubbard et al.,
802 2016; Whipple et al., 2016; Ghoshal et al., 2020, and this study). This pattern is also found valid
803 in Kumaun, Annapurna, and Bhutan Himalaya (Whipp et al., 2007; C  lerier et al., 2009; Landry
804 et al., 2016). These observations support that the same mechanism may be controlling the
805 tectonomorphology of the whole orogen; the along-strike variation in the location of the active-
806 uplift front can therefore be interpreted as an indicator of the along-strike heterogeneity in the

807 location of the MHT mid-lower crustal ramp (Fan and Murphy, 2021). The newly reported ages
808 along the northern Himalaya antiform and close to the hinge of the Dolpo synform in western
809 Nepal are older (9.5-12.2 Ma) than those reported along-strike correspondent ages along the
810 modeled transect. In contrast, the ages to the south of the Dolpo synform, generally along the
811 eastern limb of the topographic embayment in mid-western Nepal, are of younger ages (3.1-7.4
812 Ma), and the youngest age of 3.1 Ma is consistent with the ZHe ages reported along the high
813 slope zone in other sectors of the orogen (Fig. 4). In the Thakkhola graben, published ages also
814 show that the youngest cooling ages are at the high-slope zone and the ages to the north in the
815 inner wedge are relatively old (Fig. 4). Although the southward younging trend of the cooling
816 ages along the footwall of the graben-bounding fault is originally thought to indicate the
817 southward development of the graben (Brubacher et al., 2020), we interpret it as a result of the
818 southward propagation of the crustal thickening for two reasons. First, the graben-bounding fault
819 is very steep and has only accommodated 2.2 km of wedge-parallel extension (Baltz, 2012).
820 Therefore, it is unlikely to significantly affect the thermal structure in the footwall. Second, the
821 cooling ages reported from the footwall of the graben-bounding fault by Brubacher et al. (2020)
822 are very consistent with our new data from the northern Himalaya antiform and from the hinge of
823 the Dolpo syncline, far away from the graben-bounding normal fault. These data support that the
824 active-uplift front of the transitional area along the embayment from far western Nepal to central
825 Nepal is represented by the northern branch of the high-slope zone (Fan and Murphy, 2021).

826 The details of the conceptual 3D kinematic evolution of the MHT geometry including the
827 ramps and how the MHT coevolved with the morphology of the orogen is described in Fan and
828 Murphy (2021). The 3D kinematics of the fold-thrust-belt in the upper-crustal part in this region
829 is also discussed in DeCelles et al. (2020). Below we integrate these discussions with the surface
830 geology and our modeling results to briefly introduce the main along-strike variations in the
831 MHT ramp geometry of our model for the western Nepal Himalaya (Fig. 10). Along the transect
832 across the apex of the embayment the MHT is characterized by two ramps of which the mid-
833 lower crustal ramp is located further to the hinterland compared with the adjacent segments. The
834 two ramps correlate with PT2-S and PT2-N. In the adjacent areas to the west and to the east of
835 this segment, the MHT is characterized by one mid-crustal ramp located closer to the MFT and it
836 connects up-dip to a deeper flat than the flat in our modeled transect. Oblique or lateral ramps in
837 the mid-lower crust accommodate the along-strike changes in the location of the mid-lower
838 crustal ramps. These oblique or lateral ramps correlate with the abrupt forelandward shift of the
839 MCT and the active-uplift front from the embayment sector to adjacent sectors, and with the
840 along-strike transition between antiformal structures (duplex) and synformal structures (klippe or
841 half klippe). If these lateral ramps started to develop as an older mid-lower crustal ramp below
842 the northern Himalaya anticline propagated towards the foreland differently along-strike as
843 suggested by (Fan and Murphy, 2021), it may have happened at ca. 10 Ma as suggested by the
844 cooling ages along the northern Himalayan anticline (this study and Brubacher et al., 2020). To
845 the east of the town of Juphal, a small lateral ramp may accommodate the further southward shift
846 of the mid-lower crustal ramp. This lateral ramp, though is not expressed in the klippe structure
847 to the south, can be correlated with the shift of the crest of the LHS duplex, the MCT, the STD,
848 and the active-uplift front. The general trend and major step-overs of the WNFS can be
849 correlated well with the general shape of the mid-lower crustal ramps and the lateral and oblique
850 ramps in this model. This supports that the WNFS might have developed partly by exploiting the
851 structures at depth as suggested by Fan and Murphy (2021). The present mid-lower crustal ramp
852 in this model also has a similar pattern to the reconstructed initial upper-crustal frontal ramp

853 producing the LHS duplex reported in DeCelles et al. (2020), supporting that this upper-crustal
 854 ramp may initially bifurcate from a once longer mid-lower crustal ramp as suggested by Fan and
 855 Murphy (2021).



856

Figure 10 Three-dimensional conceptual model of the MHT in western Nepal Himalaya with two simplified cross-sections show the along-strike variation in first-order structures. Refer to the text for the description and the correlation with main geological features.

857

858 6.3 Implications for Orogenic Wedge Growth

859 Our thermokinematic models employed a different strategy from previous similar studies
 860 in adjacent areas to simulate crustal accretion (e.g. Herman et al., 2010; Landry et al., 2016;
 861 Grujic et al., 2020). These previous studies used a wide window of crustal accretion at a constant
 862 location in the coordinate of the upper plate. However, the location of crustal accretion may
 863 change as the orogenic wedge grows. Crustal accretion or duplexing kinematically requires the
 864 formation of new crustal slices and their advection to the base of the orogen in the hinterland
 865 over successive cycles. Thermo-mechanical models show that each accretion cycle generates
 866 transient topographic growth above the crustal ramp which translates with the crustal ramp to the
 867 hinterland after their generation, resulting in a permanent increase in the orogen width (Mercier
 868 et al., 2017). Considering this kinematics, we used a narrower window of accretion than previous
 869 studies but let it advect to the hinterland in a cycle. In this design, the total one-cycle accretion, if
 870 plotted in an uplift profile along the transect, has an isosceles trapezoid shape, consistent with the
 871 common antiformal shape of a duplex (Fig. 5).

872 In the inversions, the parameters defining crustal accretion, including the location,
873 accretion rate, and the advection velocity, are independent from the parameters defining the
874 MHT geometry. However, in the best-fit models in the two accretion-involved inversions, the
875 crustal accretion process fit well with many observations. First, the total crustal accretion is in
876 the hinterland above the present mid-lower crustal ramp and it corresponds well with the
877 northern Himalayan anticline, which are consistent with the conceptual model proposed by Fan
878 and Murphy (2021). Second, the hinterland-ward advection of the instantaneous accretion
879 window relative to the matter in upper plate is consistent with the finding of the modeling work
880 in central Nepal Himalaya that to produce the observation-fitting ages a model requires the
881 recent up-dip advection of a duplex structure over a ramp (Ghoshal et al., 2020). Third, the high-
882 slope zone is on the southern flank of antiformal crustal accretion, which is consistent with the
883 geomorphological modeling results that the enhanced uplift caused by duplexing or advection
884 over a ramp can cause the highest river steepness in the forelimb of the uplift zone (Adams et al.,
885 2016; Eizenhöfer et al., 2019). Lastly, the uplift rates caused by the crustal accretion and rock
886 advection over the mid-lower crustal ramp in the preferred models are broadly consistent with
887 previous exhumation-rate studies at different locations in the Himalaya: The preferred models
888 have accretion velocity of ~ 5.2 mm/yr, adding the uplift caused by advection over mid-lower
889 crustal ramp resulting in a total uplift rate of ~ 6.1 mm/yr. The instantaneous velocity is
890 consistent with the observed short-term deformation or erosion, such as geodetic observations at
891 various locations (Jackson et al., 1992; Jackson and Bilham, 1994; Bilham et al., 1997; Grandin
892 et al., 2012); fluvial incision rates estimated from the modern channel geometries (Lavé and
893 Avouac, 2001); and erosion rates estimated from petrographic and mineralogical data of modern
894 river sediments (Garzanti et al., 2007). If considering the time span of one accretion cycle (6
895 Ma), average uplift rate can be determined by the accretion rate, width of the accretion window,
896 and advection rate of the accretion window in the model. In our preferred models, the temporally
897 average uplift rates of the crest part of the antiformal accretion are about 1.81 mm/yr (WNP02)
898 and 2.24 mm/yr (WNP03), which adding the uplift component caused by advection over the mid-
899 lower crustal ramp results in 2.73 mm/yr and 3.17 mm/yr respectively. They are also broadly
900 consistent with studies on long-term erosion rates, including the results from bedrock
901 thermochronology (Burbank et al., 1996; Burbank et al., 2003; Huntington et al., 2006; Blythe et
902 al., 2007; Adams et al., 2009; Thiede et al., 2009; Thiede and Ehlers, 2013), bedrock
903 thermochronology with thermokinematic modeling (Herman et al., 2010; McCallister et al.,
904 2014; Adams et al., 2015; Landry et al., 2016; Stübner et al., 2018), detrital thermochronology
905 (Brewer et al., 2006; Huntington and Hodges, 2006; Szulc et al., 2006; Copeland et al., 2015).
906 They are also broadly consistent with studies on current average erosion rate in catchment scale
907 determined by cosmogenic nuclides in quartz from river sediments (Vance et al., 2003; Scherler
908 et al., 2014; Ojha et al., 2019) and on geochemical mass-balance of erosion fluxes of modern
909 rivers (Galy and France-Lanord, 2001).

910 A modified critical taper model predicts that a wedge with brittle-ductile transition in
911 both the wedge and decollement parts can maintain a high slope zone connecting a taper-like
912 outer wedge and a plateau inner wedge at a critical state (Williams et al., 1994). But like the
913 original critical taper model that only considers frictional material (Davis et al., 1983; Dahlen et
914 al., 1984), this model does not address the internal deformation that maintains the wedge shape
915 and self-similar growing process. Many of the previous thermokinematic models highlighting the
916 role of crustal accretion in other Himalayan regions correlate the accretion in their models with
917 the LHS duplex in the footwall of the MCT. Different from these models, our models of the

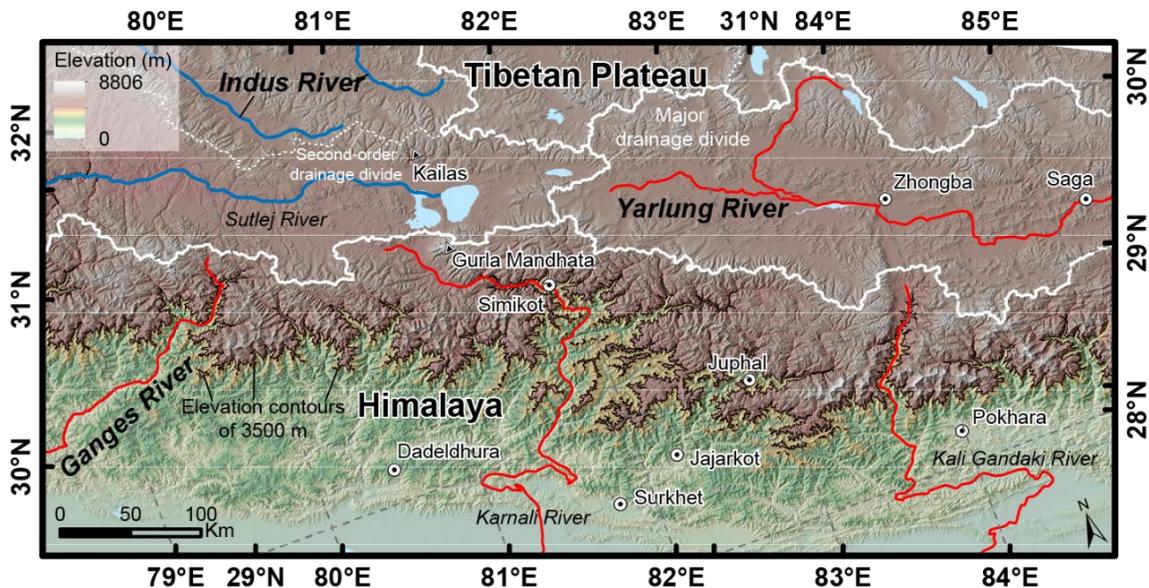
918 western Nepal Himalaya suggest the mid-lower crustal accretion causes the thickening the
919 northern Himalayan anticline (Fig. 9), where high-grade metamorphic rocks with protolith from
920 both the GHS and LHS are exposed (Murphy, 2007). However, to the south of the high-slope
921 zone or the active crustal accretion zone above the mid-lower crustal MHT ramp, the LHS
922 duplex is also well-developed and is interpreted to develop by the southward propagation of
923 crustal accretion on an upper-crustal flat (e.g. DeCelles et al., 2001; DeCelles et al., 2020) (Fig.
924 10). Recently, the LHS duplex is interpreted to be active in central Nepal based on the seismicity
925 data (Mendoza et al., 2019). Further to the foreland, between the MBT and MFT, the older
926 foreland basin strata group (SG) also exhibits duplexes (e.g. Mugnier et al., 1999). A recent
927 study at the rear side of the wedge suggests that thrust duplexing at depth is also active in South
928 Tibet (Taylor et al., 2021). These observations demonstrate that material accretion from the
929 lower plate to the upper plate occurs at different depths with a complex megathrust geometry; the
930 crustal accretion in the mid-lower crust, controlled by the brittle-ductile transition zone along the
931 MHT, may be determining the most active-uplift front and thus the edge of the plateau.

932 Crustal thickening via multi-layer duplexing is observed or proposed not only in Tibet-
933 Himalayan orogenic system, such as Kunlun range (Wang et al., 2011), Qilian Shan (Zuza et al.,
934 2018), northern Indo-Burma range (Haproff et al., 2020), Yarlung suture (Laskowski et al.,
935 2018), Lhasa terrane (Shi et al., 2020), but also in many other convergent wedges, such as the
936 Franciscan Complex (Platt, 1975), the Appalachian (Ando et al., 1984), and Alaska (Sample and
937 Fisher, 1986; Fuis et al., 1997; Moore et al., 1997; Wissinger et al., 1997; Fuis et al., 2008).
938 These studies provide detailed information about deformation styles and kinematics of wedge or
939 plateau growth. However, how rheologic change and material accretion affect the formation and
940 evolution of a mid-lower crustal ramp and the duplexes above it remains unclear. Due to the
941 limitation of the resolution and computing ability, numerical simulations at the lithospheric scale
942 usually ignore the complex ramp geometries and thus simulate accretion along a single orogen-
943 scale ramp which is seldomly truth in real life (e.g. Mercier et al., 2017; Menant et al., 2020).
944 Numerical and analogue simulations designed for investigating the kinematics of deformation
945 and the evolution of megathrust ramps usually ignore the rheology change along the megathrust
946 and the lower plate deformation as it subducts (e.g. Mugnier et al., 1997; Malavieille, 2010; Dal
947 Zilio et al., 2020b; Ghosh et al., 2020; Shen et al., 2020). These models usually require the
948 preexistence of weak layers in the lower plate to generate duplexes, and the weak layers usually
949 behave as décollements after being involved in the wedge deformation. In these models, once a
950 ramp forms it can only underthrust to the hinterland without deformation in the footwall till a
951 new ramp forms in the foreland. This kinematics may be consistent with the upper-crustal
952 structures but are not true for the evolution of the mid-lower crustal ramps. The formation and
953 sustaining of a mid-lower crustal ramp have been conceptually explained by flexural depression
954 of the lower plate driven by gravity load of the orogen or plateau (Coward, 1983), and
955 mechanical models suggest the strength of the underthrusting lithosphere affects the geometry of
956 the fault (Lyon-Caen and Molnar, 1985; Molnar and Lyon-Caen, 1988).

957 6.4 Effects of Crustal Accretion on the Landscape of the Himalayan Orogen

958 The growth of the inner wedge caused by mid-lower crustal accretion may have
959 significantly affected the drainage systems and landscape of the southern edge of the Tibetan
960 plateau. The present landscape of the Himalaya, from north to south, is characterized by the
961 longitudinal Indus and Yarlung river systems, a plateau of high-elevation and low-relief, and an

962 outer wedge of high-relief landscape, and the extent of these features varies along strike (Fig.
 963 11). The Indus and Yarlung drainage systems are separated from drainage systems within the
 964 Tibetan plateau to the north and the outer Himalayan wedge to the south by drainage divides
 965 (Fig. 11). Moreover, in the Late Oligocene-Early Miocene, the landscape between the Gangdese
 966 range in south Tibet and the Himalaya to the south was characterized by extensive lakes rather
 967 than by longitudinal river systems (e.g. DeCelles et al., 2018), and trans-Himalaya rivers existed
 968 bringing sediment from the Gangdese range to the Himalayan foreland in the Miocene (e.g. Cina
 969 et al., 2009; Lang and Huntington, 2014).



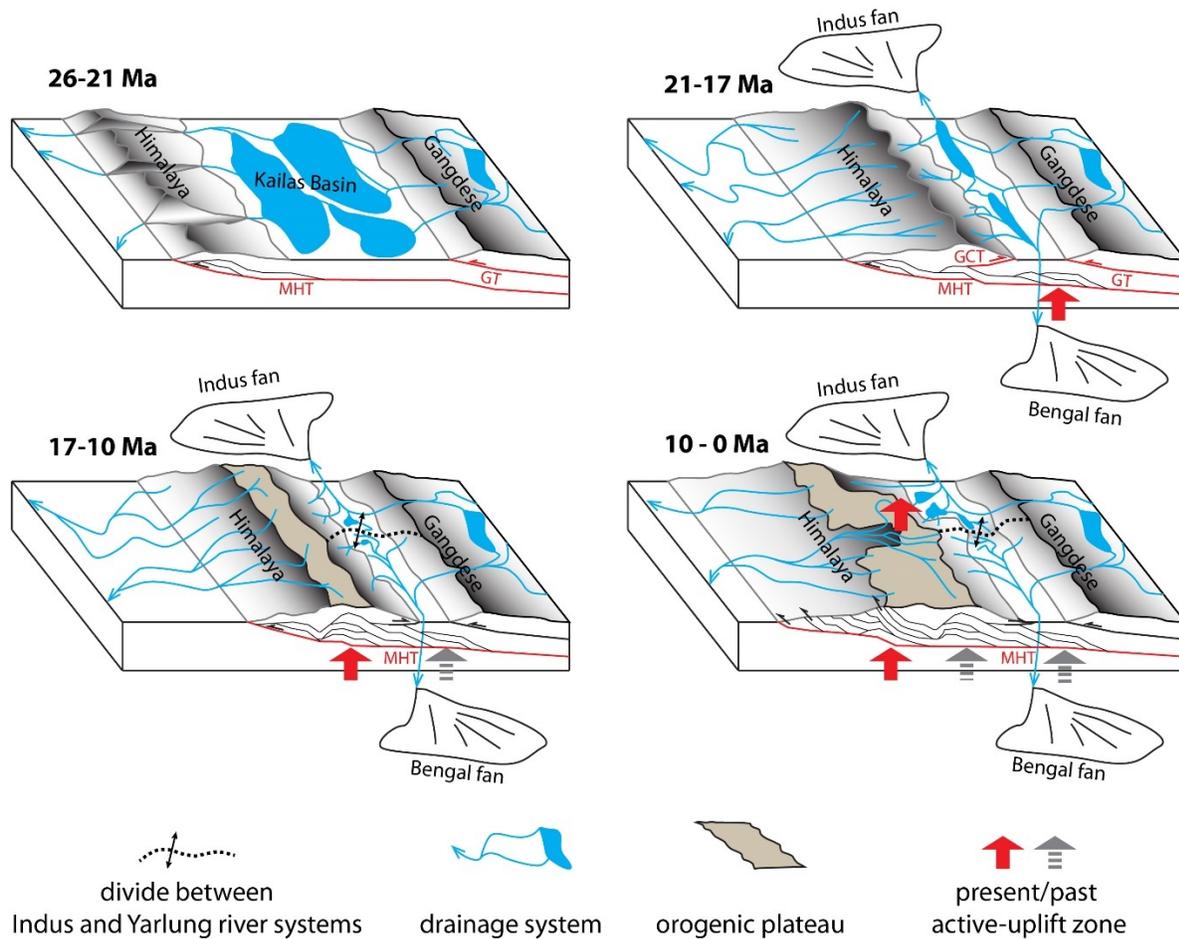
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Figure 11 Map shows the topography, landform, and the drainage systems of the central-western Nepal Himalayan and adjacent areas. The blue and red lines represent rivers that flow to Indus fan and Bangel fan respectively. Notice the along-strike change in the width of the orogenic plateau landscape, approximately the area between the 3500 m elevation contours and the southern divides of Indus and Yarlung river systems, and the major drainage divide between the Indus river system and Yarlung river system to the north of the topographic embayment in Simikot segment.

971

972 On the northern side of the Himalaya range, the formation of the longitudinal Indus and
 973 Yarlung drainage systems may be caused by uplift of the northern Himalaya and shortening on
 974 both sides of the present drainage system. On the northern side of the Himalaya along the Indus-
 975 Yarlung suture zone, low-altitude warm-water great lakes existed during Oligocene-Miocene
 976 time (26-21 Ma) (DeCelles et al., 2018). Early deposition of the Kailas Formation during this
 977 period is interpreted to be caused by regional extension as the subducting slab rolls back
 978 (DeCelles et al., 2011; Carrapa et al., 2014), and the detrital sources are mostly the Gangdese
 979 range to its north. Late stages of the deposition recorded the addition of the THS source from its
 980 south. The change in source is interpreted to be caused by the activation of the north-directed
 981 Great Counter thrust at the rear of the Himalaya range. Cooling ages of the Kailas basin suggest
 982 post-depositional exhumation at around 17 ± 1 Ma (Carrapa et al., 2014), and is interpreted to
 983 reflect the initial rapid incision of the longitudinal river systems. The northern Himalaya
 984 anticline may have developed coevally due to mid-lower crustal accretion and ceased in some
 985 segment at ca. 10 Ma as our data suggest. Along the suture, studies also suggest crustal duplex at

986 depth developed from ca. 21-23 Ma (Laskowski et al., 2018). Therefore, we posit that the crustal
 987 accretion may have facilitated the development of the river systems (Fig. 12). North-south
 988 shortening along the Great Counter thrust and the Gangdese thrust made the area between the
 989 Gangdese range and Himalaya range narrow and therefore facilitated localization of the drainage
 990 systems. Subsequent crustal accretion along the suture and later the Northern Himalaya anticline,
 991 and probably also slab detachment of the subducting plate, caused elevation gain in the area and
 992 rapid incision of the Indus and Yarlung river systems. Development of the northern Himalayan
 993 anticline may have also terminated some the trans-Himalaya rivers presently represented by wind
 994 gaps along the drainage divide between the Himalaya plateau (inner wedge) and the two
 995 longitudinal river systems (e.g. Murphy and Burgess, 2006; Buceta et al., 2020).



996

Figure 12 Conceptual evolution model of the drainage system and landscape of the area across the western Nepal Himalaya and adjacent areas.

997

998 The Indus and Yarlung river systems extends approximately parallel with the Himalaya at
 999 the rear with the Indus river system draining the western part to the west to the Indus fan and the
 1000 Yarlung river system draining the central-eastern segment to the east to the Bengal fan. The
 1001 divide between the two river-system catchments is to the north of the topographic embayment in
 1002 western Nepal (Fig. 11). This raises two questions; how did this divide form and when did it

1003 evolve to its present location? Yin (2006) originally proposed that this drainage divide may have
1004 been controlled by subduction of basement ridges in the Indian plate and it might have shifted
1005 from ridge to ridge to its current position by headwater erosion. Given that the topographic
1006 embayment developed on a structurally high MHT segment which might have pushed the
1007 hinterland mid-lower crustal accretion further to the north compared along-strike (this study and
1008 Fan and Murphy, 2021), we propose a modified basement-ridge-control model in which the
1009 divide may be controlled by this structurally high segment of the MHT, not necessarily the large
1010 basement ridges. If this is correct, the divide should have established its present location by ca.
1011 10 Ma, when the mid-lower crustal lateral ramps of the MHT in the embayment area developed.
1012 If the drainage systems were affected by the wedge-parallel extension that initiated at ca. 15 Ma
1013 in the Gurla Mandhata and Xiao Gurla areas (McCallister et al., 2014), the divide could start to
1014 develop at its present location as early as then. This suggests that the divide between the Indus
1015 river system and the Yarlung river system may have established its present location
1016 approximately right after the river systems formed or localized (Fig. 12).

1017 To the south of the Indus and Yarlung drainage systems, across a drainage divide, the
1018 growth of the Himalayan plateau landscape might be controlled by the evolution of mid-lower
1019 crustal accretion. In western Nepal, the drainage divide between Yarlung-Brahmputra drainage
1020 system and the Himalayan plateau approximately follows the northern Himalaya anticline and
1021 the southern edge of the plateau landscape is approximately along the northern limit of the high-
1022 slope zone which varies in its location along strike (Fig. 11). Therefore, the width of the inner
1023 wedge also varies along the strike. In far-western Nepal along the apex of the topographic
1024 embayment, the plateau part is narrow, while in the segment from Dolpo area to the Thakkhola
1025 graben, approximately corresponding to the Dolpo THS syncline, the plateau landscape is
1026 relatively wide (Fig. 11). If, as discussed in this study and Fan and Murphy (2021), the active-
1027 uplift front caused by mid-lower crustal accretion migrated from the northern Himalaya anticline
1028 to its present location as lateral or oblique ramps in MHT at mid-lower crustal depth developed,
1029 the inner wedge or the plateau landscape might have expanded at the same time with along-strike
1030 variation. Two factors could contribute to this process; (1) the southward migration of the mid-
1031 lower crustal accretion contributed to build the topography; (2) due to the orographic
1032 precipitation effect (Bookhagen and Burbank, 2006), the aridification to the north of the active-
1033 uplift front could facilitate the preservation of the plateau landscape built by earlier mid-lower
1034 crustal accretion. The coevolution of the mid-lower crustal accretion and plateau expansion is
1035 similar to the well-studied forming process of intermontane basins or piggyback basins resulted
1036 from the foreland-ward propagation of deformation in fold-and-thrust belts (e.g. Sobel et al.,
1037 2003; Hilley and Strecker, 2005; Coutand et al., 2006; Ruetenik et al., 2018; Pingel et al., 2020).
1038 This proposed process is also supported by the landscape evolution modeling of the forming
1039 process of high-elevation, low-relief landscape patches in the lower Himalaya area via duplex
1040 deformation at depth (Adams et al., 2016). The history of the trans-Himalaya rivers that once
1041 transported detritus from Gangdese to the Himalayan foreland is not well known, but these rivers
1042 could have been blocked as the northern Himalayan anticline and the plateau landscape
1043 developed.

1044 7 Conclusion

1045 Our thermokinematic models suggest that the MHT along the Karnali river transect is
1046 characterized by two ramps connected by a long upper-crustal flat and a mid-lower crustal ramp

1047 in the hinterland farther to the north with a gentler slope compared with the mid-lower crustal
1048 ramps in adjacent segments. The modeling results also suggest that the best-fit models require
1049 crustal accretion on the mid-lower crustal ramp to successfully produce observation-fitting
1050 cooling ages of MsAr, AFT, ZHe, and AHe dating systems. The youngest cluster of the cooling
1051 ages and the high-slope zone of the orogenic wedge are above the forelimb of the antiformal
1052 crustal accretion, which is coincidence with the northern Himalaya anticline in far-western Nepal
1053 in our model. These findings suggest that the crustal accretion along the mid-lower crustal ramp
1054 is controlling the location of the active-uplift front of the orogenic wedge and the migration of
1055 the location controls the outward growth of an orogenic plateau. The coexistence of the duplex
1056 structures in the upper-crustal depth with the mid-lower crustal accretion we modeled suggests
1057 that crustal accretion at different depths or multi-layer duplexing is an important mechanism of
1058 maintaining the wedge shape predicted by the critical taper theory.

1059 The ZHe ages reported here from western Nepal Himalaya support that the northern
1060 branch of the high-slope zone or the PT2-N is the active-uplift front of the plateau in the area and
1061 that its along-strike change in its location is an expression of the along-strike change in MHT
1062 geometry. The lateral and oblique ramps in mid-lower crust of a 3-D MHT geometric model for
1063 the western Nepal Himalaya on the base of this can be correlated well with the surficial
1064 geological features. Cooling ages in this region also suggest that the mid-lower crustal lateral and
1065 oblique ramps in MHT started to develop at least by ca. 10 Ma, when the crustal accretion
1066 building the northern Himalaya anticline ceased in the segment between the modeled transect
1067 and the Thakkhola graben and started to migrate southward to its present location. Integrating
1068 other geologic data across the Himalaya, we propose that the deep tectonic process may have
1069 controlled the first-order evolution of the landscape and drainage systems, including the
1070 transition from the landscape featured by low-elevation big lakes to the localized river systems,
1071 the cessation of the trans-Himalaya rivers, the development of the drainage divide between the
1072 Indus river system and the Yarlung river system, the southward growth of the inner-wedge
1073 plateau landscape with along-strike variations, and the formation of the topographic embayment
1074 in western Nepal.

1075 **Acknowledgements**

1076 This study was funded by National Science Foundation grant (no. 1827863) to Michael
1077 Murphy. The authors acknowledge the use of the Sabine Cluster and the advanced support from
1078 the Research Computing Data Core at the University of Houston to carry out the research
1079 presented here.

1080

1081 *References cited in the figure captions are included in the references.*

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