

1 **Observed diurnal cycles of near-surface shear and stratification in the**  
2 **equatorial Atlantic and their wind dependence**

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14 **Key Points:**

- 15 • In the upper 15 m of the equatorial Atlantic Ocean, a strong diurnal cycle of velocity  
16 differences of more than 10 cm s<sup>-1</sup> is observed.
- 17 • Wind speed controls amplitude and timing of the diurnal cycles of shear and stratification  
18 in the upper 20 m of the equatorial Atlantic.
- 19 • Wind speed dependence of descent rates of diurnal shear and stratification can explain  
20 the varying onset of deep-cycle turbulence.

## 21 **Abstract**

22 The diurnal cycles of near-surface shear and stratification, also known as diurnal jet and diurnal  
23 warm layer (DWL), are ubiquitous in the tropical oceans, affecting the heat and momentum  
24 budget of the ocean surface layer, air-sea interactions, and vertical mixing. Here, we analyse the  
25 presence and descent of near-surface diurnal shear and stratification in the upper 20 m of the  
26 equatorial Atlantic as a function of wind speed using ocean current velocity and hydrographic  
27 data taken during two trans-Atlantic cruises along the equator in autumn 2019 and spring 2022,  
28 data from three types of surface drifters, and data from PIRATA moorings along the equator.  
29 The observations during two seasons with similar wind speeds but varying net surface heat  
30 fluxes reveal similar diurnal jets with an amplitude of about  $0.11 \text{ m s}^{-1}$  and similar DWLs when  
31 averaging along the equator. We find that higher wind speeds lead to earlier diurnal peaks,  
32 deeper penetration depths, and faster descent rates of DWL and diurnal jet. While the diurnal  
33 amplitude of shear is maximum for intermediate wind speeds, the diurnal amplitude of  
34 stratification is maximum for minimal wind speeds. The presented wind dependence of the  
35 descent rates of DWL and diurnal jet is consistent with the earlier onset of deep-cycle turbulence  
36 for higher wind speeds. The DWL and the diurnal jet not only trigger deep-cycle turbulence but  
37 are also observed to modify the wind power input and thus the amount of energy available for  
38 mixing.

## 39 **Plain Language Summary**

40 Variations in solar radiation over the course of the day cause a diurnal cycle of temperature,  
41 stratification, current velocity, and velocity shear in the near-surface ocean. These diurnal cycles  
42 are ubiquitous in the tropical oceans and are important for understanding the heat and momentum  
43 budget of the ocean surface layer and for understanding vertical mixing. Here, we analyse the  
44 diurnal cycles of stratification and velocity shear in the upper 20 m of the equatorial Atlantic,  
45 focussing on their presence, depth structure, and wind dependence. We use data taken during two  
46 trans-Atlantic cruises along the equator in autumn 2019 and spring 2022, data from surface  
47 drifters, and data from mooring sites along the equator. These observations indicate that the wind  
48 speed influences the amplitude, timing, and vertical structure of the diurnal cycles. The wind  
49 speed dependence of the depth propagation of the diurnal cycles of stratification and velocity  
50 shear is consistent with the wind speed dependence of mixing below the mixed layer. We further

51 show that the diurnal cycle of near-surface current velocities also leads to a diurnal cycle of the  
52 amount of wind energy released into the ocean.

## 53 **1 Introduction**

54 Oceanic parameters vary close to the surface with the diurnal cycle of solar radiation,  
55 including stratification, shear, and mixing. After sunrise, a diurnal (buoyantly isolated) warm  
56 layer (DWL) forms that traps heat and also wind-forced momentum close to the surface, creating  
57 a highly-sheared near-surface diurnal jet due to the ‘slippery layer’ effect. The stratified shear  
58 layer descends during late afternoon and evening, transmitting heat and momentum below the  
59 mixed layer into the deeper ocean. After sunset, cooling of the sea surface and convective  
60 overturning destroy the DWL (Kudryavtsev & Soloviev, 1990; Price et al., 1986; Smyth et al.,  
61 2013; Woods & Strass, 1986). This diurnal variability of shear and stratification is linked to  
62 diurnal variability of turbulent dissipation both within the DWL (St. Laurent & Merrifield, 2017;  
63 Sutherland et al., 2016) and below (Moum et al., 2022; Peters et al., 1988). Therefore, the near-  
64 surface diurnal cycle modifies near-surface heat and momentum budgets and plays a role in air-  
65 sea interactions and vertical mixing.

66 Understanding the diurnal cycle is of particular interest in equatorial regions for several  
67 reasons:

- 68 1. The equatorial Atlantic and Pacific are characterised by a zonal current system and a  
69 highly-sheared Equatorial Undercurrent (EUC), eventually leading to the presence of  
70 marginal instability. Marginal instability is defined as a state in which shear and  
71 stratification vary almost proportionally so that the Richardson number remains close to  
72 its critical value (Smyth & Moum, 2013; Smyth et al., 2019). It is observed in the  
73 equatorial Pacific that, when the descending diurnal shear layer merges with the  
74 marginally unstable shear above the EUC core, shear instabilities are induced that can  
75 trigger turbulence, the so-called deep-cycle turbulence (DCT) (Smyth & Moum, 2013;  
76 Smyth et al., 2013; Pham et al., 2013). DCT also occurs in the equatorial Atlantic, but it  
77 is still an open question whether there are fundamental differences in the nature of  
78 instabilities leading to DCT in the Atlantic and Pacific (Moum et al., 2023).

- 79        2. Diurnal jet and DWL are observed to reach and thus impact the ocean far deeper near the  
80        equator than away from it (Masich et al., 2021). This discrepancy arises due to a  
81        combination of Coriolis rotational effects that are vanishing towards the equator disabling  
82        the rotation of horizontal velocities with depth (Hughes et al., 2020a) and the presence of  
83        very high background shear near the equator supporting the descent of the shear layer  
84        (e.g., Lien et al., 1995). Note that longitudinal differences in background shear and the  
85        presence of marginal instability can also lead to longitudinal differences in the  
86        penetration depth of the diurnal jet (Masich et al., 2021).
- 87        3. Equatorial cold tongue regions are critical for the global heat balance and the near-  
88        surface diurnal cycle presents a key mechanism there for the heat uptake from the  
89        atmosphere to the stratified ocean below the surface mixed layer (Moum et al., 2013;  
90        Whitt et al., 2022). Upwelling and mixing in these regions define not only the downward  
91        heat flux but similarly the upward nitrate flux (Radenac et al., 2020; Brandt et al., 2023),  
92        stressing the importance of diurnal variability at the equator also for biological  
93        productivity.
- 94        4. It is expected that in the tropics and subtropics, which often present a larger partial  
95        pressure of CO<sub>2</sub> in the ocean than in the atmosphere, diurnal variability of turbulence  
96        within the DWL increases the flux of CO<sub>2</sub> from the ocean to the atmosphere (Sutherland  
97        et al., 2016).

98        Wind speed influences the formation and the pattern of the DWL and the diurnal jet as  
99        indicated by observational (Hughes et al., 2020b; Wenegrat & McPhaden, 2015; Masich et al.,  
100        2021) and modelling studies (Hughes et al., 2020a, 2021), where the diurnal cycle of the wind  
101        itself can be neglected as it is at least one order of magnitude smaller than the daily-mean wind  
102        signal magnitude (Masich et al., 2021; Smyth et al., 2013). It has been suggested that DWLs and  
103        diurnal jets do not exist for wind speeds exceeding a threshold ranging between 6 m s<sup>-1</sup> and 8 m  
104        s<sup>-1</sup> (Hughes et al., 2021; Kudryavtsev & Soloviev, 1990; Matthews et al., 2014; Thompson et al.,  
105        2019). Furthermore, Wenegrat & McPhaden (2015) observed a seasonal variability in the  
106        equatorial Atlantic with pronounced descending diurnal shear layers and limited diurnal sea  
107        surface temperature variability in steady trade wind conditions during boreal summer and  
108        autumn, and opposite patterns in weak wind conditions during boreal winter and spring. More

109 comprehensive analyses of the interaction between the wind and DWL and diurnal jet have been  
110 performed in the tropical Pacific. For higher wind speeds, the penetration depth of both DWL  
111 and diurnal jet becomes deeper (Hughes et al., 2020b; Masich et al., 2021; Price et al., 1986), and  
112 the descent rate of the DWL increases (Hughes et al., 2020b). However, little is known about the  
113 diurnal amplitudes as a function of wind speed. Masich et al. (2021) found a linear relationship  
114 between the wind speed and the strength of the diurnal cycle of current velocities at locations  
115 where marginal instability was present. Price et al. (1986) suggested that the diurnal jet  
116 amplitude is solely dependent on the net surface heat flux and followingly independent of the  
117 wind speed. Hence, there is a lack of a comprehensive analysis of the interplay between wind  
118 and diurnal jet regarding descent rates and diurnal amplitudes as well as a lack of a confirmation  
119 of the processes observed in the tropical Pacific for the tropical Atlantic.

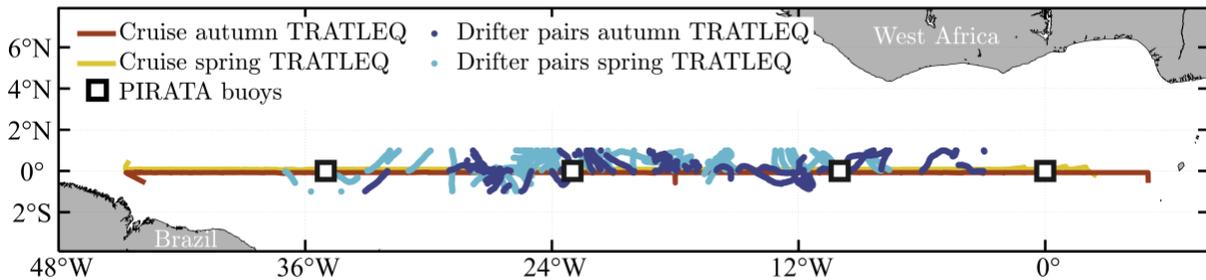
120 Measurements of ocean currents in the upper 10 m are still rare because of measurement  
121 constraints and noise, e.g., shipboard measurements with acoustic Doppler current profilers  
122 (ADCP) typically cover a depth range below 15 m depth and moored measurements with upward  
123 looking ADCPs are contaminated by interference with surface reflections or aggregation of fish  
124 (Röhrs et al., 2021; Elipot & Wenegrat, 2021). However, near-surface estimates within the upper  
125 10 m are necessary to properly capture the diurnal dynamics. Only few studies provide  
126 observational estimates within the upper 10 m at diurnal time scales. These studies are located in  
127 the tropics to subtropics and are based on different types of surface drifters (Kudryavtsev &  
128 Soloviev, 1990), on current meters and/or ADCPs attached to a mooring or surface buoy (Price  
129 et al., 1986; Cronin & Kessler, 2009; Wenegrat & McPhaden, 2015; Sutherland et al., 2016;  
130 Pham et al., 2017), or on a SurfOtter (Hughes et al., 2020a). The resulting diurnal jet amplitudes  
131 vary from  $10 \text{ cm s}^{-1}$  to  $20 \text{ cm s}^{-1}$  with different associated depth intervals, different locations as  
132 well as different seasons and prevailing background conditions. The vertical structure of near-  
133 surface shear and factors influencing the diurnal jet are still poorly understood.

134 This study combines observational data sets from the TRATLEQ expeditions, which are  
135 two trans-Atlantic equatorial cruises with dedicated en-route measurements and drifter  
136 deployments, and data sets from specially-instrumented PIRATA moorings to capture diurnal  
137 stratification and shear in the upper 15 m of the equatorial Atlantic Ocean. With these data sets,  
138 we aim to assess near-surface diurnal dynamics focussing on the influence of background  
139 conditions and, in particular, on the wind speed dependency. Followingly, the study addresses

140 the lack of near-surface measurements and the knowledge deficit about the wind dependency of  
 141 diurnal jet and the DWL in the equatorial Atlantic. The paper is organised as follows. Data and  
 142 methodology are described in sections 2 and 3, respectively. Results about the near-surface  
 143 diurnal cycle in the equatorial Atlantic and impacts of different background conditions are  
 144 presented in section 4. The impact of the wind speed on diurnal shear and stratification is  
 145 examined in more detail in section 5. The results are then discussed in terms of descent rates of  
 146 diurnal shear and stratification and are linked to the wind dependence of DCT found by Moum et  
 147 al. (2023) in section 6.1. The impact of the described diurnal cycles on the wind power input  
 148 (WPI) is discussed in section 6.2.

## 149 2 Observational data

150 This study focusses on different observational data sets from the TRATLEQ expeditions,  
 151 consisting of two research cruises and associated surface drifter experiments, and data sets from  
 152 PIRATA moorings (Figure 1).



153  
 154 *Figure 1. Geographical map of the observations used in the present study. Displayed are mean positions of drifter*  
 155 *pairs within 1° north and south of the equator deployed during TRATLEQ I in autumn 2019 (dark blue dots) and*  
 156 *TRATLEQ II in spring 2022 (light blue dots), the equatorial transects of the autumn (red line) and spring (yellow*  
 157 *line) TRATLEQ cruises, and the locations of the PIRATA buoys (black-bordered squares).*

### 158 2.1 TRATLEQ Cruises

159 Shipboard measurements were carried out during the cruises M158 and M181 with the  
 160 research vessel Meteor, the so-called TRATLEQ (Trans-Atlantic Equatorial) cruises I & II. The  
 161 cruises provide equatorial Atlantic transects from 5°E to 45°W between September 29 and  
 162 October 22, 2019 and from 2°E to 45°W between April 30 and May 20, 2022. In the following,  
 163 TRATLEQ I will be termed (boreal) autumn TRATLEQ and TRATLEQ II (boreal) spring  
 164 TRATLEQ. During both cruises, near-surface stratification is estimated from 10 s sea surface

165 temperature and sea surface salinity measurements by the ship's dual thermosalinograph (TSG)  
166 as well as 10 s pitch and roll data from the ship (more details in section 3.2). The vessel  
167 measures the speed and direction of the wind at 30 m height as well as global short-wave  
168 radiation (SWR) with a temporal resolution of 1 min. Direct shipboard velocity measurements  
169 from a marine radar and a vessel-mounted ADCP (vmADCP) are only considered for autumn  
170 TRATLEQ. A coherent-on-receive marine X-band (9.4 GHz) radar developed at the Helmholtz-  
171 Zentrum Hereon (Horstmann et al., 2021) was installed during autumn TRATLEQ. The  
172 instrument was set to operate at a pulse length of 50 ns (i.e., short-pulse mode), providing a  
173 range resolution of 7.5 m. It is equipped with a vertical transmit and receive (VV) polarised  
174 antenna of 2.3 m (7.5 ft) with a beam width of  $1.1^\circ$ , a rotational period of 2 s and a pulse  
175 repetition frequency of 2 kHz. The obtained image sequences are analysed with respect to the  
176 surface wave properties such as wave directions, wave lengths, and phase velocities. The surface  
177 current vector is then resulting from the difference between the observed phase velocities and the  
178 phase velocities given by the linear dispersion relation of surface gravity waves (Horstmann et  
179 al., 2015; Lund et al., 2018). The retrieved surface current layer varies between 1 m and 5 m  
180 depth, depending on the surface wave length. Here, a mean depth of 3 m is assumed for the  
181 marine Radar measurements. A validation study in the Gulf of Mexico showed a root-mean-  
182 square error of  $4 \text{ cm s}^{-1}$  compared to velocities of surface drifters representing the upper 0.4 m  
183 depth (Lund et al., 2018). There are no data between  $18^\circ\text{W}$  and  $25^\circ\text{W}$ , that is, from October 09 to  
184 October 12, 2019 and only data between  $0^\circ 0.6' \text{ S}$  and  $0^\circ 0.6' \text{ N}$  are considered. In addition to the  
185 marine radar, a vmADCP, a 75 kHz RDI Ocean Surveyor, was installed during autumn  
186 TRATLEQ with the bin size set to 8 m (Brandt et al., 2022). Here, only data from the uppermost  
187 bin centred at 17 m depth are considered. Hourly velocity data from the vmADCP have an  
188 accuracy of  $1 \text{ cm s}^{-1}$  for on-station and  $2\text{-}4 \text{ cm s}^{-1}$  for underway measurements depending on  
189 wave and wind conditions (J. Fischer et al., 2003). For comparison of marine radar and  
190 vmADCP data, 10 min averages were calculated. Outliers of the velocity differences between the  
191 two data sets were determined using a criterion of three standard deviations off the median and  
192 were eliminated. Direct shipboard velocity measurements are not considered for spring  
193 TRATLEQ. Due to a malfunction of the OS75kHz system (vmADCP), a 75kHz LR was installed  
194 in the sea chest during that cruise, which had a reduced signal-to-noise ratio for the uppermost  
195 bin, leading to a distortion of the diurnal cycle.

## 196           **2.2 TRATLEQ drifter experiments**

197           During both TRATLEQ cruises, drifter experiments were carried out that consisted of the  
198 pairwise deployments of two types of surface drifters about every 1° longitude along the equator.  
199 For autumn TRATLEQ, 31 SVP (Surface Velocity Program) drifters, drifting with velocities at  
200 about 15 m depth, and 27 CARTHE (Consortium for Advanced Research on Transport of  
201 Hydrocarbon in the Environment) drifters, providing velocities at about 0.5 m depth, were  
202 deployed between September 29 and October 18, 2019. For spring TRATLEQ, 18 SVP drifters  
203 and 44 Hereon drifters, which are similar to CARTHE drifters and provide velocities at about 0.5  
204 m depth, were deployed between May 04 and May 17, 2022. Both trajectory data sets were  
205 quality-controlled and interpolated to hourly values. Estimates of the velocity difference between  
206 0.5 m and 15 m were then based on drifter pairs that are separated in time by less than 1 hour and  
207 in distance by less than 100 km (details in Text S1 and Figure S1). This study considers 7633  
208 drifter pairs between October 02 and October 29, 2019 in the area from 33°W to 3°W and 1°S to  
209 1°N as well as 9602 drifter pairs between May 04 and June 02, 2022 in the area from 37°W to  
210 8°W and 1°S to 1°N. The mean distance of the paired drifters is 46 km for autumn TRATLEQ  
211 and 54 km for spring TRATLEQ.

## 212           **2.3 PIRATA moorings**

213           Near-surface temperature and salinity data from the PIRATA (Prediction and Research  
214 Moored Array in the Tropical Atlantic) buoys at 0°N, 23°W and 0°N, 10°W are used. Moorings  
215 were progressively equipped with temperature and conductivity sensors at 1 m, 5 m, 10 m, 20 m,  
216 and 40 m depth from 1999 to 2022. Wind data at 4 m height are taken from the PIRATA sites  
217 0°N, 0°W (1999 - 2022), 0°N, 10°W (1999 - 2022), 0°N, 23°W (1999 - 2022), and 0°N, 35°W  
218 (1998 - 2022) (Bourlès et al., 2019). The net surface heat flux at 0°N, 10°W and 0°N, 23°W is  
219 estimated as the sum of SWR, long-wave radiation, latent, and sensible heat flux provided by  
220 ePIRATA (Foltz et al., 2018) for the two TRATLEQ periods. Additionally, a Teledyne-RDI  
221 Sentinel Workhorse 600 kHz ADCP was deployed at 0°N, 23°W from October 13, 2008 until  
222 June 18, 2009, providing hourly velocity averages from a depth of 4.3 m to 38.8 m (see  
223 Wenegrat et al., 2014). The data set is masked according to  
224 ([https://www.pmel.noaa.gov/tao/drupal/disdela/adcp\\_0n23w/index.html](https://www.pmel.noaa.gov/tao/drupal/disdela/adcp_0n23w/index.html)). This period of near-

225 surface high vertical resolution moored velocity measurements will be referred to as enhanced  
 226 monitoring period (EMP).

## 227 **2.4 Satellite wind data**

228 Winds at 10 m height are taken from the gridded 6-hourly Cross-calibrated Multi-  
 229 Platform (CCMP) near-real time wind satellite product provided by Remote Sensing Systems for  
 230 the period January 2000 to November 2022. The product is also used to estimate the wind at the  
 231 drifter locations. The CCMP V2.0 product is processed to L3 standard, has a horizontal  
 232 resolution of  $0.25^\circ \times 0.25^\circ$  and a temporal resolution of 6 h. It is averaged daily for the following  
 233 analysis.

## 234 **3 Methods**

### 235 **3.1 Stratification from the PIRATA moorings**

236 For the PIRATA moorings, temperature and salinity data are given on regular pressure  
 237 and time grids. The stratification,  $N^2$ , is given as squared Brunt-Väisälä frequency and can be  
 238 calculated according to IOC et al. (2010) as

$$239 \quad N^2 = g^2 * \rho * \frac{\beta * \Delta S_A - \alpha * \Delta \theta}{\Delta P} \quad (1)$$

240 where  $\theta$  is the conservative temperature,  $S_A$  the absolute salinity,  $\rho$  the in-situ density,  $g$   
 241 the gravitational acceleration,  $\alpha$  and  $\beta$  the coefficients of thermal expansion and saline  
 242 contraction, respectively, and  $P$  the pressure in Pa. The respective parameters are computed using  
 243 the Gibbs SeaWater Oceanographic Toolbox of TEOS-10.

### 244 **3.2 Stratification from the vessel-mounted thermosalinograph**

245 The stratification,  $N^2$ , at the depth of the TSG inlet can be estimated using data taken at a  
 246 high sampling rate (here  $0.1 \text{ s}^{-1}$ ) for temperature, salinity, and the vertical movement of the inlet  
 247 position relative to the water column. This method was first described in T. Fischer et al. (2019).  
 248 The vertical distance of the inlet relative to the mean sea level is evaluated as

$$249 \quad d_{inlet,sealevel} \approx (y_{inlet,com} * \sin(\psi) - z_{inlet,com} * \cos(\psi)) * \cos(\gamma) - x_{inlet,com} * \sin(\gamma) +$$

$$250 \quad d_{com,sealevel} \quad (2)$$

251 where  $(x, y, z)_{\text{inlet,com}}$  is the inlet position relative to the center of mass in ship's coordinates,  
 252 positive for (bow, starboard, up), and  $d_{\text{com,sealevel}}$  is the distance of the center of mass to sea level.  
 253 For the RV Meteor III,  $(x, y, z)_{\text{inlet,com}} = (40 \text{ m}, -3 \text{ m}, -2 \text{ m})$  and  $d_{\text{com,sealevel}} = 1 \text{ m}$ . Moreover,  $\psi$  is  
 254 the roll angle positive for starboard down, and  $\gamma$  is the pitch angle positive for bow up. This  
 255 calculation is only an estimate, being accurate to at least the order. Not considered are surface  
 256 waves and the actual flow along the ship's hull which causes uncertainties in the actual depth of  
 257 the sampled water and in the measured properties. The mean  $d_{\text{inlet/sealevel}}$  is  $4.1 \text{ m} \pm 0.4 \text{ m}$  during  
 258 the equatorial section of autumn TRATLEQ and  $4.0 \text{ m} \pm 0.4 \text{ m}$  during the equatorial section of  
 259 spring TRATLEQ. Neglecting  $\Delta S_A$  in Equation (1),  $N^2$  at the inlet can be approximated to

$$260 \quad N^2 \approx g^2 * \rho * \alpha * \frac{T_z}{10^4} \text{ with } T_z = \frac{\sqrt{\text{var}(T)}}{\sqrt{\text{var}(d_{\text{inlet,sealevel}})}} \quad (3)$$

261 where  $T_z$  is the vertical temperature gradient and  $T$  is the temperature measured at the TSG inlet.

### 262 **3.3 Vertical shear of horizontal velocities**

263 In this study four different velocity data sets are considered: 1. Marine radar and  
 264 vmADCP data during autumn TRATLEQ, 2. CARTHE and SVP drifter experiment during  
 265 autumn TRATLEQ, 3. Hereon and SVP drifter experiment during spring TRATLEQ, and 4. the  
 266 EMP at the PIRATA site  $0^\circ\text{N}, 23^\circ\text{W}$ . For all four data sets, the zonal and meridional ocean  
 267 velocities are transformed into an along- and across-wind coordinate system. This transformation  
 268 allows an easier identification of the diurnal jet as, according to its definition, the jet is created  
 269 by wind that is trapped in the DWL. A positive across-wind component corresponds to velocities  
 270 to the left of the wind direction. The chosen wind value (satellite winds for 1-3, PIRATA winds  
 271 for 4) is the daily-mean value that is closest in time and space to the velocity measurements. The  
 272 vertical shear of horizontal velocities in along-wind direction,  $Sh_{\text{AI}}$ , is defined as the vertical  
 273 derivative of the along-wind velocities. In the following, vertical differences of horizontal  
 274 velocities and  $Sh_{\text{AI}}$  are considered as defined above.

### 275 **3.4 Diurnal cycle diagnostics**

276 Mean diurnal cycles are created by taking the mean of hourly bins. The time is  
 277 considered in Solar Apparent Time (SAT) so that solar noon is centred at 12:00, using a  
 278 conversion from Universal Time Coordinated to SAT (Koblick, 2021). The standard error is

279 computed as  $\frac{std}{\sqrt{f}}$  where *std* is the standard deviation and for the degrees of freedom, *f*, one  
 280 independent value per day is assumed. Diurnal patterns are compared in terms of the diurnal  
 281 timing and the diurnal amplitude. The peak (timing and value) is determined by a sinusoidal fit  
 282  $f(t) = \alpha * \sin(\omega * t + \varphi)$  as a function of time *t* [days] considering  $\pm 3.5$  h around the maxima  
 283 of the hourly means (i.e., 7 values of the hourly time series are used). Only periods ( $\frac{2*\pi}{\omega}$ ) between  
 284 0.5 and 2 days and phases ( $\varphi$ ) smaller than 1 day are considered. A symmetric fit is assumed to  
 285 be a good enough approximation, though there might be a tendency for a slower increase and a  
 286 faster decrease. The amplitude is calculated as the difference of the peak value determined by the  
 287 fit and the minimum of the hourly means between 6:00 SAT (sun rise) and the peak time. These  
 288 two characteristics are calculated for all robust diurnal cycles where the robustness is determined  
 289 using a signal-to-noise ratio. The signal is defined as the amplitude, and the noise is defined as  
 290 the arithmetic mean of the hourly computed standard errors. If the signal-to-noise ratio exceeds  
 291 2.5 for  $Sh_{AI}$  and 10 for  $N^2$ , we assume the presence of a robust diurnal cycle for  $Sh_{AI}$  and  $N^2$  as  
 292 well as the presence of a DWL and a diurnal jet, respectively. In order to determine the accuracy  
 293 of the sinusoidal fits, the bootstrapping method is utilised. This allows to establish confidence  
 294 intervals (CI) for the estimated parameters without prior knowledge of the shape of the  
 295 underlying distribution (Efron, 1979). For each diurnal fit, 10,000 resamples are taken from the  
 296 original data set with replacement and the same probability for each datapoint to be selected.  
 297 Each set of resamples has the sample size of the original data set. From the resulting distribution  
 298 of parameters for the diurnal fit, a 95% CI is given by taking the 2.5% and 97.5% quantiles.

### 299 **3.5 Wind speed, wind stress, and wind power input**

300 For comparability, the shipboard wind measurements from the TRATLEQ cruises in 30  
 301 m height and the PIRATA buoy wind measurements in 4 m height are scaled to 10 m wind  
 302 velocities using a logarithmic wind profile for neutral conditions. For a given height, *z*, the 10 m  
 303 winds can be calculated as

$$304 \quad W(10 \text{ m}) = W(z) * \frac{\ln(10 \text{ m}) - \ln(z_0)}{\ln(z) - \ln(z_0)} \quad (4)$$

305 where *W* is the wind velocity and  $z_0$  the surface roughness length (Fleagle & Businger, 1980)  
 306 with offshore assuming  $z_0 = 0.0002$  (Dutton, 1995). In the following, the horizontal wind vector

307 at 10 m height is denoted as  $\mathbf{u}_{10}$ . The wind stress vector,  $\boldsymbol{\tau}$ , is then defined as (Pacanowski,  
308 1987)

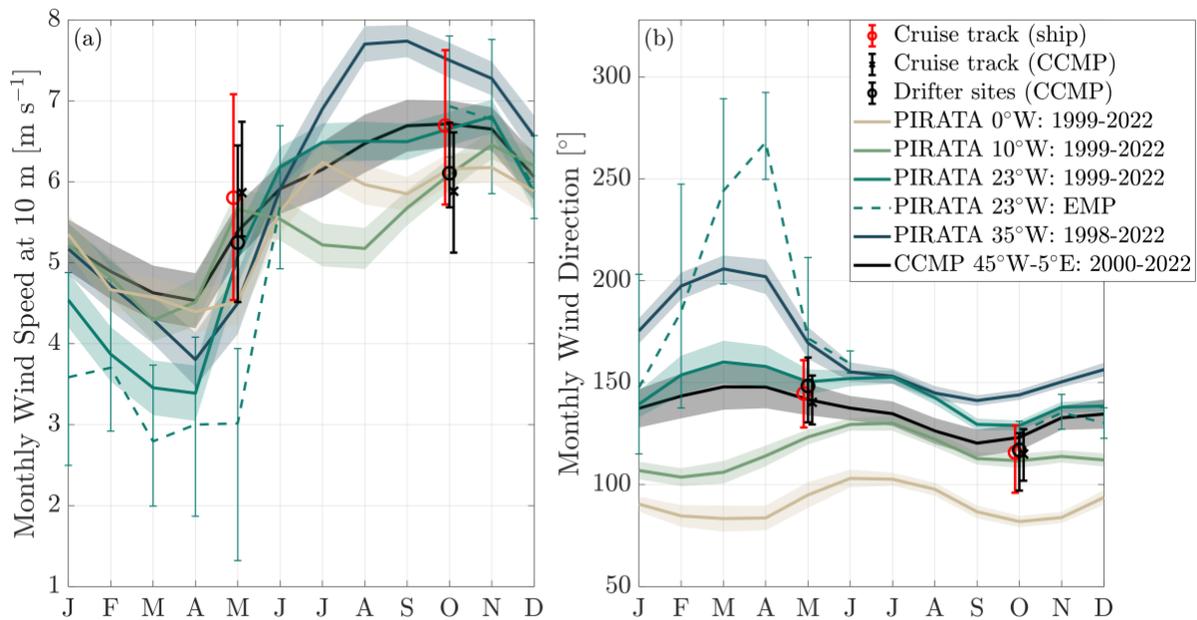
$$309 \quad \boldsymbol{\tau} = \rho_a * c_D * (\mathbf{u}_{10} - \mathbf{u}) * |\mathbf{u}_{10} - \mathbf{u}| \quad (5)$$

310 where  $\rho_a = 1.223 \text{ kg m}^{-3}$  is the density of air,  $c_D = 0.0013$  the drag coefficient, and  $\mathbf{u}$  the  
311 observed ocean surface velocity vector.

312 The WPI is the mechanical energy transferred by winds into the ocean. Part of this energy  
313 drives upper-ocean turbulence and is locally dissipated (Moum & Caldwell, 1985). The wind  
314 stress works on the ocean flow, so that the WPI is defined as

$$315 \quad WPI = \boldsymbol{\tau} * \mathbf{u} * \rho_w^{-1} \quad (6)$$

316 where  $\rho_w = 1025 \text{ kg m}^{-3}$  is the density of sea water. Note that ignoring the effect of the ocean  
317 surface velocity on  $\boldsymbol{\tau}$ , i.e. using  $\mathbf{u}_{10}$  instead of the velocity difference  $(\mathbf{u}_{10} - \mathbf{u})$  in Equation (5),  
318 leads to a 3% / 5% / 6% increase in the mean  $\boldsymbol{\tau}$  if  $\mathbf{u}$  were velocities at 0.5 m depth from the  
319 CARTHE drifters / velocities at 0.5 m depth from the Hereon drifters / velocities at 4.3 m depth  
320 of the EMP at 0°N, 23°W. This increase is derived using daily-mean wind speeds and hourly  
321 ocean velocities.

322 **4 Diurnal cycle in the equatorial Atlantic during two contrasting seasons**323 **4.1 Background conditions**

324

325 *Figure 2. Seasonal wind climatologies for the equatorial Atlantic and wind conditions during measurement*  
 326 *campaigns. Mean seasonal (a) wind speed at 10 m height and (b) wind direction in polar coordinates ( $0^\circ$*   
 327 *corresponds to westerlies,  $90^\circ$  to northerlies) derived from wind measurements at four different PIRATA sites*  
 328 *(coloured lines) and from CCMP winds (black lines) at the equator. The dashed line corresponds to PIRATA winds*  
 329 *measured at  $0^\circ\text{N}$ ,  $23^\circ\text{W}$  during the EMP (from October 2008 until June 2009). Values for the TRATLEQ expeditions*  
 330 *are derived from the ship's sensors along the cruise tracks (red circle) and by interpolating CCMP winds on the*  
 331 *drifter (black cross) and ship positions (black circle). Shading denotes  $\pm$  one standard error of the monthly mean*  
 332 *assuming one independent value per month, and error bars denote the interquartile range.*

333 Wind and net surface heat flux are assumed to potentially govern the pattern of the  
 334 diurnal cycles of near-surface shear and stratification. We start by investigating these two  
 335 atmospheric fields as background conditions to classify diurnal shear and stratification obtained  
 336 during the two TRATLEQ expeditions.

337 Monthly climatologies of wind speed and direction in the equatorial Atlantic are derived  
 338 from PIRATA buoys at four different longitudes and from the CCMP product zonally averaged  
 339 between  $45^\circ\text{W}$  and  $5^\circ\text{E}$ , using the complete available time series (Figure 2). The CCMP winds

340 have a clear seasonal cycle with a minimum speed of  $4.5 \text{ m s}^{-1}$  in April and a maximum speed of  
341  $6.7 \text{ m s}^{-1}$  in October associated with the meridional migration of the Intertropical Convergence  
342 Zone (Waliser & Gautier, 1993). The wind is directed towards the northwest, ranging from  $120^\circ$   
343 (SSE) in September to  $148^\circ$  (SE) in March. CCMP winds along the equator are mostly consistent  
344 with PIRATA measurements at  $23^\circ\text{W}$  and  $10^\circ\text{W}$ . The PIRATA measurements further indicate  
345 that whilst the amplitude of seasonal variations in wind speed and direction is decreasing from  
346 the western to the eastern basin, the wind direction is changing from dominantly easterlies to  
347 dominantly southerlies across the basin.

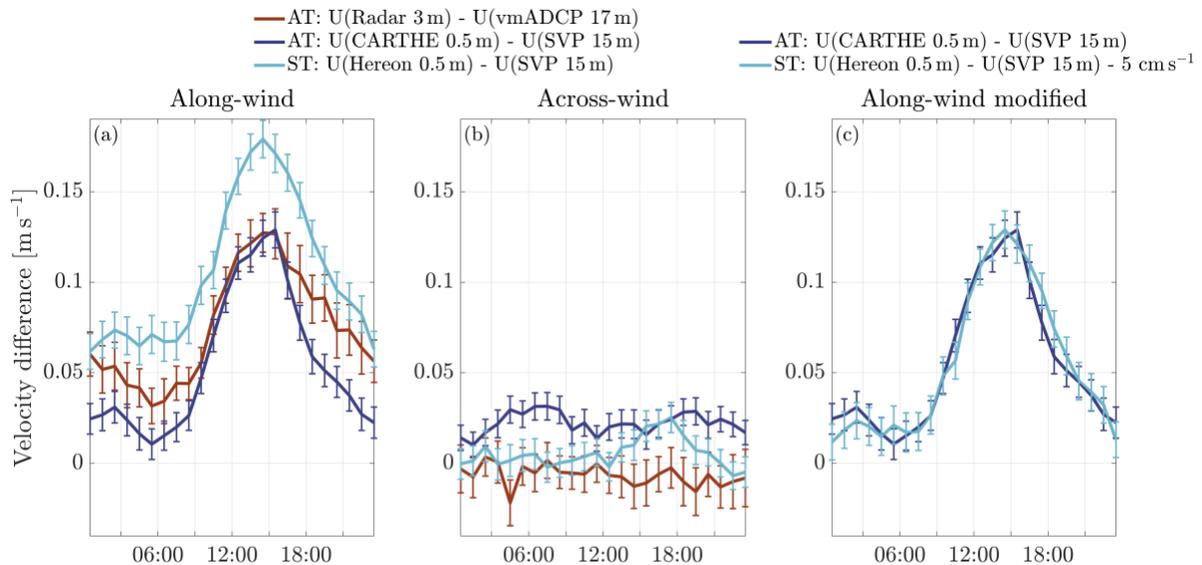
348         This seasonal wind variability is the main reason for the different wind conditions that  
349 prevail during the two TRATLEQ expeditions in boreal autumn 2019 and boreal spring 2022,  
350 shown as zonal means in Figure 2. Three wind estimates are provided to determine wind  
351 conditions during the TRATLEQ expeditions: an average of the wind from the ship's sensors  
352 during the TRATLEQ cruises and averages of the CCMP product interpolated on the ship as well  
353 as drifter positions. The mean wind speed differed by  $0.9 \text{ m s}^{-1}$  between the two cruises  
354 considering both CCMP and ship's sensor winds. The mean wind speed at the location of the  
355 drifter pairs is the same for the two experiments with similar wind conditions for the cruise and  
356 associated drifter experiment during autumn TRATLEQ and an offset of  $0.6 \text{ m s}^{-1}$  during spring  
357 TRATLEQ. The mean wind direction changed about  $30^\circ$  between the two cruises and between  
358 the two drifter experiments, yielding more northward (meridional) winds for spring TRATLEQ  
359 and more westward (zonal) winds for autumn TRATLEQ. A comparison of the conditions during  
360 the TRATLEQ expeditions with the zonal mean CCMP climatology suggests that the observed  
361 wind direction was typical and the wind speed atypical with respect to the climatological  
362 seasonal cycle. Note further that the mean CCMP wind speed is  $0.6 \text{ m s}^{-1}$  lower than the mean  
363 ship's sensor wind speed for both cruises, possibly pointing to a general offset between the  
364 directly measured wind speed and the CCMP product.

365         The net surface heat flux is computed as the sum of SWR, long-wave radiation, latent,  
366 and sensible heat flux. While the mean SWR differs only by  $14 \text{ W m}^{-2}$  between the two cruises,  
367 there is a higher difference between the drifter experiments. Averaging the SWR measured at the  
368 PIRATA buoys at  $0^\circ\text{N}$ ,  $10^\circ\text{W}$  and  $0^\circ\text{N}$ ,  $23^\circ\text{W}$  for the periods of the two drifter experiments

369 yields  $31 \text{ W m}^{-2}$  (14 %) higher SWR during autumn compared to spring TRATLEQ. The net  
 370 surface heat flux is higher by  $39 \text{ W m}^{-2}$  (45 %) respectively.

371 Hence, the two TRATLEQ periods are characterised by comparable mean wind speeds  
 372 for the two drifter experiments and a difference in mean wind speeds of  $0.9 \text{ m s}^{-1}$  between the  
 373 two cruises. The mean wind direction is shifted by about  $30^\circ$  comparing both the drifter  
 374 experiments and cruises. Furthermore, the net surface heat flux is noticeably weaker and the  
 375 SWR slightly weaker during spring TRATLEQ compared to autumn TRATLEQ.

#### 376 4.2 Observed near-surface diurnal shear



377  
 378 *Figure 3. Mean diurnal cycles of vertical differences of horizontal velocities in (a) along- and (b) across-wind*  
 379 *direction as a function of SAT. The velocity differences are obtained from the marine radar and the uppermost bin of*  
 380 *the vmADCP during autumn TRATLEQ (AT, dark red) and from the drifter experiments during autumn TRATLEQ*  
 381 *(AT, dark blue) and spring TRATLEQ (ST, light blue). Further, (c) compares the drifter experiments in (a) with a*  
 382 *background velocity difference of  $0.05 \text{ m s}^{-1}$  removed for spring TRATLEQ. The error bars represent the standard*  
 383 *error.*

384 The mean diurnal cycles of vertical differences of horizontal velocities are investigated  
 385 close to the surface, measured between 0.5 m and 15 m depth by CARTHE/Hereon and SVP  
 386 drifters and between 3 m and 17 m depth by the marine radar and the vmADCP (Figure 3). For  
 387 all data sets, there are stronger diurnal and background signals in the along-wind component than

388 in the across-wind component (Figure 3a,b), implying that the vertical differences of near-  
389 surface velocities are mainly wind driven. Therefore, only the along-wind component is  
390 considered in the following. A clear diurnal cycle is seen in the along-wind velocity differences  
391 (Figure 3a). Velocity differences are at their minimum in the morning hours (from 0:00 to 6:00  
392 SAT), increase after sunrise (6:00 SAT) until they reach their maximum between 14:00 and  
393 16:00 SAT (2 – 4 h after solar noon) and decrease in the evening - corresponding to the features  
394 of a diurnal jet. This results in diurnal amplitudes of  $11.8 \text{ cm s}^{-1}$  (autumn TRATLEQ drifter),  $9.6$   
395  $\text{cm s}^{-1}$  (autumn TRATLEQ shipboard), and  $11.7 \text{ cm s}^{-1}$  (spring TRATLEQ drifter). For the  
396 autumn TRATLEQ expedition, the diurnal amplitude obtained from the drifter experiment  
397 exceeds the one from shipboard measurements by 22%. While the magnitude of velocity  
398 differences is the same in the afternoon, there is a  $2.0 \text{ cm s}^{-1}$  weaker night-time velocity  
399 difference for the drifter measurements compared to the shipboard measurements. Overall, the  
400 two measurement techniques are in good agreement, showing the robustness of the diurnal cycle  
401 and giving confidence in the usage of both techniques. The distance criterion of 100 km used for  
402 the collocation of pairs of the two drifter types reveals that the diurnal cycle is also a horizontally  
403 large-scale feature, matching previous findings (Bellenger & Duvel, 2009).

404 The drifter experiments during the autumn and spring TRATLEQ expeditions yield  
405 closely resembling diurnal cycles for the vertical differences of horizontal velocities between 0.5  
406 m and 15 m depth (Figure 3c). While the pattern of the diurnal cycles agrees well, there is an  
407 offset of  $5 \text{ cm s}^{-1}$ . This offset can be better interpreted in terms of background shear when  
408 looking at the western, central, and eastern basin separately (Table 1). Background shear is  
409 analysed as the velocity difference between 0.5 m and 15 m depth averaged between midnight  
410 and sunrise, that is, 0:00 to 6:00 SAT. During this period, the DWL and hence the diurnal jet are  
411 removed by nocturnal mixing, and only the background shear remains. The seasonal and  
412 longitudinal variations of this background shear yield a range from  $-0.4 \text{ cm s}^{-1}$  to  $11.8 \text{ cm s}^{-1}$  in  
413 the along-wind component with more zonal background shear during spring TRATLEQ and  
414 more meridional background shear during autumn TRATLEQ. A comparison with the depth and  
415 strength of the EUC, which was generally shallower and stronger during spring compared to

416 autumn TRATLEQ (Brandt et al., 2023), indicates that the presence of a higher zonal  
 417 background shear goes along with a shallower and stronger EUC.

	Western: 37°W to 25°W	Central: 25°W to 17°W	Eastern: 17°W to 2°W
AT: wind speed [ $\text{m s}^{-1}$ ], wind direction [ $^{\circ}$ ]	$6.0 \pm 0.7$ , $128 \pm 15$	$6.0 \pm 0.7$ , $116 \pm 13$	$5.7 \pm 1.1$ , $100 \pm 18$
AT: diurnal amplitude of along-wind velocity differences [ $\text{cm s}^{-1}$ ]	13	13	12
AT: 0 – 6 SAT mean velocity difference along-wind (zonal, meridional) [ $\text{cm s}^{-1}$ ]	$-0.4 \pm 1.6$ ( $-3.2 \pm 1.5$ , $-3.1 \pm 1.4$ )	$3.5 \pm 1.5$ ( $-0.4 \pm 1.3$ , $4.2 \pm 1.5$ )	$3.4 \pm 1.2$ ( $-3.7 \pm 1.2$ , $1.3 \pm 1.1$ )
ST: wind speed [ $\text{m s}^{-1}$ ], wind direction [ $^{\circ}$ ]	$6.1 \pm 0.8$ , $158 \pm 14$	$5.5 \pm 1.2$ , $135 \pm 32$	$6.0 \pm 1.0$ , $127 \pm 16$
ST: diurnal amplitude of along-wind velocity differences [ $\text{cm s}^{-1}$ ]	12	13	14
ST: 0 – 6 SAT mean velocity difference along-wind (zonal, meridional) [ $\text{cm s}^{-1}$ ]	$8.4 \pm 2.2$ ( $-9.2 \pm 2.1$ , $-0.1 \pm 1.2$ )	$1.2 \pm 1.5$ ( $-2.2 \pm 1.7$ , $0.4 \pm 1.7$ )	$11.8 \pm 1.4$ ( $-9.2 \pm 1.6$ , $7.3 \pm 1.0$ )

418 *Table 1: Background conditions in the western, central, and eastern equatorial Atlantic basin during the TRATLEQ*  
 419 *drifter experiments. The mean CCMP wind speed and direction as well as the diurnal amplitude and the mean for*  
 420 *the period 0:00 to 6:00 SAT of the velocity differences between 0.5 and 15 m depth are calculated for three*

421 *longitudinal groups during the autumn TRATLEQ (AT) and spring TRATLEQ (ST) drifter experiments. For the 0 – 6*  
422 *SAT mean velocity difference, the zonal and meridional components are additionally given in brackets.*

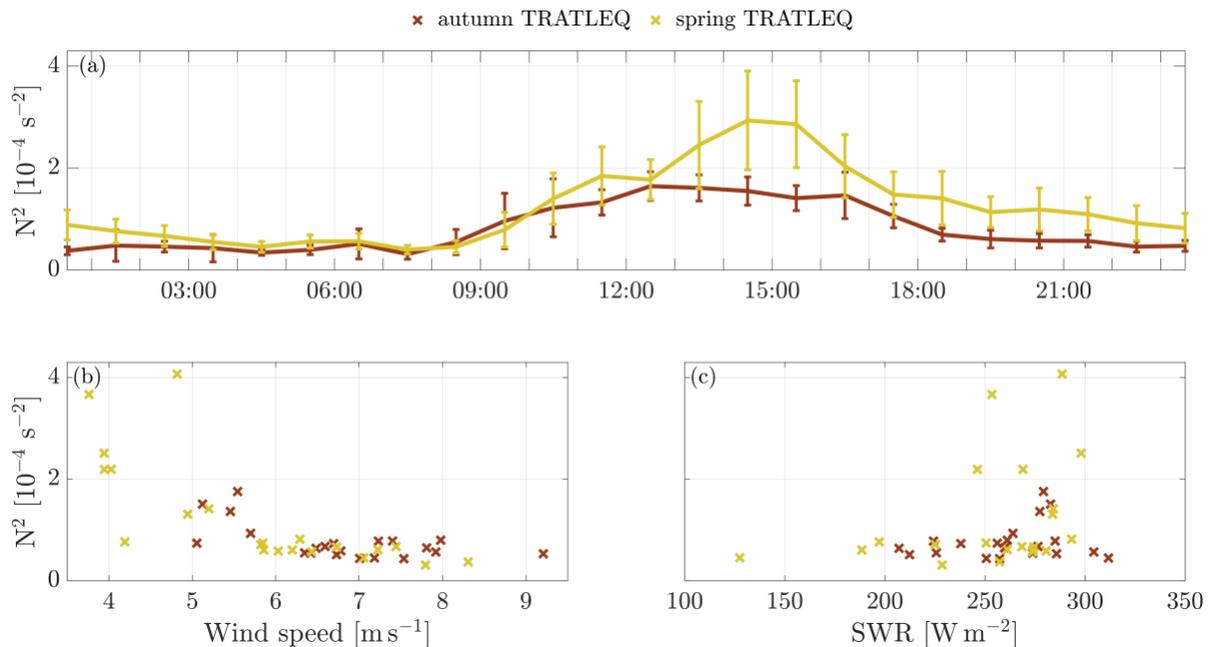
423         The meridional background velocity difference is strongest during phases of enhanced  
424 meridional winds, i.e., during autumn TRATLEQ and generally in the eastern basin. Marine  
425 radar and vmADCP measurements during the autumn TRATLEQ cruise reveal a strong positive  
426 correlation between the daily-mean meridional wind stress (ignoring the effect of the ocean  
427 surface velocity on  $\tau$ ) and the daily-mean vertical differences of meridional velocities, yielding a  
428 Pearson correlation coefficient of 0.81 using CCMP winds and 0.79 using ship's sensor winds.  
429 These correlations are significant at the 99%-level according to a Student's t-test. The  
430 correlations imply that, directly at the equator, a higher meridional wind stress results in higher  
431 meridional background shear. This is in line with the presence of the equatorial roll which is a  
432 shallow cross-equatorial overturning cell in the upper 80 m of the ocean with northward surface  
433 flow and a velocity reversal at about 25 m depth (Heukamp et al., 2022).

434         The wind speed during both drifter experiments is comparable in the three basins, while  
435 the wind direction is turning 30° westward from the eastern to the western basin. Like the wind  
436 speed, the diurnal amplitudes are comparable across the basin (Table 1).

437         These results suggest that the seasonal and longitudinal differences in background  
438 conditions mainly impact the background shear. With the wind speed being the only considered  
439 background condition that was similar for both TRATLEQ drifter experiments, the wind speed

440 might control the strength of the diurnal amplitude, i.e. the diurnal jet. This hypothesis about the  
 441 wind speed influence will be tested in section 5.

### 442 4.3 Observed near-surface diurnal stratification



443  
 444 *Figure 4. Stratification ( $N^2$ ) at 4 m depth as a function of SAT, wind speed, and SWR. Mean diurnal cycles of  $N^2$  are*  
 445 *shown (a) as a function of SAT. The daily-mean  $N^2$  is shown as a function of (b) wind speed at 10 m height and (c)*  
 446 *SWR. All parameters are derived from shipboard measurements during the spring (yellow) and autumn (red)*  
 447 *TRATLEQ cruises. The error bars in (a) represent the standard error.*

448 Both cruises are also characterised by a diurnal cycle of near-surface stratification ( $N^2$ )  
 449 estimated at a depth of about 4 m which shows weak stratification at night and maximum  
 450 stratification in the afternoon (Figure 4a), indicating the presence of a DWL for both cruises.  
 451 More precisely, the stratification is weakest about 1.5 h after sunrise and reaches its maximum  
 452 0.5 h (autumn TRATLEQ) to 2.5 h (spring TRATLEQ) after solar noon. While the diurnal cycles  
 453 calculated for the two cruises are aligned during the morning hours, there is higher near-surface  
 454 stratification in the afternoon and at night for spring TRATLEQ compared to autumn

455 TRATLEQ. Maximum and minimum  $N^2$  are  $1.6 * 10^{-4} \text{ s}^{-2}$  and  $0.3 * 10^{-4} \text{ s}^{-2}$  for autumn  
456 TRATLEQ and  $2.9 * 10^{-4} \text{ s}^{-2}$  and  $0.4 * 10^{-4} \text{ s}^{-2}$  for spring TRATLEQ.

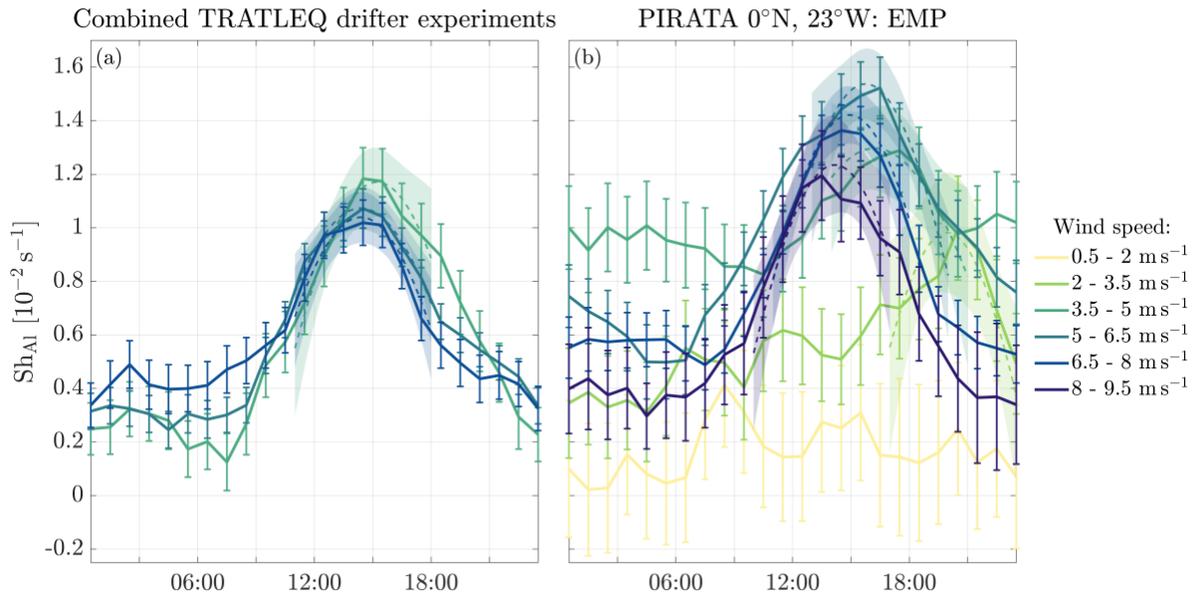
457 The effect of wind speed and SWR on daily-mean  $N^2$  is shown in Figure 4b,c including  
458 the contribution of both the diurnal amplitude and night-time values. There are strong negative  
459 Pearson correlation coefficients between daily-mean wind speed and daily-mean stratification of  
460  $-0.60$  for autumn TRATLEQ and  $-0.73$  for spring TRATLEQ, which are significant at the 99%-  
461 level according to a Student's t-test, suggesting higher stratification for lower wind speeds and  
462 vice versa. Although the mean wind speed only differed by  $0.9 \text{ m s}^{-1}$  between the two cruises,  
463 days with low wind speeds ( $< 5 \text{ m s}^{-1}$ ) occurred only during spring TRATLEQ, resulting in a  
464 higher spread of daily-mean stratification values compared to autumn TRATLEQ (Figure 4b).  
465 This might also explain the higher  $N^2$  maximum for spring compared to autumn TRATLEQ.  
466 Furthermore, there appears to be a wind speed threshold of about  $6 \text{ m s}^{-1}$  that separates the higher  
467 and lower spread of daily-mean stratification values. A higher spread of stratification values can  
468 also be found for higher SWR (Figure 4c). As the occurrence of high daily-mean stratification  
469 coincides with both the presence of low wind speeds and high SWR, the wind speed threshold  
470 can be interpreted as a threshold below which SWR becomes important. A similar wind speed  
471 threshold of  $6 \text{ m s}^{-1}$  was found off Peru for near-surface stratification to persist overnight,  
472 possibly creating multi-day near-surface stratification (T. Fischer et al., 2019). The effect of the  
473 wind speed on the diurnal stratification will be further analysed in section 5.

474 The observations indicate that a DWL and a diurnal jet were present during both  
475 TRATLEQ expeditions. While the diurnal cycles (amplitude and phase) of near-surface velocity  
476 differences were identical, the amplitude of diurnal near-surface stratification was larger for  
477 spring TRATLEQ compared to autumn TRATLEQ. In this comparison it should be noted that  
478 the mean wind speed of the two TRATLEQ drifter experiments, from which diurnal velocity  
479 cycles were derived, differed by only  $0.2 \text{ m s}^{-1}$  while the mean wind speed of the two TRATLEQ  
480 cruises, from which the diurnal stratification cycles were derived, differed by  $0.9 \text{ m s}^{-1}$ . Hence,  
481 the wind speed might determine the amplitudes of not only the diurnal jet, as hypothesised  
482 earlier, but also of the DWL. In the following, the influence of the wind speed on the diurnal

483 cycles of shear and stratification will be examined focussing mainly on longer-term observations  
 484 from the PIRATA buoy at 0°N, 23°W.

## 485 5 Wind speed dependence of the diurnal cycles

486 In this section, we explore the dependence of the near-surface shear and stratification on wind  
 487 speed. During the TRATLEQ drifter experiments, 98% of the observed daily-mean wind speeds  
 488 ranged from 3.4 m s<sup>-1</sup> to 7.7 m s<sup>-1</sup> and during the EMP at the PIRATA buoy at 0°N, 23°W from  
 489 0.4 m s<sup>-1</sup> to 9.3 m s<sup>-1</sup>. The effect of the wind speed on mean diurnal cycles of  $Sh_{AI}$  is analysed by  
 490 subsampling  $Sh_{AI}$  for each hour of the day into 1.5 m s<sup>-1</sup> wind speed intervals for drifters and  
 491 mooring data (Figure 5).

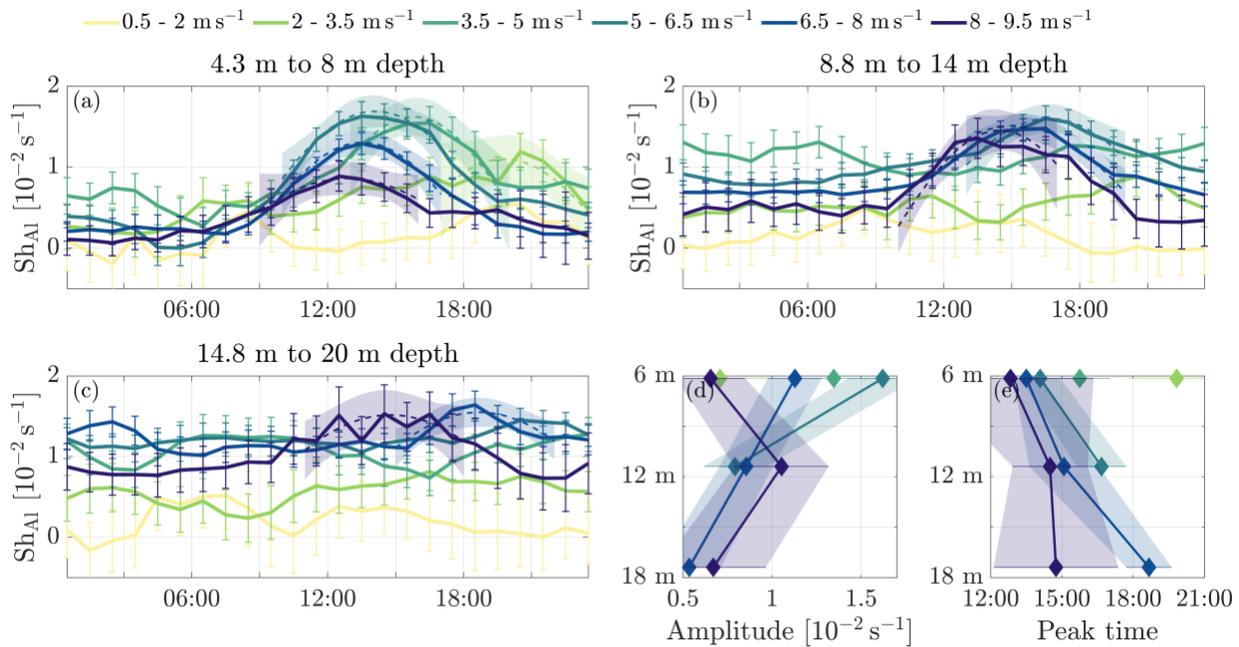


492

493 *Figure 5. Mean diurnal cycles of along-wind shear ( $Sh_{AI}$ ) as a function of SAT and wind speed.  $Sh_{AI}$  is derived (a)*  
 494 *between 0.5 m and 15 m depth for both drifter experiments and (b) between 4.3 m and 14.8 m depth for the EMP at*  
 495 *the PIRATA site at 0°N, 23°W. The colours correspond to different wind speed ranges. Wind speeds at 10 m height*  
 496 *are taken from (a) CCMP and (b) PIRATA. For robust diurnal patterns, the peak is fitted to a sinusoidal function*  
 497 *which is displayed by the dashed line. The error bars represent the standard error, and the shading marks the 95%*  
 498 *CI to estimate the fitted peak.*

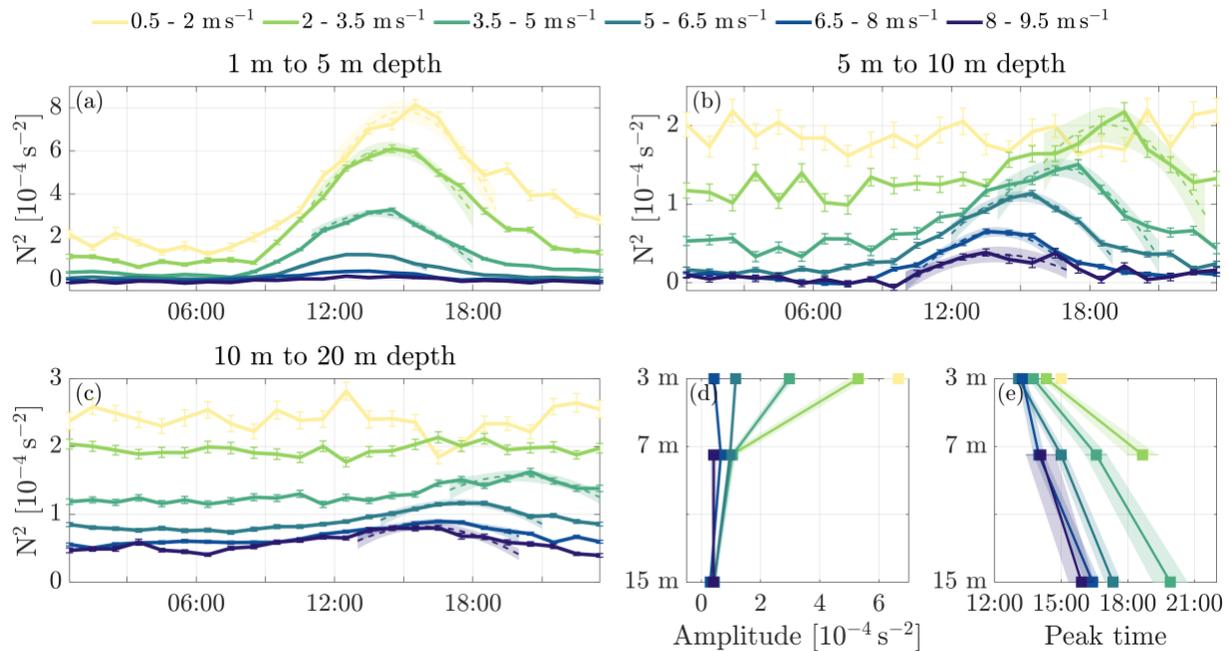
499 Diurnal cycles of  $Sh_{AI}$  averaged between 0.5 m and 15 m depth (4.3 m and 14.8 m depth) can be  
 500 observed for winds stronger than 2 m s<sup>-1</sup> as derived from the combined drifter experiments (the  
 501 PIRATA EMP at 0°N, 23°W). With increasing wind speed, the diurnal peak occurs earlier and

502 tends to last for a shorter time. The latter relation indicates the tendency of a shorter persistence  
 503 of the diurnal jet with higher wind speeds. The amplitude of the diurnal cycle of  $Sh_{AI}$  (Figure 5)  
 504 varies between  $4.7 (6.3) \text{ s}^{-1}$  and  $10.3 (10.5) * 10^{-3} \text{ s}^{-1}$  with the maximum being reached at  
 505 moderate winds of  $5 \text{ m s}^{-1}$  to  $6.5 \text{ m s}^{-1}$  ( $3.5 \text{ m s}^{-1}$  to  $5 \text{ m s}^{-1}$ ) considering the PIRATA EMP  
 506 (combined TRATLEQ drifter experiments). Note that the difference of about  $1.5 \text{ m s}^{-1}$  in the  
 507 wind speed at which maximum  $Sh_{AI}$  occurs in the two data sets might reduce considering the  
 508 wind speed offset between the shipboard measurements and the CCMP winds of  $0.6 \text{ m s}^{-1}$ . The  
 509 wind speed dependence of the diurnal pattern of  $Sh_{AI}$  is consistent in both data sets, though the  
 510 associated differences in amplitude and timing are more distinct in the mooring data set. This  
 511 might be a consequence of the larger range of seasons in the mooring data defining the  
 512 distribution of not only wind speed but also other possible influencing parameters such as heat  
 513 fluxes or near-surface Richardson number.



514  
 515 *Figure 6. Mean diurnal cycles of along-wind shear ( $Sh_{AI}$ ) as a function of SAT and wind speed.  $Sh_{AI}$  is derived*  
 516 *between (a) 4.3 m and 8.0 m, (b) 8.8 m and 14.0 m and (c) 14.8 m and 20.0 m depth for the EMP at the PIRATA site*  
 517 *at  $0^\circ\text{N}$ ,  $23^\circ\text{W}$ . For robust diurnal pattern, the peak is fitted to a sinusoidal function which is displayed by the dashed*  
 518 *line. For these cases, (d) the diurnal amplitude of  $Sh_{AI}$  and (e) the timing of the diurnal peak are displayed for the*  
 519 *three different depth ranges. The colours correspond to different wind speed ranges. The vertical error bars*  
 520 *represent the standard error, and the shading marks the 95% CIs to estimate the fitted peak.*

521 To address the vertical structure and descent of the diurnal shear signal,  $Sh_{AI}$  is evaluated  
 522 at 6.1 m depth (averaged between 4.3 m and 8.0 m, Figure 6a), 11.4 m depth (8.8 m and 14.0 m,  
 523 Figure 6b), 17.4 m depth (14.8 m and 20.0 m, Figure 6c) as well as at 22.6 m depth (20.0 m and  
 524 25.3 m, not shown) derived from the PIRATA EMP at  $0^\circ N$ ,  $23^\circ W$ . For the latter depth interval  
 525 no wind group shows robust diurnal characteristics as defined in section 3.5. Wind speeds need  
 526 to exceed  $2 \text{ m s}^{-1}$  for the diurnal jet to reach about 6 m depth, exceed  $5 \text{ m s}^{-1}$  to reach about 11 m  
 527 depth, and exceed  $6.5 \text{ m s}^{-1}$  to reach about 17 m depth. While at 6 m depth the diurnal amplitude  
 528 of  $Sh_{AI}$  (Figure 6d) is largest for moderate winds (as seen in Figure 6 for  $Sh_{AI}$  averaged for the  
 529 upper 15 m depth), from 11 m depth downward, the diurnal amplitude increases with increasing  
 530 wind speed. In general, the amplitude of  $Sh_{AI}$  decreases with depth for every wind group, except  
 531 for  $8 \text{ m s}^{-1}$  to  $9.5 \text{ m s}^{-1}$  winds where the amplitude just reaches its maximum at 11 m depth. The  
 532 diurnal peak of  $Sh_{AI}$  occurs later in the day with depth for all wind groups (Figure 6e), resulting  
 533 in mean descent rates of  $2.0 \text{ m h}^{-1}$ ,  $2.2 \text{ m h}^{-1}$ , and  $5.9 \text{ m h}^{-1}$  for wind speeds of  $5 \text{ m s}^{-1}$  to  $6.5 \text{ m s}^{-1}$ ,  
 534  $6.5 \text{ m s}^{-1}$  to  $8 \text{ m s}^{-1}$ , and  $8 \text{ m s}^{-1}$  to  $9.5 \text{ m s}^{-1}$ , respectively.

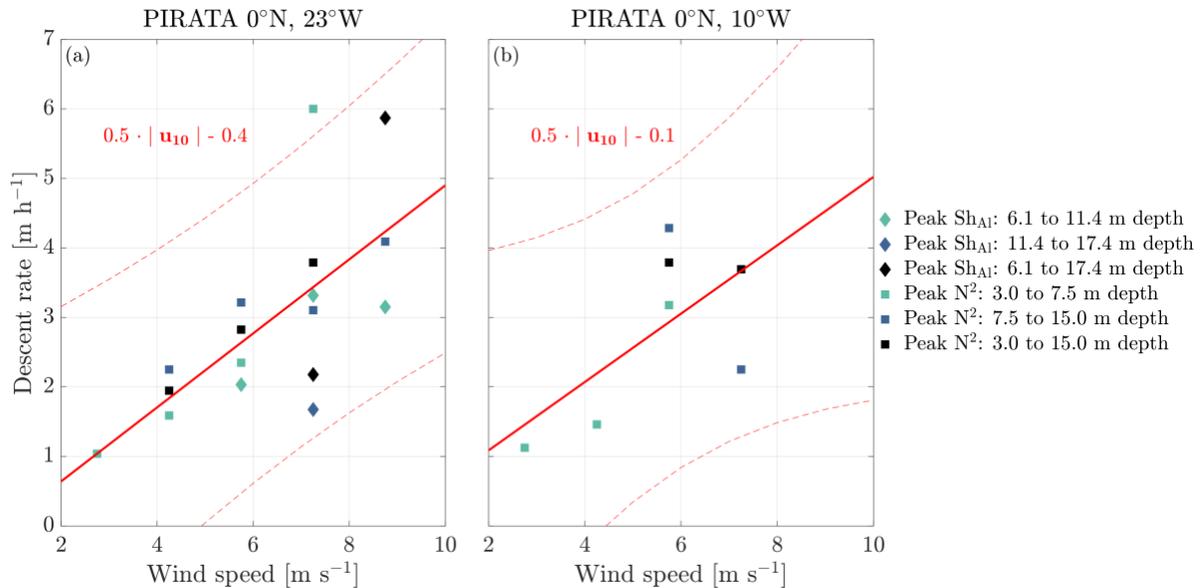


535  
 536 *Figure 7. Mean diurnal cycles of stratification ( $N^2$ ) as a function of SAT and wind speed.  $N^2$  is derived between (a) 1*  
 537 *m and 5 m, (b) 5 m and 10 m, and (c) 10 m and 20 m depth for the PIRATA site at  $0^\circ N$ ,  $23^\circ W$ . For robust diurnal*  
 538 *patterns, the peak is fitted to a sinusoidal function which is displayed by the dashed line. For these cases, (d) the*  
 539 *diurnal amplitude of  $Sh_{AI}$  and (e) the timing of the diurnal peak are displayed for the three different depth ranges.*

540 *The colours correspond to different wind speed ranges. The vertical error bars represent the standard error, and the*  
541 *shading marks the 95% CIs to estimate the fitted peak.*

542 To address the vertical structure and descent of the diurnal stratification signal,  $N^2$  is  
543 analysed at 3 m depth (averaged between 1 m and 5 m, Figure 7a), 7.5 m depth (10 m and 15 m,  
544 Figure 7b), 15 m depth (10 m and 20 m, Figure 7c), and 30 m depth (20 m and 40 m, not shown)  
545 derived from the PIRATA buoy at 0°N, 23°W. At all these depths  $N^2$  increases with decreasing  
546 wind speed, valid all day through. The standard error also increases with decreasing wind speed,  
547 indicating a higher spread in  $N^2$  values for weaker winds. This can be explained by the surface  
548 layer being more responsive to the heat flux at low wind speeds (Matthews et al., 2014). At 3 m  
549 depth, the diurnal cycle of  $N^2$  (as described in section 4.3) is more pronounced with weaker wind  
550 speeds, reaching amplitudes of  $6.7 \cdot 10^{-4} \text{ s}^{-2}$  for the weakest and  $0.4 \cdot 10^{-4} \text{ s}^{-2}$  for the strongest wind  
551 group (Figure 7d). This relation between the diurnal amplitude of  $N^2$  and the wind speed also  
552 applies to the other depths. However, the spread of the amplitudes reduces with depth. Wind  
553 speeds need to exceed  $2 \text{ m s}^{-1}$  for the DWL to reach about 7.5 m depth and exceed  $3.5 \text{ m s}^{-1}$  to  
554 reach about 15 m depth. There is no robust diurnal signal visible anymore at 30 m depth. Hence,  
555 the DWL reaches deeper for stronger winds but also becomes weaker. Besides, the wind affects  
556 the timing of the diurnal cycle with an earlier peak occurring for stronger winds (Figure 7e). The  
557 resulting spread of the peak times increases with depth. The descent of the maximum of  $N^2$   
558 becomes faster with increasing wind speeds with descent rates of  $1.0 \text{ m h}^{-1}$ ,  $1.9 \text{ m h}^{-1}$ ,  $2.8 \text{ m h}^{-1}$ ,

559 3.8 m h<sup>-1</sup>, and 4.1 m h<sup>-1</sup> for wind speeds of 2 m s<sup>-1</sup> to 3.5 m s<sup>-1</sup>, 3.5 m s<sup>-1</sup> to 5 m s<sup>-1</sup>, 5 m s<sup>-1</sup> to 6.5  
 560 m s<sup>-1</sup>, 6.5 m s<sup>-1</sup> to 8 m s<sup>-1</sup>, and 8 m s<sup>-1</sup> to 9.5 m s<sup>-1</sup>, respectively.



561

562 *Figure 8. Descent rates as a function of wind speed. The descent rates are calculated from the peaks of N<sup>2</sup> (square)*  
 563 *and Sh<sub>AI</sub> (diamond) which are shown in Figures 6e and 7e for (a) the PIRATA site at 0°N, 23°W and (b) the PIRATA*  
 564 *site at 0°N, 10°W. The different depth intervals are indicated in colour. The red lines represent a linear fit to the*  
 565 *data (solid) and the associated 95% prediction interval (dashed).*

566 According to the timing of the peaks, descent rates of both diurnal Sh<sub>AI</sub> and N<sup>2</sup> increase  
 567 with increasing wind speed (Figure 8a). There seems to be a linear relationship between the wind  
 568 speed and the descent rate: for every increase of 2 m s<sup>-1</sup> in wind speed, the descent rate appears  
 569 to increase by 1 m h<sup>-1</sup>. Note that this linear regression is computed excluding one Sh<sub>AI</sub> peak as an  
 570 outlier. The timing and amplitude of the diurnal cycle of N<sup>2</sup> at the PIRATA buoy at 0°N, 10°W  
 571 (not shown) are similar to the ones at 0°N, 23°W presented before. The descent rates are also  
 572 similar, yielding the same slope of the linear fit for the descent rate as a function of the wind  
 573 speed (Figure 8b). The main differences between the mooring sites are a higher background

574 stratification at 30 m depth as well as enhanced variability at 15 m depth disguising the diurnal  
575 signal at 10°W compared to 23°W.

## 576 **6 Summary and Discussion**

577 This study focusses on the diurnal cycles of shear and stratification, respectively called  
578 the diurnal jet and the DWL, in the upper 20 m of the equatorial Atlantic and on their wind  
579 dependence. Shear and stratification are primarily derived from drifter experiments and  
580 shipboard measurements during the TRATLEQ expeditions in October 2019 and May 2022.  
581 These two seasons differed in wind direction, net surface heat flux as well as strength and depth  
582 of the EUC but were comparable in wind speed. Despite these partly contrasting conditions,  
583 similar diurnal jets with an amplitude of about 11 cm s<sup>-1</sup> and similar DWLs are observed. The  
584 main difference between the expeditions is the 5 cm s<sup>-1</sup> background velocity difference between  
585 0.5 m and 15 m depth in May 2022, associated with a shallower and stronger EUC (Brandt et al.,  
586 2023). We suggest that zonal background shear is mainly related to the vertical migration of the  
587 EUC core (Brandt et al., 2016, 2023) and meridional background shear to the presence of the  
588 equatorial roll (Heukamp et al., 2022), with zonal background shear being dominant. The  
589 generalizability of this statement is limited as we only considered two points in time. Potential  
590 constraints of the analysed velocity and associated shear estimates are the various vertical ranges  
591 (drogue length of drifters, wave length of the dominant surface waves defining the penetration  
592 depth of the marine radar, and bin size of the vmADCP), which are averaged to obtain the  
593 assigned depth values, and wind slip for drifter measurements (see Text S1 for more details).  
594 Still, the observed diurnal jets are in the range of previous observational results in the tropical  
595 and subtropical Atlantic and Pacific (Price et al., 1986; Kudryavtsev & Soloviev, 1990; Cronin &  
596 Kessler, 2009; Wenegrat & McPhaden, 2015; Sutherland et al., 2016), although a direct  
597 comparison of diurnal jet diagnostics is complicated by the use of different depth levels and by  
598 different background and in particular wind conditions. The comparison of the two TRATLEQ  
599 expeditions suggests that, at least for wind speeds around 6 m s<sup>-1</sup>, the diurnal jet amplitude is  
600 independent of the surface heat flux and the wind direction but possibly dependent on the wind  
601 speed. For lower wind speeds, it is possible that the surface heat flux plays a role again  
602 (Matthews et al., 2014). The hypothesis of a wind speed dependence of the diurnal jet amplitude

603 is tested and supported by examining observational records with a larger spread in wind  
604 conditions taken at the PIRATA buoys.

605

606 At the PIRATA buoy at 0°N, 23°W, diurnal jets and DWLs are observed for wind speeds  
607 ranging from 2 m s<sup>-1</sup> to 9.5 m s<sup>-1</sup> and from 0.5 m s<sup>-1</sup> to 9.5 m s<sup>-1</sup>, respectively. Both diurnal jet  
608 and DWL descend deeper and reach their peak earlier for stronger winds. For the DWL, the  
609 diurnal stratification amplitude increases with decreasing wind speed. For the diurnal jet, the  
610 diurnal shear amplitude at 6 m depth is maximum for moderate wind speeds of 5 m s<sup>-1</sup> to 6.5 m s<sup>-1</sup>  
611 <sup>1</sup>. At 11 m depth and below, the diurnal shear amplitude is maximum for maximum wind speeds.  
612 In the following, we will discuss 1. wind speed thresholds, 2. wind speed dependence of the  
613 penetration depth, and 3. wind speed dependence of diurnal amplitudes.

- 614 1. Our findings match with model results suggesting a minimum wind speed threshold of 2  
615 m s<sup>-1</sup> for the existence of a diurnal jet (Hughes et al., 2020a, 2021). However, our  
616 observations do not support a maximum wind speed threshold (at least up to 9.5 m s<sup>-1</sup>  
617 winds) for both DWL and diurnal jet as found by Hughes et al. (2021) with 8 m s<sup>-1</sup> in the  
618 subtropical Pacific, by Thompson et al. (2019) with 7.6 m s<sup>-1</sup> in the equatorial Indian  
619 Ocean, by Matthews et al. (2014) with 6 m s<sup>-1</sup> in the equatorial Indian Ocean, and by  
620 Kudryavtsev & Soloviev (1990) with roughly 6.6 m s<sup>-1</sup> in the equatorial Atlantic. In  
621 contrast, our observations indicate even for wind speeds of 8 m s<sup>-1</sup> to 9.5 m s<sup>-1</sup> (no  
622 observations for higher wind speeds) the presence of a weak but sufficient DWL to trap  
623 wind momentum and to generate a diurnal jet. However, we do see a wind speed  
624 threshold of about 6 m s<sup>-1</sup> above which the daily-mean stratification is clearly reduced.  
625 Followingly, the discrepancy between our findings and previous studies might be a result  
626 of varying definitions of the DWL, especially with respect to thresholds such as  
627 minimum penetration depth and minimum diurnal amplitude. The existence of diurnal  
628 dynamics also for high wind speeds, as suggested in this study, indicates that also DCT  
629 can occur in those wind conditions.
- 630 2. Our results indicate an increase in the penetration depth of the diurnal jet and the DWL  
631 with increasing wind speed. This is in line with earlier observations (Price et al., 1986;  
632 Hughes et al., 2020b; Masich et al., 2021). Note that our results give the impression that  
633 the DWL reaches deeper than the diurnal jet. We assume that this is solely a consequence

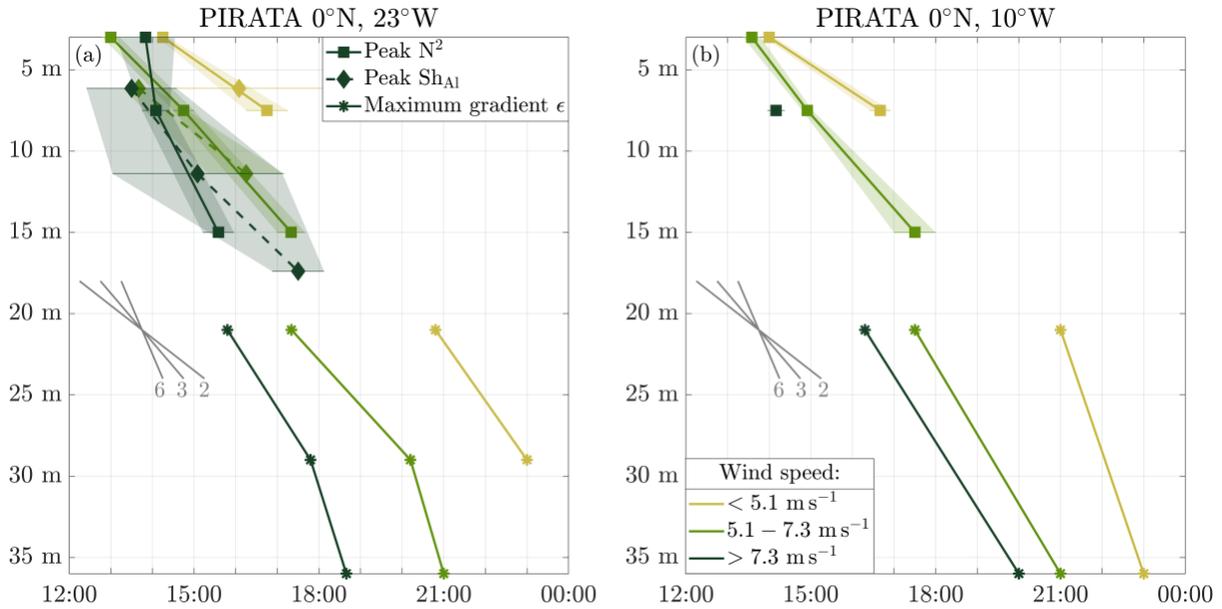
634 of a reduced signal-to-noise ratio of the velocity data, leading to the signal of the  
635 descending diurnal jet being lost earlier than the one of the DWL. The maximum  
636 penetration depth for the DWL and the diurnal jet should be the same, as shown, e.g., in  
637 Smyth et al. (2013).

638 3. We find that a stronger wind stress does not necessarily generate stronger vertical shear  
639 between fixed depth levels in the upper ocean, a behaviour opposite to that of a classical  
640 wall layer, as Price et al. (1986) already observed. As a function of wind speed, we find  
641 small but noticeable variations in the diurnal jet amplitude between 0.5 m and 15 m depth  
642 and distinct variations considering smaller depth intervals. In contrast to our  
643 observations, a near-uniform diurnal jet amplitude has been previously suggested which  
644 only depends on the net surface heat flux and is independent of the wind stress (Price et  
645 al., 1986; Sutherland et al., 2016). An idealised simulation by Hughes et al. (2020a)  
646 showed a dependence of the maximum shear on the wind speed for all considered mixing  
647 schemes with the maximum shear decreasing with increasing wind speed for winds  
648 stronger than  $2 \text{ m s}^{-1}$ . A main difference between our observations and those used for  
649 previous studies is the duration with only a few days of measurements in the previous  
650 studies. Therefore, one possible explanation for the different conclusions on wind  
651 dependence might be the hypothesis of a memory of previous diurnal events (Sutherland  
652 et al., 2016). If a memory exists, changes in wind speed will have little influence on the  
653 diurnal diagnostics considering a time span of a few days but will be apparent in longer  
654 observational records. However, the observed decrease in the diurnal amplitude with  
655 depth for all wind speeds (except for the diurnal jet at winds of  $8 \text{ m s}^{-1} - 9.5 \text{ m s}^{-1}$ )  
656 suggests that the amplitude of both the diurnal jet and DWL will be underestimated if a  
657 shallowest usable depth of 11 m (Masich et al., 2021) or 10 m (Smyth et al., 2013) is  
658 used. This stresses the importance of near-surface measurements to properly evaluate the  
659 near-surface heat and momentum budget.

### 660 **6.1 Descent rates of diurnal jet and diurnal warm layer and their relation to deep-cycle** 661 **turbulence as a function of wind speed**

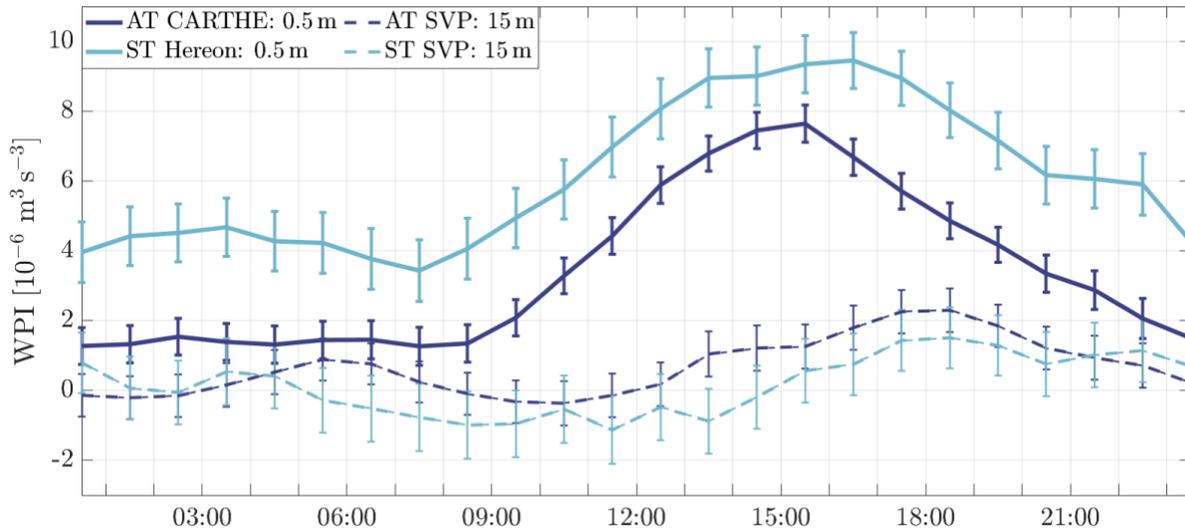
662 This study demonstrates the wind dependence of the timing of the diurnal peak for both  
663 shear and stratification. In line with observations by Smyth et al. (2013), the timing of the two

664 parameters is usually similar. We find at all considered depth levels earlier shear and  
665 stratification peaks for stronger winds, consistent with the simulation by Hughes et al. (2020a).  
666 Furthermore, we find that the descent rates of shear and stratification peaks increase with higher  
667 wind speed with values of  $1.0 \text{ m h}^{-1}$ ,  $1.9 \text{ m h}^{-1}$ ,  $2.8 (2.0) \text{ m h}^{-1}$ ,  $3.8 (2.2) \text{ m h}^{-1}$ , and  $4.1 (5.9) \text{ m h}^{-1}$   
668 for wind speeds of  $2 \text{ m s}^{-1}$  to  $3.5 \text{ m s}^{-1}$ ,  $3.5 \text{ m s}^{-1}$  to  $5 \text{ m s}^{-1}$ ,  $5 \text{ m s}^{-1}$  to  $6.5 \text{ m s}^{-1}$ ,  $6.5 \text{ m s}^{-1}$  to  $8 \text{ m s}^{-1}$ ,  
669  $1$ , and  $8 \text{ m s}^{-1}$  to  $9.5 \text{ m s}^{-1}$  considering stratification (shear), respectively. According to a linear  
670 regression, the descent rate increases by  $1 \text{ m h}^{-1}$  every  $2 \text{ m s}^{-1}$  wind speed. The observed descent  
671 rates are in line with observations for the DWL descent by Hughes et al. (2020b) who found  
672 descent rates in the upper 8 m of the ocean of  $0.3 \text{ m h}^{-1}$ ,  $1 \text{ m h}^{-1}$ , and  $4 \text{ m h}^{-1}$  for wind speeds of  
673  $1.6 \text{ m s}^{-1}$ ,  $4.0 \text{ m s}^{-1}$ , and  $7.6 \text{ m s}^{-1}$ , respectively. The observed descent rate of  $2 \text{ m h}^{-1}$  in the upper  
674 20 m of the ocean for mean winds of  $6 \text{ m s}^{-1}$  by Sutherland et al. (2016) is also in agreement with  
675 our results. The fact that these two experiments were conducted away from the equator ( $12^\circ\text{N}$  to  
676  $18^\circ\text{N}$  and  $25.6^\circ\text{N}$ , respectively) and match our results suggests that the descent rate is  
677 independent of the Coriolis parameter at least up to subtropical regions. Furthermore, the  $6 \text{ m h}^{-1}$   
678 descent rate for both the DWL and the diurnal jet observed by Smyth et al. (2013) in 15 m to 50  
679 m depth in the equatorial Pacific for a mean wind speed at 10 m height of about  $8 \text{ m s}^{-1}$  is at the  
680 upper limit of the above-mentioned relations. The multi-monthly mean (May 2004 to February  
681 2005) descent rate of the diurnal jet of  $5 \text{ m h}^{-1}$  observed between 7.5 m and 17.5 m depth also in  
682 the equatorial Pacific for mean winds at 10 m height of  $7 \text{ m s}^{-1}$  (Pham et al., 2017) also exceeds  
683 our observations. The higher descent rates observed in the equatorial Pacific compared to our  
684 findings in the equatorial Atlantic could indicate the presence of background conditions in the  
685 equatorial Pacific that facilitate the descent. We expect that marginal instability could be such a  
686 condition as it was found to be more present in the equatorial Pacific than Atlantic (Moum et al.,  
687 2023) and it is assumed to facilitate the descent of the diurnal jet (Lien et al., 1995; Masich et al.,  
688 2021). However, marginal instability has not been analysed in this study and further research is  
689 needed to identify the causalities of possible different descent rates in the equatorial Atlantic and  
690 Pacific.



691  
 692 *Figure 9. Descent of DWL, diurnal jet, and DCT as a function of SAT, depth and wind speed at the PIRATA sites (a)*  
 693 *0°N, 23°W and (b) 0°N, 10°W.  $N^2$  peaks (square),  $Sh_{AI}$  peaks (diamond), and the times of maximum temporal*  
 694 *dissipation ( $\epsilon$ ) gradient (asterisk, estimated from Figure 9 of Moum et al. (2023)) are presented for three different*  
 695 *wind groups in colour. The shading marks the 95% CIs to estimate the fitted diurnal peak. As a reference for the*  
 696 *descent rates, nominal slopes of 2 m h<sup>-1</sup>, 3 m h<sup>-1</sup> and 6 m h<sup>-1</sup> are indicated by the grey lines.*

697 The observed timing of the diurnal stratification and shear peaks as a function of depth  
 698 and wind speed can be used to examine the hypothesis of Moum et al. (2023) that the wind-  
 699 dependent delay of DCT may be a direct result of the wind-dependent DWL deepening. Both the  
 700 maximum temporal dissipation gradient found by Moum et al. (2023) and the peak of  
 701 stratification and shear at the PIRATA mooring sites (Figure 9) show an earlier onset or peak,  
 702 respectively, for stronger winds. This indicates that the wind-dependent descent rates of the DWL  
 703 and the diurnal jet indicated in this study also reflect in the timing of DCT. However, it remains  
 704 unclear when and where instabilities are triggered. The exact timing of the onset and the peak of  
 705 diurnal shear, shear instabilities and DCT might also depend on background stratification and  
 706 shear. Further studies are needed to better understand the processes. We suggest that studying  
 707 DCT as a function of wind speed can help to relate the wind-dependent diurnal jet to DCT, for  
 708 which Moum et al. (2023) found a wind dependence of the strength but not of the descent rate.  
 709 Note that also the wind-dependent strength of DCT might be explained by the wind-dependent  
 710 penetration depth and amplitude of the diurnal jet.

711 **6.2 Impact of the diurnal cycle on the wind power input**

712  
 713 *Figure 10. Mean diurnal cycles of WPI as a function of SAT. The WPI is computed for spring (ST, light blue) and*  
 714 *autumn (AT, dark blue) TRATLEQ drifter experiments with velocities at 0.5 m and at 15 m depth depicted by solid*  
 715 *and dashed lines, respectively. The error bars represent the standard error.*

716 The near-surface diurnal dynamics described in this study also reflect in the WPI (Figure  
 717 10) and thus impact the amount of mechanical energy transferred by winds into the ocean. There  
 718 is a diurnal cycle in the WPI derived from the autumn (spring) TRATLEQ drifter velocities at  
 719 0.5 m depth, leading to a  $1.62$  ( $1.60$ )  $\times 10^{-6} \text{ m}^3 \text{ s}^{-3}$ , i.e., 59% (32%) increase of the diurnal mean  
 720 WPI compared to the night-time WPI. The calculation of a fictive WPI using the autumn (spring)  
 721 TRATLEQ drifter velocities at 15 m depth leads to a reduction of the WPI by 80% (96%). This  
 722 underestimation of the available surface kinetic energy stresses the relevance of considering the  
 723 diurnal jet and of actually observing surface velocities instead of taking, e.g., 15 m velocities as  
 724 surface velocities. Furthermore, it shows that DWL and diurnal jet not only impact energy  
 725 transfer into the mixed layer but also impact air-sea fluxes and the amount of energy within the  
 726 DWL and below which is available for mixing.

727 **7 Conclusion**

728 This study examines the diurnal jet and DWL in the equatorial Atlantic, focussing on the  
 729 impact of the wind speed. Our analysis demonstrates that the wind speed influences timing,  
 730 amplitude, penetration depth, and descent rate of DWL and diurnal jet. The presented wind-

731 dependent descent rate of the diurnal jet and DWL can explain the wind-dependent onset of  
732 DCT. Furthermore, the diurnal dynamics impact the energy input into the ocean through the  
733 WPI. The question of how much of this energy is used to enhance turbulence during the descent  
734 of the DWL in the DCT layer remains open. Our results enhance the understanding of diurnal  
735 dynamics and stress the importance of near-surface measurements of, in particular, velocity. We  
736 want to emphasize that satellite missions aiming to resolve absolute ocean currents could provide  
737 additional data for better regional characterization of diurnal surface velocity variability  
738 (Ardhuin et al., 2019; Villas Bôas et al., 2019). Our results and in particular the TRATLEQ  
739 velocity data sets, which allow for a basin-scale view of velocities and shear in the upper metres  
740 of the ocean and which are presented in this study for the first time, can contribute to calibrate  
741 and validate satellite missions that aim to resolve absolute ocean currents like the current SWOT  
742 mission (Morrow et al., 2019) or possible future missions based on advanced Doppler-radar  
743 techniques as suggested for Odysea (Rodríguez et al., 2019) or SKIM (Ardhuin et al., 2018). Our  
744 results will enable the examination of possible offsets of satellite measurements due to sampling  
745 at various hours of the day. This study can also facilitate the validation of ocean models that aim  
746 to resolve diurnal dynamics (Bernie et al., 2007), aim to be energetically consistent (Eden et al.,  
747 2014; Gutjahr et al., 2021), or aim to correctly represent surface currents for other applications,  
748 e.g., to deduce Sargassum drift (Van Sebille et al., 2021). Moreover, this study points out that the  
749 diurnal cycle can be captured by vessel-mounted observation systems, which might be useful for  
750 further studies on spatial pattern of diurnal dynamics.

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769

### 770 **Open Research**

771 The marine radar and vmADCP measurements used to derive 3 m and 17 m depth on-track ocean  
772 velocities during autumn TRATLEQ are available at Pangaea in Carrasco & Horstmann (2024)  
773 and in Brandt et al. (2022), respectively. The drifter data used to derive velocities at 0.5 m and 15  
774 m depth for autumn TRATLEQ in the study are available at Pangaea in Hans & Brandt (2021).  
775 For spring TRATLEQ, the Hereon drifter positions are available at Pangaea in Horstmann et al.  
776 (2023) and the SVP drifter positions at NOAA's OSMC ERDDAP via  
777 <https://www.aoml.noaa.gov/phod/gdp/data.php> with the relevant ID/WMO numbers listed in  
778 Table S2. The ID/WMO numbers of the SVP drifters deployed during autumn TRATLEQ are  
779 listed in Table S1. TSG, pitch and roll data to derive a stratification estimate at 4 m depth as well  
780 as wind and radiation data for the two TRATLEQ cruises are available at the Dship system via  
781 [dship.bsh.de](http://dship.bsh.de). Temperature, salinity and wind data from the PIRATA buoys used in this study are  
782 available from the Global Tropical Moored Buoy Array at  
783 <https://www.pmel.noaa.gov/tao/drupal/disdell/>. The access to the heat flux data used from  
784 ePIRATA is described at <https://www.aoml.noaa.gov/phod/epirata/>. The velocity data at the  
785 PIRATA site at 0°N, 23°W during the EMP are available at  
786 [https://www.pmel.noaa.gov/tao/drupal/disdell/adcp\\_0n23w/index.html](https://www.pmel.noaa.gov/tao/drupal/disdell/adcp_0n23w/index.html). The satellite CCMP V2.0  
787 wind data are available at REMSS via [www.remss.com](http://www.remss.com) (Wentz et al., 2015). All analyses were  
788 performed and all figures created using MATLAB R2021a.

789

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