

1 **Implementation and evaluation of SNICAR snow albedo scheme in Noah-MP (version 5.0)**
2 **land surface model**

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14 **Key points**

- 15 • We enhance Noah-MP snow albedo modeling by implementing physical snow radiative
16 transfer and aging processes from the latest SNICAR
17 • The Noah-MP/SNICAR simulated snow albedo exhibits superior performance compared
18 to the default snow albedo scheme at validation sites
19 • Noah-MP/SNICAR quantifies the impact of snow grain size, shape, and light-absorbing
20 particles on snow albedo and radiative forcing

21

22 **Abstract**

23 The widely-used Noah-MP land surface model (LSM) currently adopts snow albedo
24 parameterizations that are semi-physical in nature with nontrivial uncertainties. To improve
25 physical representations of snow albedo processes, a state-of-the-art snowpack radiative transfer
26 model, the latest version of Snow, Ice, and Aerosol Radiative (SNICAR) model, is integrated into
27 Noah-MP in this study. The coupled Noah-MP/SNICAR represents snow grain properties (e.g.,
28 shape and size), snow aging, and physics-based snow-aerosol-radiation interaction processes. We
29 compare Noah-MP simulations employing the SNICAR scheme and the default semi-physical
30 Biosphere-Atmosphere Transfer Scheme (BATS) against in-situ snow albedo observations at three
31 Rocky Mountain field stations. The agreement between simulated and in-situ observed ground
32 snow albedo in the broadband, visible, and near-infrared spectra is enhanced in Noah-

33 MP/SNICAR simulations relative to Noah-MP/BATS simulations. The SNICAR scheme
34 improves the temporal variability of modeled broadband snow albedo, with a nearly twofold higher
35 correlation with observations ($r=0.66$) than the default BATS snow albedo scheme ($r=0.37$). The
36 underestimated variability in Noah-MP/BATS is a result of inadequate physical linkage between
37 snow albedo and environmental/snowpack conditions, which is substantially improved by the
38 SNICAR scheme. Importantly, the Noah-MP/SNICAR model, with constraints of snow grain size
39 from the MODIS snow covered area and grain size (MODSCAG) satellite data, physically
40 represents and quantifies the snow albedo and absorption of shortwave radiation in response to
41 snow grain size, non-spherical snow shapes, and light-absorbing particles (LAPs). The coupling
42 framework of the Noah-MP/SNICAR model provides a means to reduce the bias in simulating
43 snow albedo.

44

45 **Plain Language Summary**

46 Snow albedo, defined as the proportion of sun light reflected on the snowpack influences
47 snowpack growth, melting rate, surface energy and water cycles, as well as regional and global
48 hydrological and climate change. The community Noah-MP land surface model uses semi-
49 physical snow albedo parameterizations with critical uncertainties. In this study, the integration of
50 the latest version of the Snow, Ice, and Aerosol Radiative (SNICAR) model into Noah-MP aims
51 to enhance the accuracy of snow albedo simulations. The coupled Noah-MP/SNICAR model
52 encompasses various aspects of snow, including its grain properties, aging, and interactions among
53 snow, aerosols, and radiation. We evaluate Noah-MP simulations using SNICAR and the default
54 Biosphere-Atmosphere Transfer Scheme (BATS) snow albedo schemes against albedo
55 observations at three Rocky Mountain stations. The SNICAR scheme enhances accuracy of snow
56 albedo simulations, showing a correlation with observations that is twice as high as simulations
57 using the default BATS snow albedo scheme. The variability of snow albedo in Noah-MP/BATS
58 is underestimated due to the inadequate physical relationship between snow albedo and
59 environmental/snowpack conditions. This issue is addressed and significantly improved by Noah-
60 MP/SNICAR. It is worth noting that the Noah-MP/SNICAR model takes into account the
61 constraints of fresh snow grain size from satellite data.

62

63 **1. Introduction**

64 Snow albedo is defined as the ratio of the snow-reflected solar radiation to the total solar
65 radiation incident on the snowpack. The observed high snow albedo is a result of the considerable
66 portion of the visible spectrum of solar energy on the snow surfaces (Cohen & Rind, 1991). The
67 albedo of snow exerts a substantial influence on multiple facets in the Earth system, including the
68 evolution of the snowpack, rates of melting, surface energy and water cycles, and regional and
69 global hydrological and climate change (Barnett et al., 2005; Flanner et al., 2011; Qian et al., 2015;
70 Skiles et al., 2018; Zhang et al., 2019). This can be attributed to the positive snow albedo feedback
71 and the complex interactions that occur between the land and the atmosphere (Hall, 2004).
72 Variations in snow albedo resulting from temperature warming or cooling can lead to

73 enhancements or reductions in the absorption of solar radiation, hence intensifying the initial
74 warming or cooling (Thackeray & Fletcher, 2016).

75 There is still limitation and uncertainty in the representation of snow albedo processes
76 within many land surface models (LSMs) coupled to regional and global weather and climate
77 models, which consequently affects the estimation of land surface energy and water balances. For
78 example, systematical snow albedo biases were found in one widely-used LSM, the Noah with
79 multi-parameterization options (Noah-MP) (Niu et al., 2011) which is a land component within
80 the Weather Research & Forecasting (WRF) model, the NOAA Unified Forecast System (UFS)
81 model, and the National Water Model (NWM) among many others, as demonstrated by various
82 studies (Abolafia-Rosenzweig et al., 2021, 2022a; Chen et al., 2014; He et al., 2019, 2021; Liu et
83 al., 2021, 2022; Wang et al., 2020; Xiao et al., 2021). The snow albedo in Noah-MP is represented
84 using semi-physical or empirical functions (Verseghy, 2007; Yang et al., 1997), which lack
85 physical treatments of snow albedo response to the evolution of snow properties such as aging and
86 metamorphism. This approach leads to an inconsistent treatment between snow albedo and other
87 snowpack properties (He and Flanner, 2020). To enhance the accuracy of snow albedo modeling,
88 it is necessary to have a comprehensive understanding of the underlying physical mechanisms that
89 influence snow albedo. Subsequently, these processes need to be represented in a more physically
90 realistic manner.

91 The albedo of snow is determined by a complex combination of multiple factors, such as
92 snow depth, the size and shape of snow grains, and the concentration of light-absorbing particles
93 (LAPs) (Warren and Wiscombe, 1980). It has been known that snow albedo is affected by LAPs
94 mainly in the visible band (Warren and Wiscombe, 1980) and by grain size mainly in the near-
95 infrared (NIR) band (Wiscombe and Warren, 1980). Following the occurrence of snowfall, snow
96 crystals experience fast alterations in their size and shape, exhibiting a tendency for snow grains
97 to progressively increase in size over time (Colbeck, 1982). The alteration in snow grain size
98 influences the interaction between the snow surface and incoming solar radiation. The presence of
99 larger grains in snow leads to an increase in the path length traveled by photons, resulting in a
100 lower albedo (Warren 1982). The Snow, Ice, and Aerosol Radiative (SNICAR) model (Flanner et
101 al, 2007, 2021) is one of the most widely used open-source snowpack radiative transfer models,
102 which resolves the aforementioned physical processes and simulates snow albedo by considering
103 snowpack properties such as grain size and shape, as well as environmental conditions including
104 the presence of LAPs.

105 Simulating snow albedo using SNICAR has many advantages compared to the current
106 semi-physical snow albedo schemes, such as the Biosphere-Atmosphere Transfer Scheme (BATS)
107 in Noah-MP (Abolafia-Rosenzweig et al., 2022a; Yang et al., 1997). (1) The study conducted by
108 Abolafia-Rosenzweig et al. (2022a) demonstrated the issue of using a constant parameter to
109 represent the fresh snow albedo in Noah-MP/BATS, which is typically adopted by empirical/semi-
110 physical schemes, whereas the measured albedo of fresh snow exhibits significant variability
111 especially in the NIR band. In contrast, fresh snow albedo is influenced by several environmental
112 conditions and physical processes in the SNICAR model, such as temperature, downward solar
113 spectrum, snow grain size and shapes, LAPs within snow, and the thickness and density of the

114 snowpack. The SNICAR treatment is more physically realistic, as highlighted by Wang et al.
115 (2020). (2) The inclusion of different snow grain shapes, such as spheres, spheroids, hexagonal
116 plates, and Koch snowflakes, is necessary to depict the types of non-spherical snow grains that are
117 commonly observed (Liou et al., 2014; He et al., 2018a, 2023a; Robledano et al., 2023). This
118 representation is currently absent in Noah-MP, while it has been included in SNICAR. (3) The
119 representation of snow aging processes is empirical and incomplete in Noah-MP due to the
120 absence of the simulation of snow grain size. Instead, the simulated snow grain size in SNICAR
121 can be validated using either in-situ measurements or remote-sensing products. This approach
122 offers the advantage of requiring less arbitrary tuning of empirical snow aging parameters which
123 is needed by current Noah-MP snow albedo schemes. (4) SNICAR simulates the interaction
124 between snow, aerosols, and radiation (Flanner et al., 2021; He et al. 2018a; Skiles & Painter,
125 2019), encompassing three LAPs: black carbon (BC), organic carbon (OC), and dust. Additionally,
126 the latest SNICAR coupled into the Community Land Model (CLM5) and the DOE's Energy
127 Exascale Earth System Model (E3SM) Land Model (ELM) also includes the internal mixing of
128 BC and dust with snow grains (He et al., 2023a; Hao et al., 2023). However, these treatments are
129 either missing or not physically represented in Noah-MP. (5) SNICAR presents the effect of solar
130 zenith angle on snow albedo (for direct radiation) physically, while Noah-MP parameterizes this
131 effect semi-empirically such as in BATS. (6) SNICAR computes vertical solar radiation absorption
132 and heating rate for individual snow layers and the top soil layer, which changes snow and soil
133 temperature profiles but is missing in the current Noah-MP albedo schemes. (7) SNICAR has a
134 hyperspectral calculation capability that is more accurate than narrowband calculations (Wang et
135 al., 2022), and can be expanded to incorporate or compare with spectral radiation obtained by
136 remote sensing, while current Noah-MP snow albedo schemes only use two (visible and NIR)
137 bands.

138 Recent studies have implemented SNICAR in some LSMs coupled to global climate
139 models, such as the CLM within Community Earth System Model (CESM) (Flanner et al., 2007;
140 He et al., 2023a) and the DOE's ELM (Hao et al., 2023). This study aims to implement the latest
141 SNICAR version that has several new features and enhancements (He et al., 2023a) into the newly
142 refactored Noah-MP version 5 by accounting for snow grain shape, size, snow-aerosol-radiation
143 interaction, and snow aging processes (Section 2), and to evaluate modeled snow albedo using in
144 situ albedo observations at three Rocky Mountain field stations (Section 3). We use in situ
145 observations of upward/downward shortwave radiation and snow depth and satellite-derived snow
146 grain size to evaluate and constrain the modeled snowpack and radiation conditions. We first
147 evaluate whether the coupled Noah-MP/SNICAR model can accurately reproduce the observed
148 mean and variability of snow albedo, and then perform the process-level model experiments to
149 quantify the effects of snow grain size, shape, and LAPs on snow albedo and radiative forcing
150 (Section 4). We also compare the snow albedo between enhanced Noah-MP/SNICAR simulation
151 and the default Noah-MP/BATS simulation (Section 4). Section 5 discusses the potential
152 uncertainties and future directions, and Section 6 concludes the study.

153

154 **2. Noah-MP and its coupling with SNICAR**

155 **2.1. Noah-MP model description**

156 Noah-MP (Niu et al., 2011; Yang et al., 2011) is one of the most widely used open-source
157 community LSMs worldwide, which has been used in various research and operational models
158 pertaining to weather, climate, and hydrology. The newest version of Noah-MP (version 5.0) (He
159 et al., 2023c) has undergone recent refactoring and incorporating contemporary Fortran code
160 styles, data structures, and standards. This refactoring has significantly improved the model's
161 modularity, interoperability, and applicability (He et al., 2023b).

162 Noah-MP is featured as a multi-parameterization LSM that enables the user to combine
163 different physics schemes for modeling individual land surface processes (Niu et al. 2011). The
164 Noah-MP snow module has the capability to simulate a maximum of three snow layers, with the
165 number of layers being dependent on snow depth. Noah-MP treats explicit snow layers when snow
166 depth is larger than 2.5 cm, and implicitly represents a very shallow (<2.5 cm) snow layer by
167 combining it with the top soil layer in energy and water balance calculations. Snow layer
168 temperature, snow depth, and snow water and ice contents are calculated based on snowpack water
169 and energy balances. The model considers many key snow processes such as snow layer division
170 and combination, liquid water holding within the snowpack, snow compaction, snow melting and
171 refreezing, frost and sublimation at the ground snow surface, and the interception of snow by
172 vegetation. The technical report by He et al. (2023c) provides a comprehensive description of the
173 various aspects related to snowpack mass and energy processes.

174 Within Noah-MP, there exist two semi-physical snow albedo schemes, namely CLASS
175 (Verseghy, 2007) and BATS (Yang et al., 1997). The mathematical equations of the two snow
176 albedo schemes are described in detail in He et al. (2023c). Both schemes simulate snow albedo in
177 the visible and NIR bands under direct and diffuse radiation, but CLASS assumes the same snow
178 albedo for direct and diffuse radiation as well as visible and NIR bands, which is physically
179 unrealistic. Furthermore, both schemes do not explicitly simulate the evolution of snow properties
180 (e.g., snow aging/metamorphism, grain size, and shape). A recent study (Abolafia-Rosenzweig et
181 al., 2022a, 2022b) has tried to optimize the BATS albedo parameters using in-situ snow albedo
182 measurements, which however still showed nontrivial remaining biases particularly for fresh snow
183 albedo due to a lack of physical representation of relevant albedo processes.

184

185 **2.2 Noah-MP/SNICAR coupling**

186 In this study, we couple the refactored Noah-MP version 5 with the latest version of
187 SNICAR (<https://github.com/ESCOMP/CTSM/pull/1861>) that has recently been implemented
188 into CLM5 (He et al., 2023a).

189 **2.2.1 Multiple physics options for SNICAR albedo calculations**

190 The SNICAR scheme we implement into Noah-MP incorporates several key physical
191 processes and updates following He et al. (2023a): (1) two options for radiative transfer solvers,
192 with one for a traditional tri-diagonal matrix two-stream solver (Toon et al., 1989) and one for a
193 new adding-doubling solver (Dang et al., 2019); (2) three options for ice optical properties (Flanner

194 et al., 2021) using different ice refractive indices from Warren (1984), Warren and Brandt (2008),
195 and Picard et al. (2016); (3) updated aerosol optical properties of BC and OC from Flanner et al.
196 (2021); (4) six options of downward solar spectra for multiple atmospheric conditions (Flanner et
197 al., 2021), including mid-latitude winter, mid-latitude summer, sub-Arctic winter, sub-Arctic
198 summer, Summit Greenland, and high mountain; (5) four types of snow grain shapes including
199 sphere, spheroid, hexagonal, and snowflake (He et al., 2017); (6) three dust types including
200 Saharan dust, Colorado dust, and Greenland dust (Flanner et al., 2021); (7) two options for either
201 internal or external mixing of dust (He et al., 2019b) or BC (He et al., 2017) with snow grains; (8)
202 two options for wavelength band setup, including 5-band and hyperspectral (480-band with 10-nm
203 spectral resolution) capabilities. Both 5-band and 480-band albedo results are then averaged to two
204 (visible and NIR) bands' values to be used in Noah-MP surface energy flux calculations. We
205 implement all these SNICAR albedo calculation processes and the ability of choosing different
206 physics options to simulate each individual processes into Noah-MP.

207 **2.2.2 Model inputs**

208 We incorporate additional input datasets that are required for SNICAR (Flanner et al.,
209 2021; He et al., 2023a) through an updated model I/O interface. SNICAR requires the input
210 variables including direct/diffuse radiation, surface downward solar spectrum, solar zenith angle
211 (only for direct radiation), albedo of the surface underlying snowpack, vertical profiles of snow
212 grain size, snow layer thickness, snow density, and mass concentrations of LAPs (BC, mineral
213 dust, and OC), snow grain shape, and optical properties of ice and LAPs. The optical properties of
214 ice and LAPs for each snow layer and spectral bands, including single-scattering albedo, mass
215 extinction cross section, and asymmetry parameter, are archived as look-up table datasets derived
216 by Flanner et al. (2021) and He et al. (2023a).

217 **2.2.3 Snow grain size and aging processes**

218 The evolution of snow effective grain size is represented by snow aging processes and
219 implemented into the Noah-MP/SNICAR model. The change in effective snow grain size is based
220 on the dry and wet snow processes, including liquid-water-induced metamorphism, dry snow
221 metamorphism, refreezing of liquid water, and the addition of freshly fallen snow (Flanner et al.,
222 2007; Lawrence et al., 2019). The liquid-water-induced metamorphism is parameterized based on
223 measured grain growth rates under different liquid water contents (Brun, 1989). The dry snow
224 metamorphism is determined by snow temperature, temperature gradient, density, and initial snow
225 grain size distribution based on a microphysical particle model that simulates diffusive vapor flux
226 amongst collections of ice crystals with various size and inter-particle spacing (Flanner and
227 Zender, 2006). This process reproduces the typical observed rapid snow aging and increased snow
228 grain size under the conditions of combined warm snow, large temperature gradient, and low
229 density. The effective radius of refrozen liquid water is set to 1000 μm (Oleson et al., 2013). The
230 air temperature is a key factor in determining the grain size of freshly fallen snow. At temperatures
231 below -30 degrees Celsius, a minimum of 54.5 μm (radius) is imposed (Lawrence et al., 2019). A
232 limit is imposed on the maximum of 204.5 μm (radius) when the temperature exceeds 0 degrees
233 Celsius. A linear ramp is employed within the temperature range between -30 to 0 degrees Celsius
234 (Lawrence et al., 2019). These maximum and minimum limits are tunable parameters. In our

235 investigation, we discover that grain size and snow albedo are sensitive to the minimum and
236 maximum values of the grain size of freshly fallen snow. In Section 3.3.3, we have optimized these
237 two parameters to match with the snow grain size acquired from a satellite product (see Section
238 3.2). In situations where there is a non-zero snow mass but an explicit snow layer has not yet been
239 established (i.e., snow depth <2.5 cm), the effective snow grain size is assigned as the effective
240 radius of freshly fallen snow (Lawrence et al., 2019). When snow layers are combined or divided,
241 the effective snow grain size is calculated as a mass-weighted mean of those of the two layers.
242 Lastly, the effective snow grain size is limited to a range of 30-1500 μm , as this range covers the
243 majority of snow grain size in reality and corresponds to the defined optical properties that are
244 archived in look-up tables (Flanner et al., 2021; He et al., 2023a).

245 **2.2.4 Snow-aerosol-radiation interactions**

246 Additionally, we implement a mass-conserving approach to account for the presence of
247 LAPs within snow, encompassing the mechanisms of atmospheric aerosol deposition on the
248 uppermost snow layer, aerosol mass reduction via inter-layer meltwater drainage, and aerosol mass
249 changes due to snow layer combination and subdivision (Flanner et al., 2007; Lawrence et al.,
250 2019). The Noah-MP/SNICAR model tracks the mass of nine aerosol particle species within each
251 snow layer including hydrophilic BC, hydrophobic BC, hydrophilic OC, hydrophobic OC, and
252 mineral dust with five particle size bins (μm in diameter, Table S2): 0.1-1.0, 1.0-2.5, 2.5-5.0, 5.0-
253 10.0, and 10.0-100.0 (Flanner et al., 2021). Each species exhibits distinct optical characteristics
254 (Flanner et al., 2021; He et al., 2023a) and meltwater removal efficiencies (Lawrence et al., 2019).

255 **2.2.5 Albedo output and snowpack heating**

256 The results simulated from the Noah-MP/SNICAR model include the spectral snow albedo
257 and the fraction of solar flux that is absorbed by each individual snow layer and the top soil layer.
258 The spectral snow albedo is partitioned into visible and NIR bands by computing the spectrally
259 weighted mean based on downward solar spectra (Flanner et al., 2007). The layer-wise snowpack
260 heating due to snow and LAPs absorption of solar radiation from SNICAR is coupled with Noah-
261 MP snow and soil temperature computations to alter the temperature for each snow layer and the
262 underlying top soil layer.

263

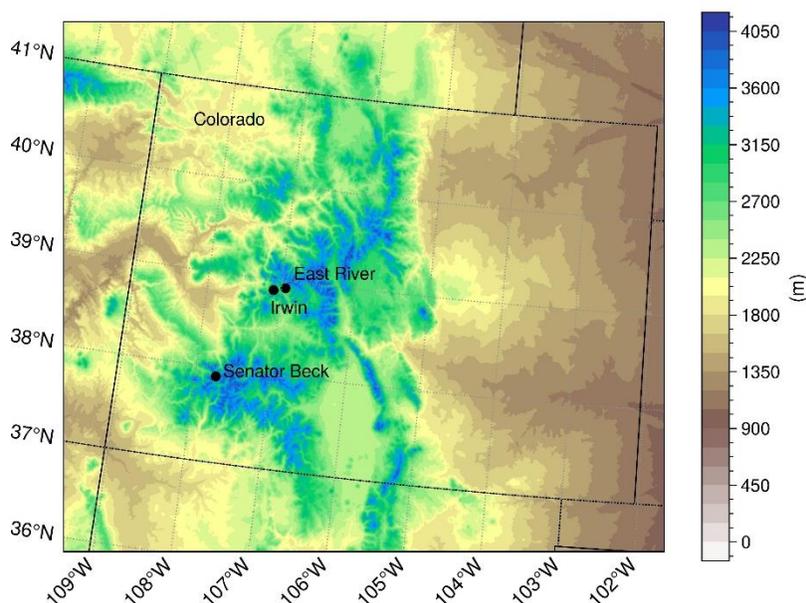
264 **3. Model experiments and evaluation data**

265 **3.1 In-situ observations**

266 In-situ observations of snow albedo and snow depth data are obtained from three high-
267 elevation locations, East River, Irwin, and Senator Beck, within the southern Rocky Mountains in
268 the state of the Colorado, United States (Figure 1; Abolafia-Rosenzweig et al., 2022a, 2022b). The
269 longitudes, latitudes, elevation, vegetation types, available observed spectrum bands of snow
270 albedo (broadband, visible, and NIR), and atmospheric forcing variables for each site are provided
271 in Table S1. The comprehensive methodologies for measuring solar radiation, albedo and snow
272 depth can be found in Abolafia-Rosenzweig et al. (2022a).

273 The snow albedo and snow depth measurements have undergone rigorous quality control
 274 procedures to ensure their accuracy and reliability. To eliminate the effect of low sun angle on
 275 albedo, we use averages of albedo and snow depth measured between 11:00 and 13:00 local time
 276 for the three sites during the periods of investigation. The analysis is further limited to periods
 277 when the observed snow depth is more than 0.5 m at the East River site and 0.2 m in the Senator
 278 Beck and Irwin sites, in order to ensure that understory vegetation is completely buried by
 279 snowpack to eliminate the influence of vegetation on snow albedo, following Abolafia-
 280 Rosenzweig et al. (2022a). Observed albedo values that exceed 1.0 or fall below 0.0 are removed.

281



282

283 **Figure 1.** Locations of the three study sites with topography.

284

285 3.2 MODIS snow covered area and grain size (MODSCAG) product

286 We use a daily 463-m MODSCAG product (Painter et al, 2009) to evaluate and optimize
 287 modeled snow grain size. Based on spectral unmixing and physically based snow radiative transfer
 288 models that remove soil and vegetation portions of the observed pixel, MODSCAG provides snow
 289 grain size at roughly 10:30 LST (local solar time). For a clear sky day, the mean absolute error
 290 (MAE) for snow grain size from MODSCAG compared to field measurements at a single site is
 291 51 μm (Painter et al., 2009). For both clear and cloud sky days, the gap-filled MODSCAG has a
 292 root mean square error (RMSE) of 118 μm for snow grain size compared to observations at three
 293 sites in the western United States (Bair et al., 2019). The MODSCAG data are obtained from the
 294 University of California, Santa Barbara (UCSB) website at
 295 <https://snow.ucsb.edu/products/MODSCAG/WUS/>. The snow grain size values are extracted from
 296 the encompassing MODSCAG grid cell of the three field sites studied in this work.

297

298 **3.3 Site-specific model setup and simulations**

299 Noah-MP simulations adopt model-physics settings from the options used in continental-
300 scale convection-permitting WRF/Noah-MP climate simulations (He et al., 2019a; Liu et al., 2017;
301 Rasmussen et al., 2023) that have reasonably captured key land surface states and fluxes over the
302 continental U.S., except for different snow albedo options tested in this study. The snow-related
303 parameters within Noah-MP follows the values used in the latest release of Noah-MP version 5.0
304 (He et al., 2023c). Leaf area index (LAI) is characterized by vegetation type based on a 10-yr
305 monthly climatology of MODIS products (Yang et al., 2011). The vegetation type for each
306 research site is grassland, and the canopy height is set as documented at each specific site
307 (Abolafia-Rosenzweig et al., 2022a). For each study site, Noah-MP is first spun up for 11-13 years
308 to get the steady-state as listed in Table S1, followed by model analysis for subsequent years
309 (Abolafia-Rosenzweig et al., 2022a). The analysis period ranges between October 2018 to August
310 2021 in Irwin, October 2011 to 2020 October in Senator Beck, and July 2017 to November 2019
311 in East River.

312 **3.3.1 Atmospheric forcing**

313 The Noah-MP simulations utilize atmospheric forcing derived from a combination of two
314 sources: the hourly forcing data obtained from the 1-km observation-constrained NOAA's
315 Analysis of Record for Calibration data set (AORC; National Weather Service, 2021) which are
316 then replaced with in-situ observed data when accessible from each study site. The atmospheric
317 forcing variables observed at each location are listed in Table S1. To minimize simulation
318 uncertainty due to downward direct/diffuse shortwave radiation in visible and NIR bands, we use
319 both the observed total downward shortwave radiation and the observed fraction of direct/diffuse
320 and visible/NIR radiation when observational data are available.

321 **3.3.2 Aerosol deposition flux**

322 All model simulations are driven by the hourly aerosol (BC, dust, and OC) wet and dry
323 deposition fluxes from the MERRA-2 reanalysis (Randles et al., 2017). MERRA-2 provides the
324 wet and dry deposition fluxes for hydrophobic and hydrophilic (aged) OC and BC as well as dust
325 with 5 size bins at a spatial resolution of $0.625^{\circ} \times 0.5^{\circ}$. The aerosol deposition fluxes at three in-
326 situ locations are determined by spatial interpolation based on the nearest neighbor grids of
327 MERRA-2 values. Additionally, the sizes of MERRA-2 dust aerosol are converted to size bins that
328 are compatible with those in Noah-MP/SNICAR (Table S2).

329 **3.3.3 Fresh snow grain size optimization**

330 We optimize the tunable minimum and maximum values of freshly fallen snow grain size
331 in SNICAR by comparing the simulated snow grain size to the MODSCAG data (Section 3.2). We
332 define the fresh snow cases based on the following criteria: 1) a daily observed increment of snow
333 depth exceeding 0.02 m following Abolafia-Rosenzweig et al. (2022a); 2) a precipitation amount
334 surpassing 0.0 mm/day following Wang et al. (2020); and 3) the model simulation indicates a
335 complete (100%) snow cover fraction. The original values are 54.526 μm for the minimum and

336 204.526 μm for the maximum. In this study, the minimal value is optimized to 33.0 μm , which is
 337 determined by identifying the smallest snow grain size from the MODSCAG data across all study
 338 sites and study periods. The maximum value is optimized to 91.0 μm to match modeled average
 339 snow grain size with the MODSCAG averaged value over all study sites and study periods.

340 3.3.4 Model experiments

341 We conduct five model experiments, as shown in Table 1. The first simulation (Exp1)
 342 serves as the baseline case using the SNICAR snow albedo scheme. The SNICAR configuration
 343 includes the utilization of an adding-doubling solver, a midlatitude winter atmosphere profile with
 344 five wavelength bands, a hexagonal snow grain shape, the Colorado dust type, ice optical
 345 properties obtained from Picard et al. (2016), the inclusion of BC, dust, and OC, and the internal
 346 mixing of BC and dust with snow grains. Additionally, the simulation incorporates optimized
 347 parameters with respect to the fresh snow grain size mentioned in Section 3.3.3. Subsequently, we
 348 perform four sensitivity simulations (Exp2-5) to comprehend the impacts of snow grain size, snow
 349 grain shapes, and LAPs on snow albedo and surface radiative balance. The second simulation
 350 (Exp2) is identical to Exp1, except that it employs the original SNICAR parameters for fresh snow
 351 grain size. The third simulation (Exp3) is identical to Exp1, except for the modification of the
 352 hexagonal snow grain shape to the spherical shape. The fourth simulation (Exp4) is identical to
 353 Exp1, except that it does not account for the influence of LAPs in snow. The final simulation
 354 (Exp5) utilizes the default Noah-MP BATS snow albedo scheme instead of SNICAR. All
 355 simulations are forced with the in-situ observed snow depth to reduce the albedo uncertainty
 356 introduced by snow depth bias, as previous studies have shown nontrivial snow depth bias
 357 simulated by Noah-MP (Abolafia-Rosenzweig et al., 2021; Chen et al., 2014; He et al., 2019a;
 358 2021; Ikeda et al., 2021). Specifically, the observed snow depth is directly inserted to ensure the
 359 simulated snow depth aligns with the observed values during periods when observations are
 360 available (Figure S1). The snow depth is subsequently transformed into snow water and ice by
 361 multiplying with the modeled snow density, as typically done in snow depth data assimilation
 362 procedures. We note that this may introduce uncertainty to snow water equivalent in the model
 363 due to the lack of direct observations of snow density.

364

365 **Table 1.** Noah-MP model configurations for different experiments.

Experiments	Snow albedo scheme	Snow grain size	Snow shape	Snow impurities
Exp1 (baseline)	SNICAR	Optimized	Hexagonal	w/ LAPs
Exp2	Same as Exp1	Original	Same as Exp1	Same as Exp1
Exp3	Same as Exp1	Same as Exp1	Sphere	Same as Exp1
Exp4	Same as Exp1	Same as Exp1	Same as Exp1	w/o LAPs
Exp5	BATS	-	-	-

366

367 3.4 Evaluation metrics

368 Statistical metrics are computed to assess the performance between simulated and observed
 369 snow albedo. Bias is employed as a metric to assess the extent to which the modeled albedo is
 370 capable of properly replicating the average condition of the observed values. The root mean square
 371 error (RMSE) is employed as a metric to assess the accuracy of the model capability in estimating
 372 the observed value. The Pearson's correlation coefficient (r) is used to quantify the temporal
 373 correlation of modeled and observed snow albedo.

374

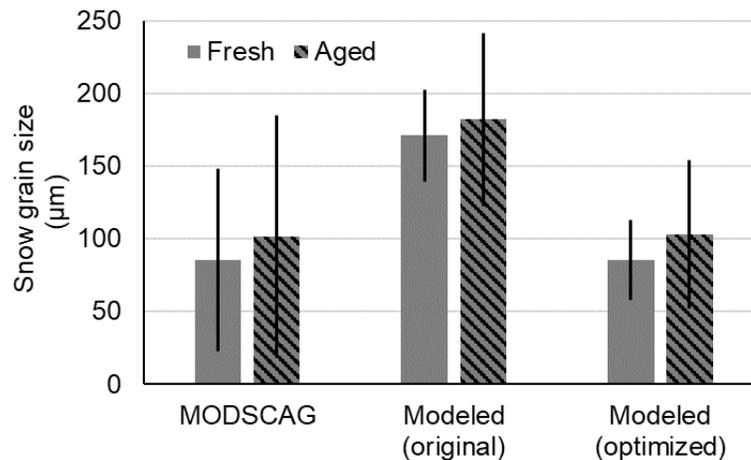
375 4. Results

376 4.1 Model evaluation

377 4.1.1 Snow grain size

378 The Noah-MP/SNICAR simulation using original SNICAR grain size parameters (i.e.,
 379 Exp2) produces systematically higher fresh and aged snow grain sizes by about two times
 380 compared to MODSCAG (Figure 2). This discrepancy can be attributed to the overestimate of
 381 fresh snow grain size. When optimizing the size parameters of the freshly fallen snow grains (i.e.,
 382 Exp1; see Section 3.3.3), the modeled average fresh snow grain sizes (85 μm) at three research
 383 sites agree very well with the values (85 μm) obtained from MODSCAG (Figure 2). Additionally,
 384 our results demonstrate that by improving the fresh snow grain size simulation in Noah-
 385 MP/SNICAR, we can further effectively decrease the bias in the simulated aged (non-fresh) snow
 386 grain size. The mean value of the modeled aged snow grain size decreases from 182 to 103 μm ,
 387 which matches very well with the observed value of 101 μm (Figure 2).

388



389

390 **Figure 2.** Comparison of the average fresh and aged snow grain size obtained from MODSCAG,
 391 and the simulated results from Noah-MP/SNICAR simulations with the original and optimized
 392 fresh snow grain size parameters at three study sites during the periods when MODSCAG data is
 393 accessible. The uncertainty bar represents the spatiotemporal variability of snow grain size within
 394 one standard deviation. The fresh snow is defined in Section 3.3.3.

395

396 **4.1.2 Snow albedo during all periods**

397 Overall, the Noah-MP/SNICAR baseline simulation (i.e., Exp1) captures the observed
398 snow albedo values, but with a higher broadband albedo by about 0.072 and less temporal
399 variability (Figure 3a-c). The broadband overestimates mainly arise from the overestimated visible
400 snow albedo by about 0.086, likely caused by the uncertainty in aerosol deposition and/or snow
401 density, because snow grain size and snow depth are constrained by observations. This also
402 explains the good agreement between modeled and observed mean NIR snow albedo (Figure 3a-
403 c), since NIR snow albedo is sensitive to grain size. Uncertainty in snow grain shape could also
404 slightly (by up to ~0.02) contribute to the overestimated visible snow albedo based on sensitivity
405 analysis (Section 4.2.2). The missing treatment of small-scale snow surface roughness in the model
406 could also contribute to the snow albedo overestimates (Manninen et al., 2021), but it generally
407 has a stronger impact on NIR albedo and hence may not be the main culprit here. Nevertheless,
408 the model reproduces the observed pattern of the visible albedo larger than the NIR albedo. The
409 underestimated temporal variability of snow albedo at both visible and NIR bands is partially
410 caused by the underestimated variability of snow grain size (Figure 2), particularly during ablation
411 periods (Figure 3g-i). This is mainly due to the uncertainty in snow aging processes, which are less
412 constrained by observations. The uncertainty in aerosol deposition and evolution in snow could
413 also contribute to the underestimated visible albedo temporal variability because the visible snow
414 albedo is more sensitive to snow impurity than snow grain size (Section 4.2.3).

415 **4.1.3 Fresh snow albedo**

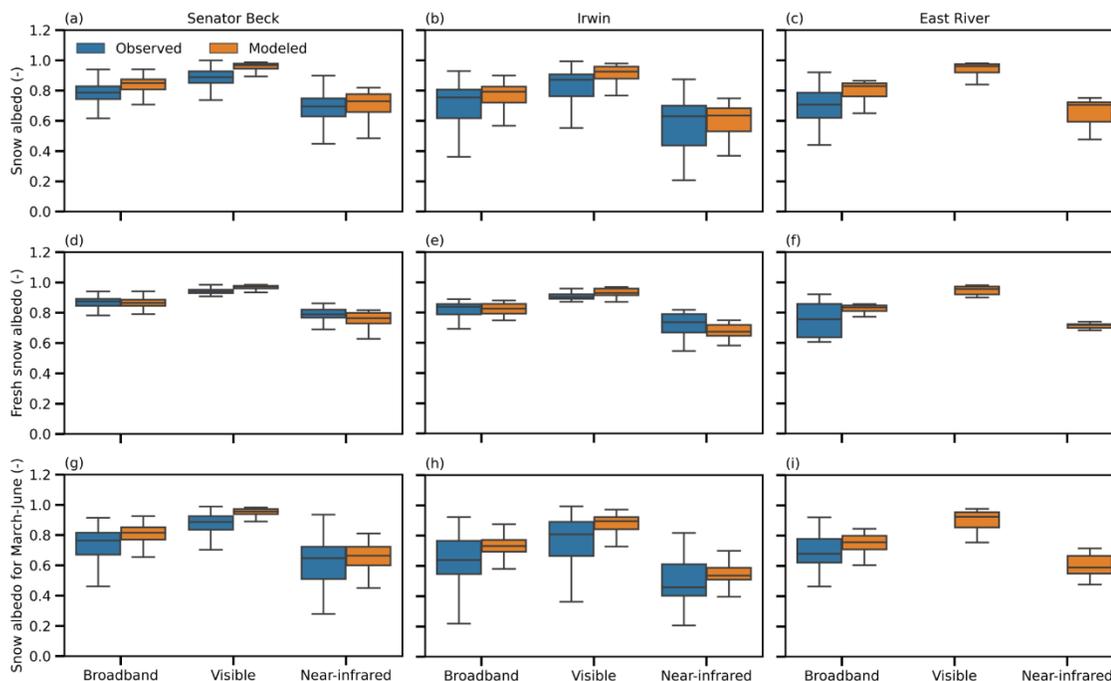
416 The Noah-MP/SNICAR baseline simulation of broadband fresh-snow albedo reproduces
417 the mean and variability of observations due to the well-captured fresh snow grain size (Figure 2),
418 with higher accuracy in the Senator Beck and Irwin sites than East River sites (Figure 3d-f). The
419 simulated median broadband value closely matches the value of observed fresh-snow albedo at the
420 Senator Beck site (0.88 observed vs. 0.87 modeled) and the Irwin site (0.84 observed vs. 0.83
421 modeled). At the East River site, the modeled median value (0.83) is higher than the observed
422 values (0.76) with underestimated temporal variability. For the visible band, the median fresh snow
423 albedo is slightly overestimated by 0.03 at both Senator Beck and Irwin sites (Figure 3d-e). For
424 the NIR band, the median fresh snow albedo is underestimated by about 0.03 at the Senator Beck
425 site and about 0.07 at the Irwin site. Thus, low broadband biases at Irwin and Senator Beck are
426 attributable to compensatory errors in visible and NIR bands.

427 **4.1.4 Snow albedo during melting periods**

428 We evaluate snow albedo during the melting period, which is delineated as the time
429 spanning from March to June (Figure 3g-i). As snow melts, its albedo decreases with increased
430 temporal variability in comparison to fresh snow albedo. The simulated snow albedo generally
431 captures the observations during melting periods at broadband, visible, and NIR bands, with a
432 similar bias pattern as that of the entire snow period (Figure 3a-c). Specifically, the overestimated
433 broadband albedo (by 0.066) is dominated by the overestimate in the visible band (mean bias =
434 0.093), with NIR albedo better simulated (mean bias = 0.043). This is likely due to the uncertainty

435 in aerosol content in snow, snow density, and/or snow grain shape as discussed in Sections 4.1.2
 436 and 5. The underestimated temporal variability of snow albedo at all bands can be explained by
 437 the uncertainty in snow aging processes and aerosol content in snow as mentioned in Section 4.1.2.

438



439

440 **Figure 3.** Site-level comparisons of snow albedo (a-c), fresh snow albedo (d-f), and snow albedo
 441 over the melt period (March-June) (g-i) from observations (Observed) and Noah-MP/SNICAR
 442 simulations (Modeled) at the Senator Beck (left panels), Irwin (middle panels), and East River
 443 (right panels) stations. The boxes are the interquartile ranges, the horizontal lines plotted in the
 444 boxes are the median values, and the whiskers indicate the maximum and minimum values of the
 445 results.

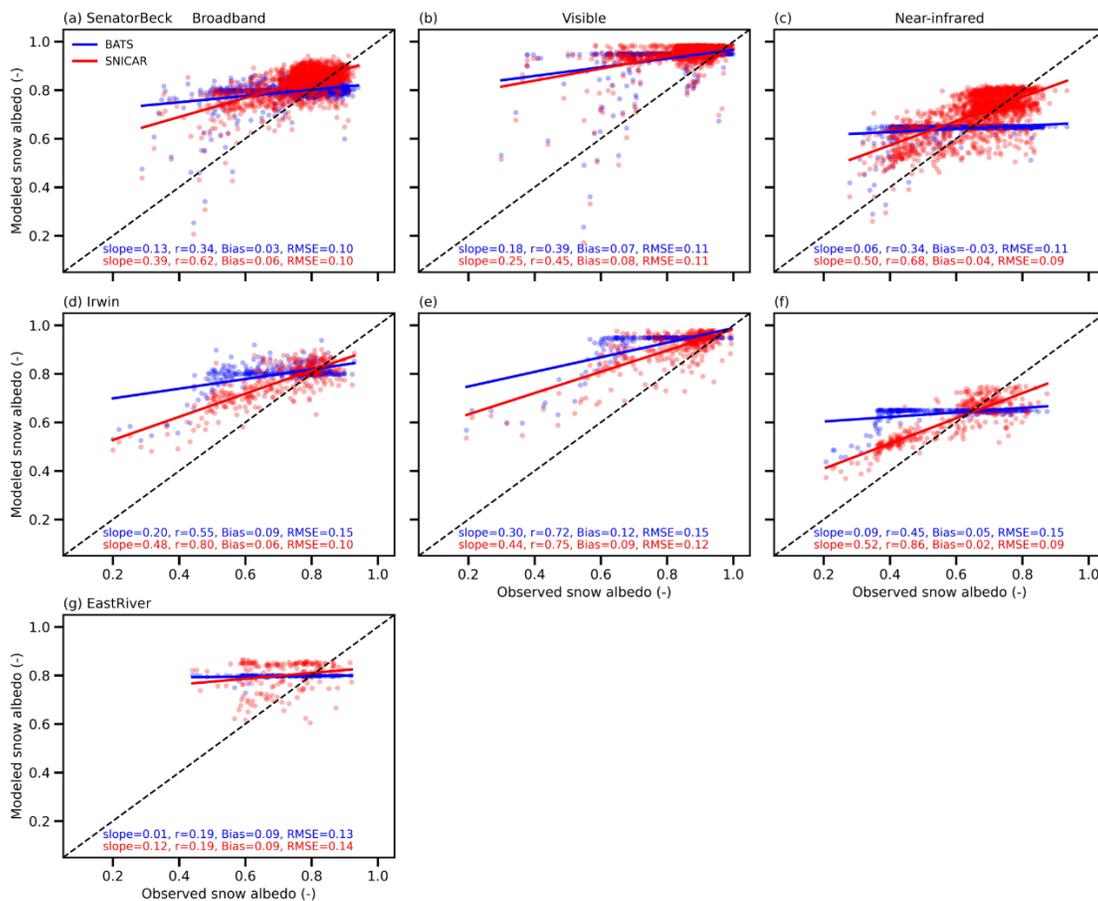
446

447 4.1.5 Comparison with default Noah-MP/BATS snow albedo scheme

448 We further compare the Noah-MP/SNICAR simulation with the Noah-MP simulation
 449 using the default semi-physical BATS snow albedo scheme that has been recently optimized by
 450 Abolafia-Rosenzweig et al. (2022a). Overall, the Noah-MP/SNICAR results outperform those of
 451 Noah-MP/BATS at all three sites (Figure 4). The SNICAR scheme improves the temporal
 452 variation (*slope* in the scatter plots) and correlation (*r* in scatter plots) with the observations for
 453 snow albedo at all bands. We note that the underestimated variability in the Noah-MP/BATS snow
 454 albedo suggests inadequate physical linkage and sensitivity between snow albedo and
 455 environmental/snowpack conditions in the BATS scheme, which is substantially improved by the
 456 SNICAR scheme. In terms of mean *Bias* and *RMSE*, there are few variations between the two
 457 simulations for the Senator Beck and East River sites, while the Irwin site shows a significant

458 improvement (about 50%). Noah-MP/SNICAR improves the issue of conditional bias existing in
 459 Noah-MP/BATS, i.e., the tendency of underestimating high albedo values and overestimating low
 460 albedo values. SNICAR enhances the variability in the visible snow albedo, which mitigates the
 461 overestimate of visible snow albedo in BATS. In the Noah-MP/BATS simulation, the visible snow
 462 albedo is consistently around 0.9 (blue dots in Figure 4b and 4e), which is not realistic. This could
 463 be because the BATS scheme uses a fixed parameter for fresh snow albedo (Abolafia-Rosenzweig
 464 et al., 2022a; Wang et al., 2020). However, in the Noah-MP/SNICAR simulation, the fresh snow
 465 albedo is dynamically dependent on environmental conditions such as changes in temperature,
 466 snow depth, snow grain size, and the concentrations of LAPs. Furthermore, the simulation of NIR
 467 snow albedo is significantly improved by Noah-MP/SNICAR relative to Noah-MP/BATS, leading
 468 to a notable decrease in the variability bias and bringing the simulated values much closer to the
 469 observed ones.

470



471

472 **Figure 4.** Scatter plots comparing observed, Noah-MP/BATS (blue dots), and Noah-
 473 MP/SNICAR (red dots) simulated ground snow albedo in broadband (a, d, and g), visible (b and
 474 e), and near-infrared (c and f) wavelengths at the Senator Beck (top panels), Irwin (middle
 475 panels), and East River (bottom panels) stations.

476

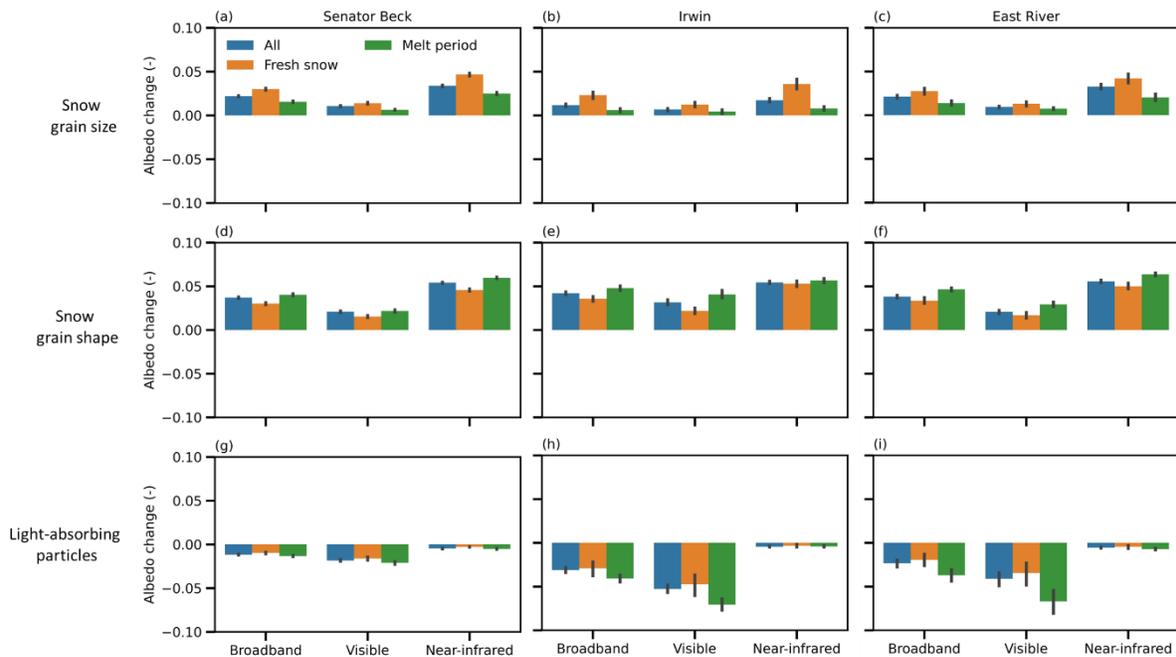
477 4.2 Effects of snow grain size, snow shape, and LAPs on albedo and radiative forcing

478 Here, we quantify the modeled snow albedo and absorbed solar radiation in response to
 479 key snow albedo factors in Noah-MP/SNICAR simulations.

480 4.2.1 Snow grain size

481 The optimization of fresh snow grain size parameters (Section 3.3.3) leads to a decrease in
 482 snow grain size, which better agrees with observations (Figure 2) and in turn increases snow albedo
 483 (Figure 5a-c). The broadband snow albedo at the Senator Beck, Irwin, and East River sites
 484 increases on average by 0.022, 0.012, and 0.021, respectively. The albedo changes induce surface
 485 radiative forcing (SRF) values of -14.1, -6.4, and -12.8 W m^{-2} (Figure 6a-c). The changes in snow
 486 grain size have a more pronounced impact on the NIR band compared to the visible band, which
 487 is consistent with previous studies showing higher NIR snow albedo sensitivity to snow grain size
 488 (e.g., Flanner et al., 2021). As a result, there are greater fluctuations in the SRF in the NIR band,
 489 leading to a decrease in the absorbed broadband solar radiation. In addition, the albedo and SRF
 490 changes induced by snow grain size changes are more pronounced for fresh snow compared to
 491 those in the melting period, mostly due to alterations in the fresh snow grain size by the parameter
 492 optimization.

493



494

495 **Figure 5.** Changes in snow albedo due to changes in snow grain size from original fresh snow
 496 grain parameters to optimized ones (a-c), snow grain shape from sphere to hexagonal shape (d-f),
 497 and light-absorbing particles (LAPs) from no LAPs to with LAPs (g-i) in three stations, Senator
 498 Beck (left panels), Irwin (middle panels), and East River (right panels). The error bars represent
 499 the range of values within one standard deviation of temporal variability. The color of the plots

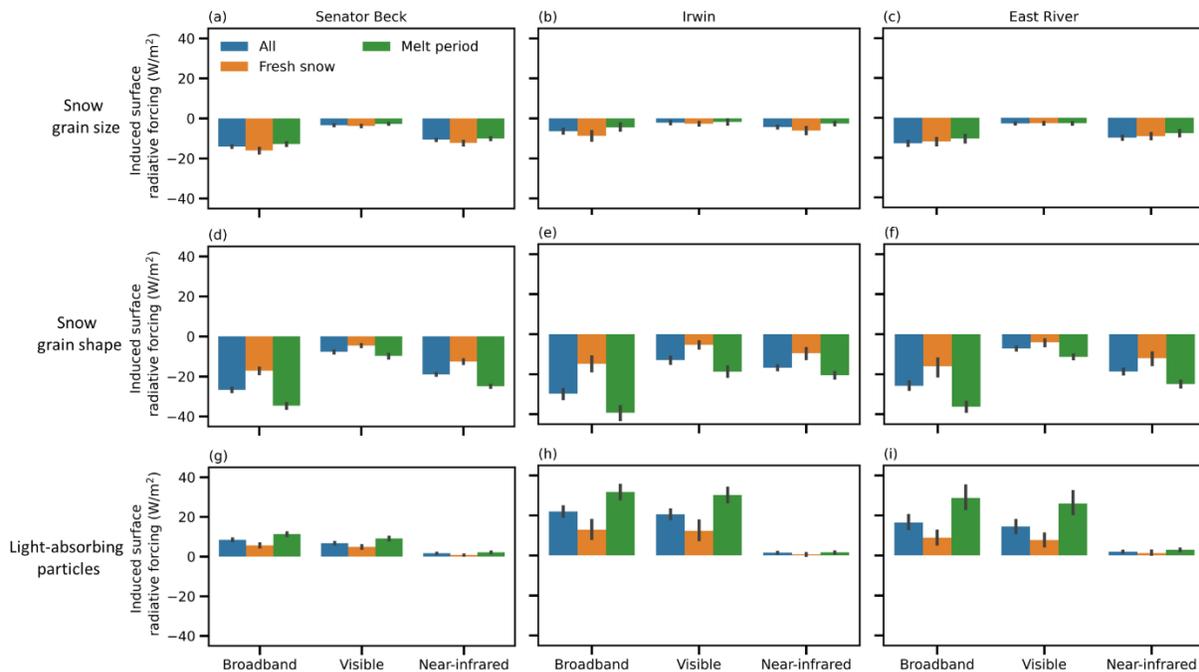
500 represents the data for the entire snow season (All), as well as the cases for fresh snow and the
 501 melting period (March-June).

502

503 4.2.2 Snow grain shape

504 In contrast to a spherical shape, a hexagonal grain shape exhibits a greater snow albedo
 505 (Figure 5d-f) and a lower SRF (Figure 6d-f). This is because non-spherical grains have a smaller
 506 asymmetry factor and weaker forward scattering compared to their spherical counterparts (Dang
 507 et al., 2016; He et al., 2017; 2018b), which is more representative of real-world conditions (Flanner
 508 et al., 2021; Hao et al., 2023; He et al., 2023a). The broadband snow albedo in the Senator Beck,
 509 Irwin, and East River stations increases by an average of 0.037, 0.042, and 0.038, respectively.
 510 This increase in albedo results in changes in surface solar radiation absorption of -26.7, -29.8, and
 511 -25.8 W m^{-2} . During the melting period, the influence of the snow non-sphericity has a greater
 512 impact on snow albedo and SRF compared to the time when the snow is fresh, due to the larger
 513 snow grain size and shallower snowpack during melting periods (He et al., 2018b; He, 2022).
 514 Although the rise in albedo and the decrease in SRF occur in both the visible and NIR bands, it is
 515 more pronounced in the NIR band, because the NIR albedo is more sensitive to snow grain shape
 516 (Dang et al., 2016; Flanner et al., 2021).

517



518

519 **Figure 6.** Similar to Figure 5, but for the induced surface radiative forcing (SRF).

520

521 4.2.3 Light-absorbing particles

522 Figures 5g-i and 6g-i display the changes in snow albedo and SRF caused by the LAPs,
523 respectively. Overall, including LAPs in snow causes a decrease in the broadband snow albedo by
524 an average of -0.012 at the Senator Beck station. This reduction in albedo results in a SRF of 8.4
525 W m^{-2} . The Irwin and East River sites exhibit greater changes in snow albedo and SRF compared
526 to the Senator Beck site, mostly due to the higher concentrations of LAPs present in these two
527 locations. The LAPs-induced SRF at the Irwin station is 21.9 W m^{-2} , while in the Easter River
528 station it is 16.4 W m^{-2} . These changes correspond to a decrease in the broadband snow albedo of
529 -0.031 in the Irwin station and -0.023 in the Easter River station. The effects of LAPs are much
530 more pronounced in the visible band than the NIR band, consistent with literature (e.g., Warren
531 and Wiscombe, 1980; Flanner et al., 2007). The melting period exhibits greater LAPs-induced
532 changes in snow albedo and SRF compared to the fresh snow period, because of larger snow grain
533 sizes and higher snowpack density during melting periods (Flanner et al., 2021; He, 2022) as well
534 as the enrichment of LAPs as snow melts (Niu et al., 2017). In addition, the higher downward solar
535 radiation during melting periods also contributes to the higher SRF compared to winter.

536

537 **5. Uncertainty discussions and future directions**

538 Snow grain shape, size, and snow LAPs all contribute to the potential uncertainty in snow
539 albedo calculations and solar radiation processes. In the three study sites, the shape of snow grains
540 has a considerable impact on snow albedo over the whole snow season. However, because there is
541 no model process that accounts for the dynamic evolution of snow grain morphologies and no
542 direct observational constraints, the model's assumption of non-spherical shape throughout the
543 period is uncertain (He et al., 2023a). In reality, the shape of snow grains demonstrates
544 geographical variation and temporal variability, which necessitates additional refinements (Hao et
545 al., 2023). During the melting phase, the snow albedo biases and the effects of LAPs on snow
546 albedo are stronger than during the accumulation period (i.e., fresh snow). Nonetheless, the coarse
547 resolution of the MERRA-2 aerosol deposition data is accompanied by uncertainty. Furthermore,
548 the LAPs associated with snow are influenced by tunable model parameters such as snow aging
549 scaling factor and inter-layer melt-water scavenging efficiency factor, both of which affect the size
550 evolution of snow grains and the concentrations of LAPs within the snow through positive
551 feedback mechanisms (Qian et al., 2014). Because of the lack of direct observed data, these model
552 parameters are poorly constrained and warrant further exploration to reduce uncertainty in
553 calculating the interactive effects of grain size and snow LAPs on snow albedo particularly during
554 melting period.

555 Furthermore, it is important to acknowledge that the input data, such as atmospheric
556 forcing, are fundamental yet unavoidable sources of uncertainties. We strive to utilize in-situ
557 observed forcing data to the greatest extent possible in order to decrease the level of uncertainty.
558 In Noah-MP, certain snowpack physical processes, such as densification, still have uncertainties
559 (e.g., He et al., 2019, 2021), which may contribute to the bias in the estimation of snow albedo. To
560 mitigate this uncertainty, we used observed snow depth data to constrain model simulations.
561 Looking beyond this study, we plan to evaluate over a larger study domain and conduct regional
562 simulations in a coupled land-atmosphere modeling system to assess the feedback induced by the

563 enhanced SNICAR snow albedo scheme, such as the western United States that experience burning
564 or/and regular dust-on-snow events (e.g., Gleason et al., 2019; Skiles et al., 2015).

565

566 **6. Conclusions**

567 We integrated the widely-used state-of-the-art snow albedo model, the latest version of
568 SNICAR, into the refactored Noah-MP version 5, and evaluated in detail using ground
569 measurements at three Rocky Mountain observation sites. The coupled Noah-MP/SNICAR model
570 physically accounts for the aerosol-snow-radiation interaction, snow grain growth and aging, and
571 effects of snow grain size and shape on snow albedo. The Noah-MP/SNICAR simulation well
572 reproduces the observed broadband, visible, and NIR snow albedo, although it slightly
573 overestimates the visible and broadband snow albedo. The SNICAR scheme significantly
574 improves the temporal variability of snow albedo (particularly in the NIR band) comparing to the
575 semi-physical BATS snow albedo scheme in Noah-MP. The remaining bias in Noah-MP/SNICAR
576 could be attributed to uncertainties in the deposition and evolution of snow impurities and snow
577 aging processes as well as atmospheric forcing and other potential snowpack physics (e.g.,
578 densification), which requires further studies. The individual impacts of snow grain size, non-
579 spherical snow grain shape, and snow impurity on snow albedo and surface radiative forcing have
580 different signs and magnitudes. Overall, the average changes in the broadband snow albedo due to
581 the optimization of fresh snow grain size, the use of non-spherical snow shape, and including LAPs
582 at three stations are 0.018, 0.039, and -0.022. This study substantially enhances the physical
583 representations of snow albedo processes in Noah-MP, which offers a stronger snow albedo
584 modeling capability for future studies considering the wide use of Noah-MP. Future efforts are
585 needed to investigate the climate effects of aerosols in snow via land-atmosphere interaction and
586 snow albedo feedback in fully coupled meteorology-chemistry-snow models.

587

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601

602 **Open Research**

603 In situ observed albedo and snow depth from the three study sites are available on
 604 <https://data.mendeley.com/datasets/5393ck97d9/3> (Abolafia-Rosenzweig et al., 2022b).
 605 Simulation data used in this manuscript are available on
 606 <https://doi.org/10.5281/zenodo.10460675> (Lin et al., 2024). Noah-MP model code updates are
 607 publicly available: https://github.com/tslin2/hrlas_snicar.git

608

609

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